

Reports of the Department of Geodetic Science
Report No. 190

**THE OHIO STATE UNIVERSITY
GEOMETRIC AND ORBITAL
(ADJUSTMENT) PROGRAM (OSUGOP)
FOR SATELLITE OBSERVATIONS**

by

J. P. Reilly, C. R. Schwarz and M. C. Whiting

Prepared for
National Aeronautics and Space Administration
Washington, D.C.

Contract No. NGR 36-008-093
OSURF Project No. 2514



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The Ohio State University
Research Foundation
Columbus, Ohio 43212

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PREFACE

This project is under the supervision of Ivan I. Mueller, Professor of the Department of Geodetic Science at The Ohio State University, and is under the technical direction of James P. Murphy, Special Programs, Code ES, NASA Headquarters, Washington, D. C. The contract is administered by the Office of University Affairs, NASA, Washington, D. C., 20546.

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THE OSUGOP PROGRAM

1. INTRODUCTION

OSUGOP is a computer program that was developed at the Ohio State University. The name OSUGOP is an acronym for Ohio State University Geometric and Orbital Program. The basic purpose of the program is to perform adjustments for ground station coordinates from observations made to satellites by stations observing from the ground. The observations can be optical or ranges, and the adjustments can be performed in either the geometric or the orbital mode. This program is based on many smaller programs developed in 1966 and 1967 and written in the SCATRAN language for use on the IBM 7094 computer [Krakiwsky, et al., 1967 and 1968]. Later programs were also written for use on the IBM 7094, but they were written in the FORTRAN IV language.

In the spring of 1969 the geometric adjustment of optical observations and the solutions of the normal equations programs were converted to the FORTRAN IV language from the original SCATRAN listings. At that time it was anticipated that further additions to the program would be necessary, and for this reason the programming was done in such a way that additional programs could be added very easily. A system of problem codes was established that would direct the computer to perform the different adjustments. In the fall of 1969 additional subroutines were added to the system to process range observations in the geometric mode.

From the fall of 1969 to the spring of 1971, the only changes made in OSUGOP were improvements in logic and additional constraint options. Then in the spring of 1971 subroutines were added to perform an orbital adjustment. An additional change was the ability to read optical data in the GEOS format.

Of the authors listed, C. R. Schwarz was a graduate student from September, 1967 to September, 1970, and is currently with the Defense Mapping Agency Topographic Center, Washington, D. C.; M. C. Whiting was a graduate student from September, 1970 to January, 1972, and currently lives in San Francisco, California.

2. PURPOSE OF THE PROGRAM

OSUGOP is an adjustment program that can be used for many different tasks. The main purpose is to perform an adjustment for observing station coordinates. The program, however, has been developed in such a way that certain specific tasks can be performed without resorting to a complete solution.

In order to control the data flow in the program, a system of "Problem Code Definitions" has been established. These codes are numbers punched in columns 1 through 20 of a data card that is read by the program near the beginning of the data deck. After the problem codes have been read the program uses these codes to branch to the subroutines needed to perform the required task. These "Problem Code Definitions" are given in Table 1. As it can be seen there are seven (7) different types of data that can be processed (see PCODE (1)). However, at the present time (September, 1972) only five of these are operational. The documentation included in Table 1 makes the table self-explanatory. If it is desired to perform a solution (i. e., PCODE (2) = 1), it is necessary to impose some constraints on the stations in the network. Table 2 is a listing of the "Constraint Code Directory." A complete description of each of these constraints is given in Section 4.

The purpose of this report is to describe how to use the OSUGOP program. In this case it is best to start with the arrangement of the card deck for each of the five possible types of adjustments designated by PCODE (1). Figures 1 through 5 are schematics of the various cases. Notice that in all cases the deck setup is the same through the station coordinate packet. Also notice that there are no program cards. The program itself is stored on a disk pack and the JCL cards at the beginning of the deck are all that is necessary to call the program.

Table 1

PROBLEM CODE DEFINITIONS

COLUMN MEANING

1. OVERALL PROBLEM CODE
 PCODE(1)=1 MEANS OPTICAL PROGRAM
 2 MEANS RANGE
 3 MEANS SOLUTION ONLY RUN
 4 MEANS ORBITAL MODE, OPTICAL OBSERVATIONS
 5 MEANS ORBITAL MODE, RANGE OBSERVATIONS
 6 MEANS ORBITAL MODE, MIXED OBSERVATIONS
 PCODE(1)=7 MEANS OPTICAL PROGRAM, GEOMETRIC MODE (CCLS FORMAT)

2. PERFORM SOLUTION?

PCODE(2)=1 MEANS YES
 0 MEANS NO

PCODE(1)=0 IMPLIES PCODE(2)=1

3. MAXIMUM NUMBER OF ITERATIONS?
 PCODE(1) MUST EQUAL 1 OR 2,
 PCODE(2) MUST EQUAL 1,
 PCODE(3) MUST EQUAL 1, FOR ONE OR MORE COMPLETE ITERATIONS

4. FERM NORMALS?

PROCESSING CODES

1 MEANS YES, 0 MEANS NO

5. SIMULATE GUIDE MATRIX?

6. PRINT NORMALS?

7. PERFORM SUMMARY BY OBSERVED LINES?

8. PUNCH NORMALS IN ASD FORMAT?

9. PRINT SATELLITE POSITION FOR EACH EVENT?

0 MEANS NO

1 MEANS PRINT XYZ AND GEODETIC COORDINATES

2 MEANS PRINT XYZ ONLY

3 MEANS PRINT GEODETIC COORDINATES ONLY

10. THIS PARAMETER DESCRIBES WHERE THE STANDARD DEVIATIONS OF THE INDIVIDUAL OBSERVATIONS (USED TO FORM THE WEIGHTS) ARE TO BE FOUND
 PCODE(12)=0 MEANS TO READ THE OBSERVATIONAL STANDARD DEVIATION FROM THE CARD CONTAINING THE OBSERVATION.

PCODE(12)=1 MEANS TO ASSOCIATE A SINGLE STANDARD DEVIATION WITH ALL OBSERVATIONS FROM A GIVEN STATION.** THE STANDARD DEVIATIONS TO BE ASSOCIATED WITH EACH STATION ARE GIVEN IN COLUMNS 73-79 OF THE CARD CONTAINING THE INPUT COORDINATES OF THE STATION.

PCODE(12)=2 MEANS TO ASSOCIATE A SINGLE STANDARD DEVIATION WITH ALL OBSERVATIONS.** THIS NUMBER IS FOUND IN COLS. 21-26 OF THE CARD CONTAINING THE TEST DISTANCE (OPTICAL) OR TEST VARIANCE (RANGE).

** IN THE CASE OF OPTICAL OBSERVATIONS, THIS NUMBER IS INTERPRETED AS THE STANDARD DEVIATION OF THE DECLINATION AND OF THE RIGHT ASCENSION TIMES THE COSINE OF THE DECLINATION, AND THE COVARIANCE IS SET TO ZERO.

SOLUTION CODES

11. WRITE NORMALS AND INVERSE DURING SOLUTION PROCESSING?

0 MEANS PRINT NOTHING

1 MEANS PRINT PIVOT ELEMENTS

2 MEANS ALSO PRINT NORMALS AND INVERSE

3 MEANS ALSO PRINT REARRANGED NORMALS AND INVERSE

12. PUNCH ADJUSTED STATION XYZ AND VARIANCES FOR INPUT TO BAPEKAS' BATCH TRANSFORMATION PROGRAM?

13. PUNCH ADJUSTED STATION POSITIONS?

14. COMPUTE EIGENVECTORS OF VARIANCE-COVARIANCE MATRIX

15. COMPUTE CORRELATION COEFFICIENTS

Table 2

3. INPUT TO THE OSUGOP PROGRAM

The input is made up of card packets, which are groups of cards containing a variable number of cards, and signal cards. Each packet is terminated by an end signal card, which is blank in columns 1-79 and contains "E" in column 80. The one exception is the optical observation packet using the GEOS format. Here the "E" must be punched in column 73 (denoted "special end card" in Figure 2). Depending on the type of run, some packets may or may not be necessary.

3.1 Card Format for Required Cards

Title Packet (always required): As many title cards as desired are permitted, containing any text in columns 1-79. This text appears verbatim on the first page of the output. An end signal card terminates this packet.

Problem Codes (always required): These codes appear on a single card and control the type of processing to be performed by the program. See Table 1 for a description of each code and the column in which it is punched. (Do not put an end signal card after this card).

Datum Card Packet (always required): This contains a list of the ellipsoids on which the input and output ellipsoidal coordinates of the stations are located. Each datum is described by 2 cards.

Card 1.

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-2	I2	Identifying number of datum.
3-15	F13.2	Semi-major axis of ellipsoid.
16-28	F13.2	Semi-minor axis of ellipsoid.

Card 2.

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-32	4A8	32 character alphabetic name of datum.

The datum packet is terminated by an end signal card.

Station Coordinate Card Packet (always required): Each card gives the input (or approximate) coordinates of a station.

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-4	I4	Identifying number of the station.
5-6	I2	Identifying number of the ellipsoid to which the ellipsoidal coordinates refer.
7-24	4A4, A2	18 character station name.
25	A1	Sign of latitude.
26-28	I3	Degrees of latitude.
29-31	I3	Minutes of latitude.
32-39	F8.4	Seconds of latitude.
40-42	I3	Degrees of longitude (+East).
43-45	I3	Minutes of longitude.
46-53	F8.4	Seconds of longitude.
54-63	F10.2	Ellipsoid height (in meters).
73-79	F7.2	Standard deviation to be used for all observations from this station (IF PCODE (12) = 1).

There is one card for each ground station in the network. This packet is terminated by an end signal card.

The identifying number of the datum for the station coordinates must correspond to the number defining the datum. For example, if one only has station coordinates on the North American Datum, the datum card packet could contain a card with the number 1 in column two, and the numbers 6378206.4 and 6356583.8 for the semi-major and semi-minor axes. Then on each station coordinate card one would have to put the number 1 in column six to show that these coordinates refer to the North American Datum.

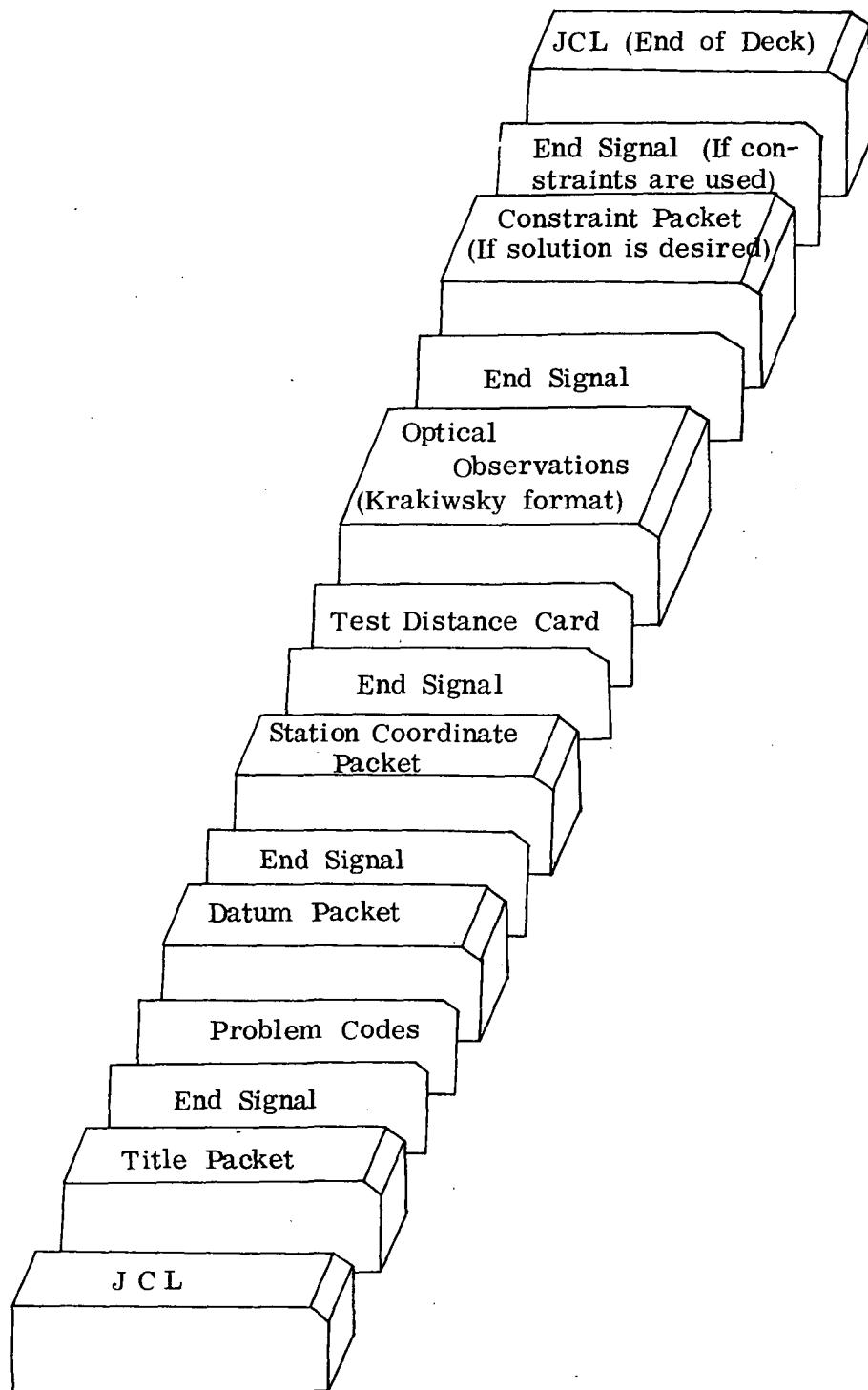


Figure 1. Deck Setup for Optical Program (OSU Format).

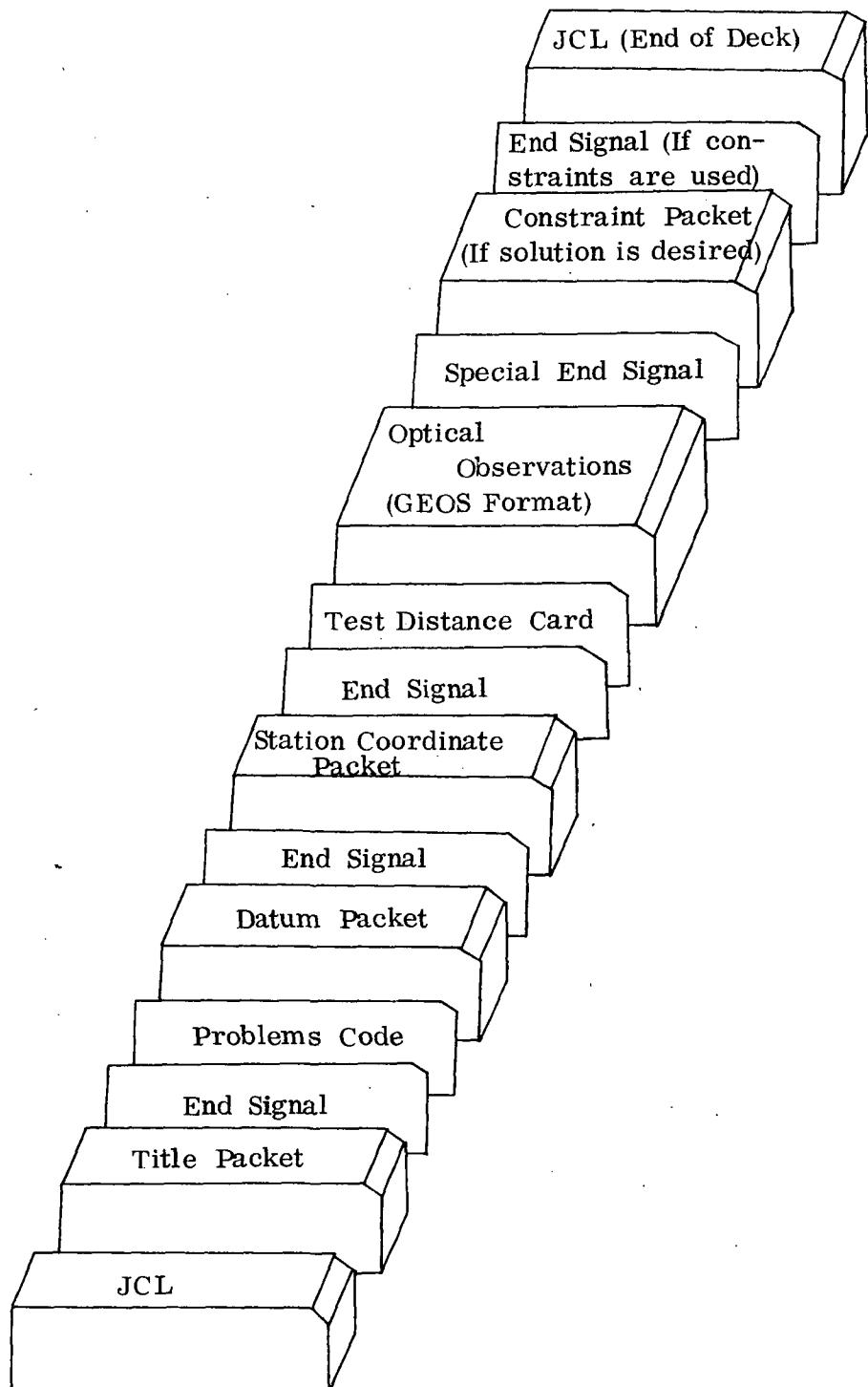


Figure 2. Deck Setup for Optical Program (GEOS Format).

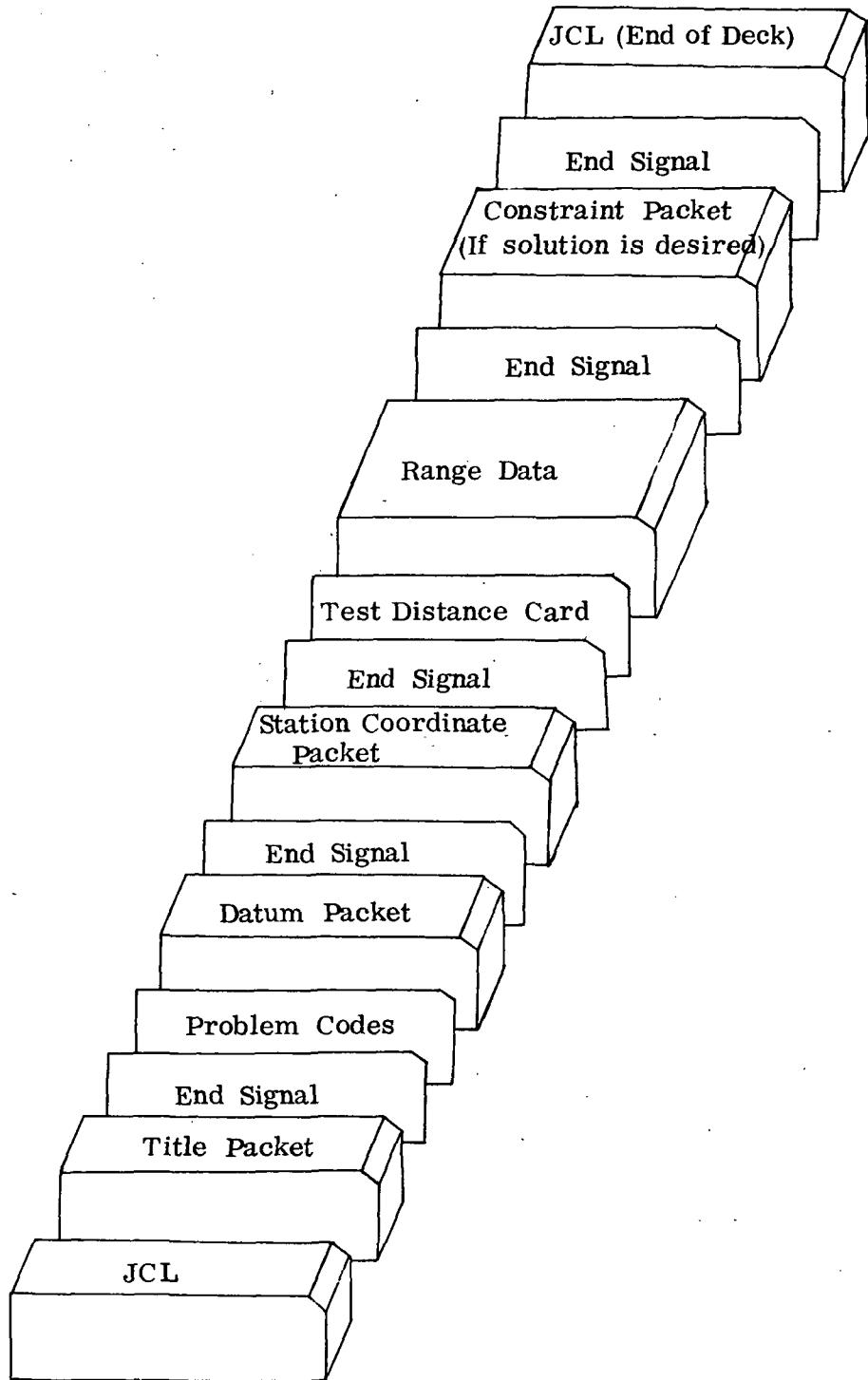


Figure 3. Deck Setup for Range Program in the Geometric Mode.

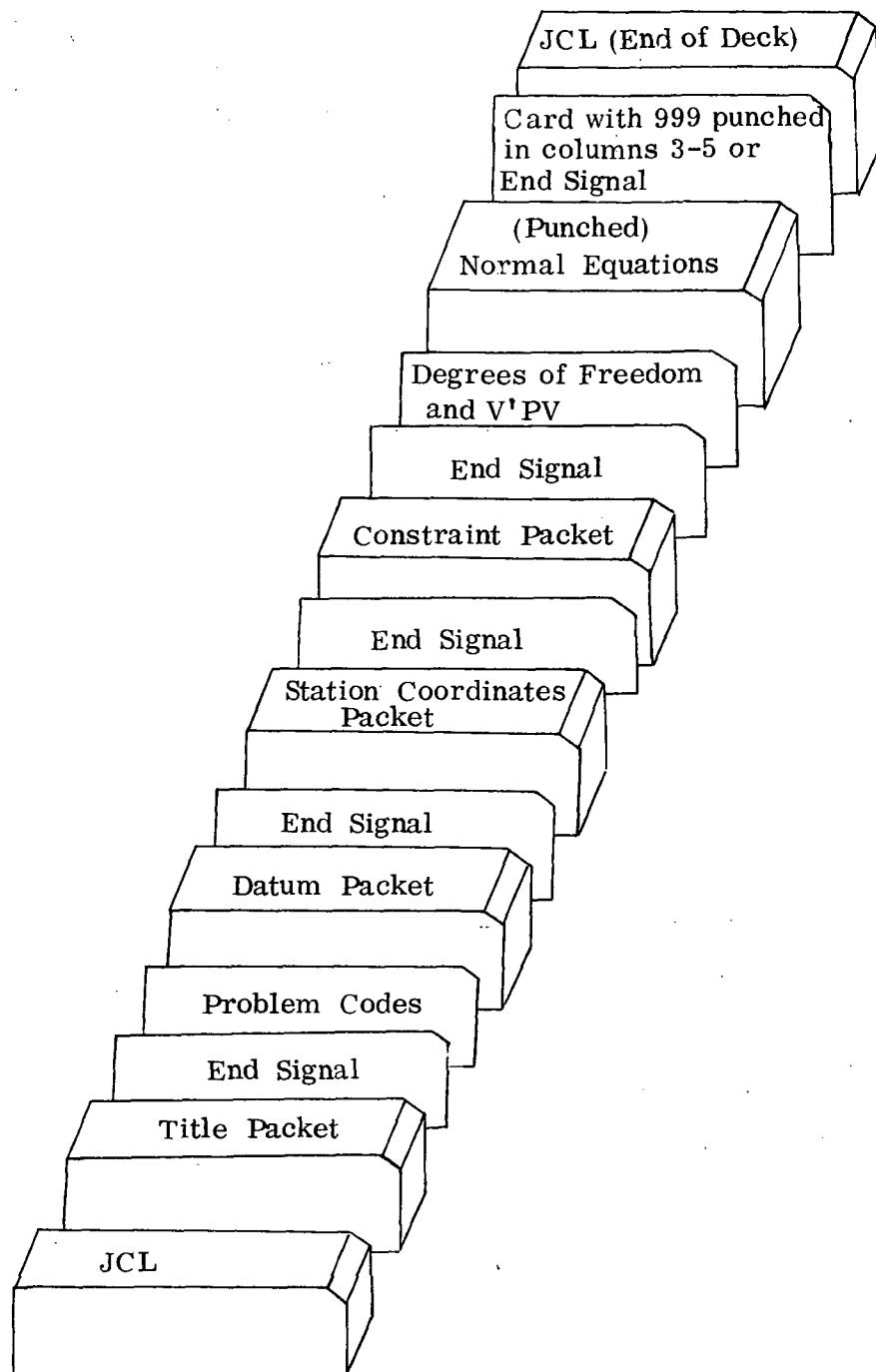


Figure 4. Deck Setup for Solution Using Normal Equations Punched on Cards.

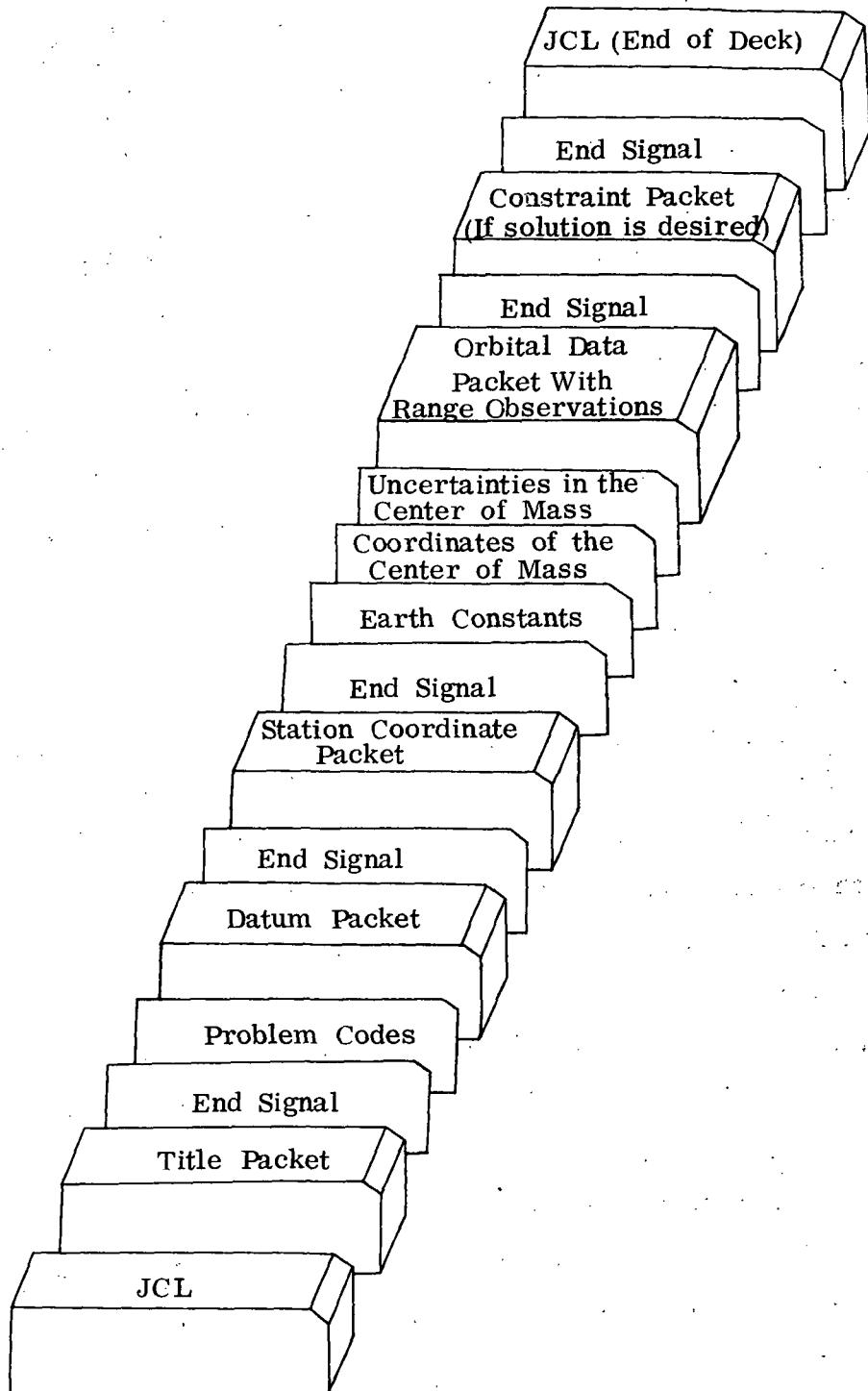


Figure 5. Deck Setup for Range Observations in the Orbital Mode.

3.2 Card Formats for the Different Options

There are five different types of input that can be handled by this program. The deck setup for each type of observation will be discussed separately.

3.2.1 Optical Observations, Geometric Mode.

There are two different card formats for the optical observations. However, the adjustment itself is the same for both. The reason for the two different formats is that the original optical programs written in SCATRAN utilized a card format that was convenient for the programmer and is referred to as the OSU format. Whenever the first optical data from the GEOS I satellite became available, it was easier to change the observations to the OSU format than to modify the program. When the optical program was later converted to FORTRAN IV the OSU format was retained. Because the GEOS format has been accepted as a standard format, the computer program was modified in 1971 to accept either card format.

As can be seen from Figures 1 and 2, it is necessary to have a test distance card in front of the optical observations. This test distance card is used to specify a rejection criteria for each observation. The purpose of the rejection criteria is to eliminate bad or questionable observations from the adjustment automatically without physically removing the observation cards from the card deck. The optical program is designed to read all the observations that have the same time of observation, and then perform an adjustment for the position of the satellite. The approximate station coordinates of the observation stations are held fixed during this adjustment. If the approximate coordinates of the observing stations are known to a certain degree of accuracy, and if the observations are known to a certain accuracy, the accuracy of the adjusted position of the satellite can be predicted.

The need to have a rejection criteria may not be very obvious, but from

past experience it has proved to be very useful. A fairly large percentage of the optical data received from the Space Science Data Center in Washington had large blunders. By setting the rejection criteria at a reasonable value, all bad observations were eliminated from the solution.

The rejection criteria can also be used when the data is reasonably free from bad data but some of the approximate station coordinates are not well known. Here the rejection criteria can be set fairly high for one or two iterations so that all observations are accepted and the questionable station coordinates are allowed to adjust.

In addition to specifying the rejection criteria, there is a place on the test distance card to insert a value for the standard deviations of all observations that will override the actual standard deviation punched on the observation cards. The standard deviations will then be the same for $\alpha \cos \delta$ and δ , and the covariance term will be zero. The need for this feature in the program became apparent when the standard deviations on the observation cards were noted to be completely out of line.

3.2.1.1 Arrangement of Optical Observations.

The only requirement in the arrangement of observation cards is that they are grouped by events. An event is all the data that has been observed on a satellite at the same instant of time. The time on the data cards should be the same, but the computer program will allow for a deviation of 0.0002 seconds. There must be at least two (2) observation cards in an event (i.e., a minimum of two stations must observe the satellite at the same time).

If there are more than two stations observing at the same instant of time the program will perform an adjustment for the satellite position starting with all observations. If any of these observations are bad the program will delete them and perform an adjustment with the remaining observations. If after deleting the bad observations there are less than two stations observing, the entire event is deleted.

The format of the optical observations must agree with the code punched on the problem code card. Using the OSU format PCODE (1) = 1, and using the GEOS format PCODE (1) = 7. The card formats peculiar to these optical observations are described below:

Test Distance Card.

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-20	F20.2	Rejection criteria, in seconds of arc, to be applied to each observation during editing.
21-30	F10.2	Standard deviation, in seconds of arc, to be used for all observations (if PCODE (12) = 2).

Optical Observations (OSU Format) Card Packet.

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1- 3	I3	Station identification number.
6-19	I2, I3, F9.4	Hours, minutes, seconds of observation (<u>expressed in UT1</u>).
20-26	I2, A3, I2	Day, month, year of observation (Note: month can be either three letters such as Jan, Feb, etc., or else the number 1, 2, . . . , 12).
27-41	2I3, F9.5	Hours, minutes, seconds of right ascension (α).
42-55	A1,I2, I3, F8.4	Sign, degrees, minutes, seconds of declination (δ).
58-62	F5.2	Standard deviation of α multiplied by the cosine of the declination, in seconds of arc.
63-67	F5.2	Standard deviation of δ , in seconds of arc.
68-72	F5.2	Covariance between $\alpha \cos \delta$ and δ , in seconds of arc squared.

This packet is terminated by an end signal card.

Optical Observations (GEOS Format) Card Packet

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
15-18	I4	Station identification number.
19-24	3I2	Year, month, day of observation (Note: Month <u>must</u> be expressed as a number)
25-34	2I2, F6.4	Hours, minutes, seconds of observation (<u>expressed in UT1</u>).
35-44	I3, I2, F5.3	Hours, minutes, seconds of right ascension (α).
45-53	A1, 2I2, F4.2	Sign, degrees, minutes, seconds of declination.
72-74	F3.2	Standard deviation of α multiplied by the cosine of the declination, in seconds of arc.
75-77	F3.2	Standard deviation of declination δ , in seconds of arc.
78-80	F3.1	Covariance between $\alpha \cos \delta$ and δ , in seconds of arc squared.

This packet is terminated by a special end signal card. This card has the letter E punched in column 71.

3.2.2 Range Observations, Geometric Mode.

As with the optical observations, a test card is required in front of the range observations. The purpose of the card is the same as described for the optical observations. The arrangement of the range observations has the same basic requirement as optical observation, and this is the grouping of observations by events. The minimum number of observations for an event is four (4). This subprogram does not have the provision for elimination of individual observations from an event; if one observation is bad, the entire event is deleted.

When using the range observations in the geometric mode, PCODE(1) = 2. The card format for the range observations is the GEOS range format. The

GEOS format as given by NASA uses all 80 columns of the data card. Since many columns are used for information such as satellite number, year of launch, etc., they are not included in the card format description that follows. The only information included in this description is data pertinent to the observation itself.

Test Distance Card

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-20	F20.2	Rejection criteria, in meters, to be applied to each event.
21-30	F10.2	Standard deviation, in meters, to be used for all observations (if PCODE (12) = 2).

Range Observation Card Packet

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
15-18	I4	Station identification number.
19-24	3I2	Year, month, day of observation.
25-34	I2, I2, F6.4	Hours, minutes, seconds of observation.
44-53	F10.3	Range, in meters.
65-70	F6.3	Standard deviation, in meters.

This packet is terminated by an end signal card.

The time (hours, minutes, seconds) of observation for range observations in the geometric mode can be in any time system. In every other type of adjustment the times must be in the UT1 time system, but in this case, the times are used only to distinguish the different events.

3.2.3 Solution Using Punched Normal Equations.

Whenever the OSUGOP Program is used with observations, it is possible to punch the normal equations on cards. These normals are punched prior to the addition of constraints, which means that the matrix of the normal equations is singular. This punching is made possible by setting PCODE (9) = 1.

The reasons for punching the normal equations are many. The most common use is that if many different solutions are to be run using the same observations, but with different constraints it is more efficient to form the normals only once, and then run the different solutions by changing only the constraints. It may take fifteen minutes of computer time to form the normal equations, and only fifteen seconds to perform the solution. Another important reason for punching normal equations is for use in a combination solution of two or more different systems of equations. The different sets of normal equations can be solved together, using constraint or ties between the systems.

Because of the need for punched normal equations, the punched output should be in such a format that the user can easily distinguish which rows and columns refer to a particular station. This has been accomplished by forming the normal equations in what is referred to as the ASD format. This is a collapsed form of normal equations where all zero elements have been eliminated. All matrix elements are in the form of 3×3 matrices (3 unknowns for each station). However, additional information other than just the elements of the matrix is punched. It is necessary to include the stations numbers and a code number to indicate the end of a row. Also, the degrees of freedom (d.o.f.) and the summation of $V'PV$ are needed. Since these normal equations are really a system of equations, the discrepancy vector, U , must be included. The discrepancy vector is the vector U in the expressions

$$NX + U = 0$$

$$X = -N^{-1}U.$$

The punched output of normal equations starts by punching the d.o.f. and $\Sigma V'PV$ on one card, and then punching each row of the normal equation. The first card for each row is the station number. The next card gives the three elements of the U vector that corresponds to this station number.

The next three cards are the elements of the 3×3 matrix that correspond to this station. Following this are the off-diagonal elements, which in this case are the 3×3 matrices corresponding to all the other stations that co-observed with the first station of this row. These are denoted by punching a card with the station number, and then three cards with the 3×3 matrix. This is repeated for all stations that co-observed. After the last set of elements for the row have been punched, a card containing the number 999 is punched to indicate the end of the row. After that the next row is punched, and so on until the end of the matrix has been reached.

A sample of the punched output can be seen in Figure 6. The first line shows the d.o.f. (4069) and $\Sigma V'PV$ (6737.147170). The next printed line indicates that the first row of normal equations has station number 9066 on the diagonal. The following card gives the three elements of the U vector, and the three cards following this are the elements of the 3×3 matrix corresponding to station 9066. After the diagonal elements come the stations that co-observed with station 9066, and the 3×3 matrix of off-diagonal elements for each of these stations. In this case, they are stations 8015 and 8019. The end of the first row is marked by the 999 (15th line).

Whenever the computer is asked to punch a set of normal equations, it is always a good idea to set the proper PCODES to print the normals also, as well as a guide matrix to indicate the layout of the matrix. This can be done with PCODE (6) and PCODE (7) (see Table 1). The guide matrix that corresponds to the normal equations in Figure 6 is given below:

	4069.000000	6737.147170	
9066			
	0.4463324330	-0.7695279730	-0.4691234316
	0.0441544859	0.0064557696	-0.0021695353
	0.0064557696	0.1576382832	-0.0109005340
	-0.0021695353	-0.0109005340	0.0317712453
8015			
	-0.0077443357	-0.0022595720	-0.0029675092
	-0.0023964059	-0.0314680666	0.0096666831
	-0.0026492987	0.0092509108	-0.0125750867
8019			
	-0.0164675851	0.0072840522	-0.0042015838
	0.0076429834	-0.0779199552	-0.0006815355
	-0.0018020912	-0.0010346222	-0.0058853776
999			
8015			
	-1.2612411155	0.1753680388	0.4569007003
	0.1472091420	0.0050456835	-0.0428749431
	0.0050456835	0.1041464281	-0.0037635258
	-0.0428749431	-0.0037635258	0.1294911894
8019			
	-0.1198067022	0.0047556960	0.0421808712
	0.0045907000	-0.0467891914	-0.0123380813
	0.0427791428	-0.0057905394	-0.1058683524
9051			
	-0.0106327398	0.0011620229	0.0030444112
	0.0011620229	-0.0001892168	-0.0005283645
	0.0030444112	-0.0005283645	-0.0014868776
999			
9080			
	-0.4712158212	0.0760527941	0.2145824528
	0.0162138738	0.0009955345	-0.0122346437
	0.0009955345	0.0046783024	-0.0012469197
	-0.0122346437	-0.0012469197	0.0242084646
999			

Figure 6.

GUIDE MATRIX

9066	8015	8019	999
8015	8019	9051	999
9080	999		

Although it cannot be easily seen from the above guide matrix, the matrix of normal equations is upper-triangular. Another example of a guide matrix is shown in Figure 16.

When using punched normal equations to perform an adjustment, the deck setup is as shown in Figure 4. After the constraint packet, the card containing d.o.f. and $\Sigma V'PV$ is first, then the normal equations, and at the end the additional 999 card or an end-signal card with the letter E punched in Column 80. PCODE(1) is set equal to 3 in this case.

3.2.3.1 Combining Different Systems of Normal Equations.

If different systems of normal equations are to be combined for a single adjustment, the only additional work required is to physically combine the normal equations together into one data deck. When doing this, there are several things that must be done:

1. There can only be one card with degrees of freedom and $\Sigma V'PV$. Therefore, the values for each set of normal equations should be added together and the total d.o.f. and $\Sigma V'PV$ punched on one card.
2. There can only be one row for any one observing station. If station 9000 is included in more than one set of normal equations, the 3×3 matrices of diagonal elements corresponding to this station in each set of normal equations

must be added together to form one 3×3 diagonal matrix. The same goes for off-diagonal elements. If station 9000 co-observed with station 9010 in more than one set of normal equations, these 3×3 matrices must be added together to form one matrix.

3. Along the same line of reasoning mentioned in 2., if station 9000 is included in more than one set of normals, but in the later set co-observed with a station that was not included in the first set, the off-diagonal matrix corresponding to that station in the later set of normals must be moved into the row of the first set of normal equations. This can best be illustrated by guide matrices for two different systems:

First Set of Normal Equations

1	2	3	4	999
2	3	4	999	
3	4	999		
4	999			

Second Set of Normal Equations

4	5	6	999
5	6	999	
6	999		

Combined Set of Normal Equations

1	2	3	4	999
2	3	4	999	
3	4	999		
4	5	6	999	
5	6	999		
6	999			

4. Make sure that the matrix of normal equations is upper-triangular. As can be seen from the guide matrix of combined normals above, Station 1 co-observed with Station 2, but when one forms the second row, the 2-1 station combination is not repeated.

3.2.4 Range Observations, Orbital Mode.

In order to use range observations in the orbital mode, the deck set-up requires a great deal of work. It is necessary to have three data cards at the beginning of the data packet that give the earth constants, coordinates of the center of mass, and the uncertainties in the center of mass (see Figure 5). After this, the observations are separated into passes, and with each pass must be included the approximate orbital elements at a particular epoch and a code to tell the program what the coordinate system is. The epoch time is also included. Each pass is then separated by an end signal card. Because of the complexity of the deck setup, each step will be described in detail.

3.2.4.1 Earth Constants.

The earth constants are the semi-major axis of the earth, GM (or kM), gravitational constant \times mass, and the rate of rotation of the earth. Also included on the earth constants card is the standard deviation of the observations if one wants to override the actual standard deviation punched on the observation cards. The card format is as follows:

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
1-20	D20.8	Semi-major axis of the earth, in meters.
21-40	D20.8	GM, in units of cm^3/sec^2 .
41-60	D20.8	Rotation rate of the earth, in radians/sec.
61-80	D20.8	Standard deviation, in meters, to be used for all observations (if PCODE (12) = 2).

3.2.4.2 Coordinates of the Center of Mass.

The coordinates of the center of mass give the location of the center of mass with respect to the origin of the ellipsoid used in the adjustment. The coordinates of the center of mass are given in the coordinate system in which the station coordinates are given.

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
<u>Card 1</u>	1-15	D15.8	X coordinate of the center of mass.
	16-30	D15.8	Y coordinates of the center of mass.
	31-45	D15.8	Z coordinate of the center of mass.
<u>Card 2</u>	1-15	D15.8	uncertainty of the X coordinate of the center of mass, in meters.
	16-30	D15.8	uncertainty of the Y coordinate of the center of mass, in meters.
	31-45	D15.8	uncertainty in the Z coordinate of the center of mass, in meters.

3.2.4.3 Orbital Data.

When inputting orbital data, the data is separated so that each orbit, or pass, is input separately. The first card of each pass gives the pass number, name and a code to indicate the type of preliminary orbital elements used to describe the orbit (see details below). The next two cards contain the orbital elements. The fourth card contains the time of epoch. After the time card, the observations for the particular pass are read. The cards do not have to be in any order, but to conserve computer time it is best to arrange the observations in the order of increasing time. The last card for the pass is an end signal card.

Each pass is arranged the same way as described above and placed one behind the other. There is no required order for arranging the different passes; the program works on each pass separately. After all passes are inserted in the deck, an extra end signal is placed after the last pass.

(Note: This end signal card is in addition to the one at the end of the last pass of data.)

The first card of the pass can have any number or name; these are used for identification purposes only. The code (IOCODE) used to indicate the type of preliminary orbital elements must be one of the numbers 0 through 4.

IOCODE = 0 - means rectangular elements are given in the True Sidereal System.

IOCODE = 1 - means rectangular elements are given in the Modified Sidereal System.

IOCODE = 2 - means the rectangular elements are given in the Earth-Fixed System.

IOCODE = 3 - means Keplerian elements are given, referred to the true equator.

IOCODE = 4 - means Keplerian elements are given, referred to the true equator and the 1950.0 equinox (i.e., the SAO Orbital System).

If the value of IOCODE is 0, 1 or 2, the orbital elements are expressed as X, Y, Z on one card, and $\dot{X}, \dot{Y}, \dot{Z}$ on a second card. If IOCODE is 3 or 4, the orbital elements are expressed as the semi-major axis, eccentricity and inclination on the first card, and right ascension of the ascending node, argument of perigee, and mean anomaly on the second card.

The fourth card of each pass is the same regardless of the type of orbital elements used. This gives the epoch time, which is the time that corresponds to the orbital elements. This particular card, at first glance, appears to be very confusing due to the fact that there are several options for specifying epoch time. Figure 7 is a sample of the fourth card layout. If the value of Z CODE is left blank, it means that the orbital elements refer

EPOCH(MJD)	IDAY	MONTH	IYR	IH	MIN	ESEC	ZCODE	(IHR)	IMIN	SEC

Figure 7.

to the epoch time given at the left side of the card, and can either be expressed in MJD or day, month, year, hour, minutes, seconds. If the value of ZCODE is anything other than a blank, it means that the epoch time is the hours, minutes and seconds given on the right-hand side of the card. The distinction between these two times is that if the epoch time is given on the right-hand side of the card, the epoch time is outside the timespan of the pass, and the desired epoch time is the time given on the left side of the card. It means that the actual epoch time is that given to the right of ZCODE, which may be as much as 24 hours away from the time of the pass. In this case, the computer program updates the elements to the time given on the left side of the cards. Care must be taken to insure that the proper day is given, since only hours, minutes and seconds are given on the right-hand side of the card.

After the time card, the range observations for that particular pass are inserted with an end card placed after the last range card.

The card formats are as follows:

Orbital Data for Each Pass

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1-4	A4	Orbit number (can be anything, used for analysts identification only).
	5-52	6A8	Orbit name (for identification only, this can be left blank, if desired).
	53	I1	IOCODE. This is the number 0, 1, 2,3 or 4 depending on the coordinate system of the orbital elements.
(i) <u>Orbital elements given in rectangular coordinates</u> (IOCODE = 0, 1 or 2).			
Card 2	1-15	D15.8	X coordinate of satellite, in meters.
	16-30	D15.8	Y coordinate of satellite, in meters.
	31-45	D15.8	Z coordinate of satellite, in meters.

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 3	1-15	D15.8	\dot{X} , the velocity component in the X direction, in meters/sec.
	16-30	D15.8	\dot{Y} , the velocity component in the Y direction, in meters/sec.
	31-45	D15.8	\dot{Z} , the velocity component in the Z direction, in meters/sec.

(ii) Orbital elements given as Keplerian elements
(IOCODE = 3 or 4).

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 2	1-15	D15.8	Semi-major axis of orbital ellipse, in meters.
	16-30	D15.8	Eccentricity of orbital ellipse.
	31-45	D15.8	Inclination of ellipse to equatorial plane, in degrees and decimal degrees.
Card 3	1-15	D15.8	Right ascension of the ascending node, in degrees and decimal degrees.
	16-30	D15.8	Argument of perigee, in degrees and decimal degrees.
	31-45	D15.8	Mean anomaly, in degrees and decimal degrees.

Epoch Time Card

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 4	1-15	D15.8	Epoch time expressed in modified Julian days (MJD).
	16-20	I5	Day of the month.
	23-25	A3	Month of the year. This can be the numbers 1 through 12 or the first three letters of the month's name, such as JAN, FEB., etc.
	26-30	I5	Year (last two digits).
	31-35	I5	Hours.
	36-40	I5	Minutes.
	41-50	D10.5	Seconds.

If the orbital elements given are at a time not in the actual pass itself, this time is punched as follows:

<u>Columns</u>	<u>Format</u>	<u>Contents</u>
51-55	I5	Hours
56-60	I5	Minutes
61-70	D10.5	Seconds

(Note: In order for the computer to use this time, something other than blanks must be in columns 54 or 55. Therefore, even if the time is zero hours, the zeros must be punched in columns 55 or 54 and 55).

It should be noted that on the format of card No. 4, there is no ZCODE as such described. The FORTRAN coding has been done in such a way that columns 54 and 55 are recognized as ZCODE and also as part of the hour value.

Observations Cards.

The format for the observation card for the orbital adjustment is identical to the observation card format described in the Range Observations, Geometric mode section of this report (3.2.2).

The end of data on each pass is marked by placing an end signal card after the last observation.

4. CONSTRAINTS

Table 2 is the directory of the constraint codes needed to apply constraints to normal equations prior to a solution. There are ten different types of constraints that can be applied, five types of weighted constraints and five types of absolute constraints. In all cases, the first card gives the constraint code, and then depending on the type of constraint, the cards following give the required information necessary to apply the constraint.

4.1 Weighted Constraints.

4.1.1 Constrain the Coordinate of a Station at its a Priori Value.

This constraint is used to weight any one or all three Cartesian coordinates of a station. It is used primarily to control the translation or to define the origin of a network of stations. The weight needed to apply this constraint is

$$W = \frac{\sigma_o^2}{\sigma_i^2}$$

where

σ_o^2 is the a priori unit variance (which, in most cases, is assumed to be 1).

σ_i^2 is the variance of the component of the station coordinate, in meters squared.

Four cards are needed to apply this constraint. The first card is the constraint code, which in this case is 1. The second card is the number of the station to be constrained. On the third card are listed the coordinates to be constrained, and the fourth card gives the weights to be applied to each of the coordinates.

If the coordinates to be constrained are the approximate coordinates given at the beginning of the program, the third card is replaced by a blank.

The card formats are as follows:

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1- 2	I2	Constraint code, which in this case is the number 1.
Card 2	1- 5	I5	Station number of the station to be constrained.
Card 3	1-16	D16.8	The X component of the station coordinate, in meters.
	17-32	D16.8	The Y component of the station coordinate, in meters.
	33-48	D16.8	The Z component of the station coordinate, in meters.
			Any one, or all three of the coordinates can be constrained. If only one or two components are to be constrained, let the field blank for the part not to be constrained. If the approximate coordinates as given at the beginning of the program are to be constrained, card 3 should be blank.
Card 4	1-16	D16.8	Weight to be applied to the X component.
	17-32	D16.8	Weight to be applied to the Y component.
	33-48	D16.8	Weight to be applied to the Z component.

4.1.2 Chord Distance Constraint.

The chord distance constraint is used primarily to apply a scale. The chord distance is computed by the program if the approximate station coordinates are to be used to compute the chord. If the chord distance is known from another source, the distance is punched onto a card.

Three cards are needed to apply this constraint. The first card is the constraint code, which is 2. The second card gives the station numbers of the two stations involved in the chord constraint. The third and last card gives the chord distance, and the accuracy of the distance. If the

accuracy of the chord distance is 1 part in 500,000, the number punched on the card is 500,000. The card formats are as follows:

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1- 2	I2	Constraint code, which is 2.
Card 2	1- 5	I5	Station number of the first station.
	6-10	I5	Station number of the second station.
Card 3	1-16	D16.8	Chord distance (in meters). This is left blank if the chord distance is to be computed from the approximate coordinates.
	17-32	D16.8	Accuracy of chord distance, expressed as the denominator of the accuracy ratio.

4.1.3 Relative Position Constraint.

This constraint is used when two or more stations have known positions with respect to each other. After the adjustment, the relative positions of these stations should remain unchanged, or the change should be within the limit of accuracy of the survey that tied the stations together. A common use for this constraint is where two or more stations are observing from the same small island where the positions of the stations are known on a local datum survey but the positions on the datum of the adjustment are not known.

The relative position constraint can only be applied between two stations at a time. If there are more than two stations involved, additional relative positions constraints must be used. As an example, if the relative positions between stations 1, 2 and 3 are to be constrained, a constraint can be applied between stations 1 and 2, and an additional constraint between stations 2 and 3. A third constraint can be applied between stations 1 and 3, but it isn't necessary.

Four cards are needed to apply this constraint. The first card is

the constraint code, which is 3. The second card gives the station numbers of the two stations involved in the relative position constraint. The third card gives the ΔX , ΔY , ΔZ coordinate difference between the two stations that is to be constrained during the adjustment. The fourth and last card gives the weights of the coordinate differences. A word of caution is necessary about the sign convention. The signs of the coordinate differences on card three must correspond to the order the station numbers appear on card two. As an example, if on card two the station numbers are 1 and 2, with 1 being punched in the first field, the sign convention for card three must be $X_1 - X_2$, $Y_1 - Y_2$, and $Z_1 - Z_2$.

The card formats are as follows:

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1-2	I2	Constraint code, which is 3.
Card 2	1-5	I5	Station number of first station (1).
	6-10	I5	Station number of second station (2).
Card 3	1-16	D16.8	Coordinate difference ΔX , expressed as $X_1 - X_2$.
	17-32	D16.8	Coordinate difference ΔY , expressed as $Y_1 - Y_2$.
	33-48	D16.8	Coordinate difference ΔZ , expressed as $Z_1 - Z_2$.
Card 4	1-16	D16.8	Weight to be applied to the ΔX coordinate difference.
	17-32	D16.8	Weight to be applied to the ΔY coordinate difference.
	33-48	D16.8	Weight to be applied to the ΔZ coordinate difference.

If the approximate coordinates are to be used to compute ΔX , ΔY and ΔZ , card 3 should be left blank.

4.1.4 Direction Constraint.

When the direction between two stations i and j is to be constrained, it can be accomplished by applying weights to two angles α and β defining the direction between them. These angles are defined as

$$\alpha = \tan^{-1} \frac{\Delta Y}{\Delta X}$$

$$\beta = \tan^{-1} \frac{\Delta Z}{R}$$

where

$$\Delta X = X_i - X_j$$

$$\Delta Y = Y_i - Y_j$$

$$\Delta Z = Z_i - Z_j$$

and

$$R = (\Delta X^2 + \Delta Y^2)^{\frac{1}{2}}$$

As with some of the other constraints, if the directions are to be computed from the approximate station coordinates, it is not necessary to precompute α and β .

Four cards are needed to apply this constraint. The card formats are as follows:

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1-2	I2	Constraint code, which is 4.
Card 2	1-5	I5	Station number of first station.
	6-10	I5	Station number of second station.
Card 3	1-16	D16.8	Alpha (α), in seconds.
	17-32	D16.8	Beta (β), in seconds.
Card 4	1-16	D16.8	Standard deviation of α , in seconds of arc.
	17-32	D16.8	Standard deviation of β , in seconds of arc.
	33-48	D16.8	Covariance term, in seconds of arc ² .

If the approximate coordinates are to be used to compute α and β , Card 3 should be left blank.

4.1.5 Constraint on Geodetic Latitude, Longitude and Height.

This constraint can be applied to the latitude, longitude and height or to any one of the three. The main use for this constraint has been to apply height constraints to island stations where the orthometric height has been well determined, and the separation between the geoid and the reference ellipsoid is known to a certain degree of accuracy. It can also be used to define the origin of a network. This is identical in concept to constraint code 1 except here the coordinates constrained are φ, λ, h .

Four cards are needed to apply this constraint. The card formats are as follows:

	<u>Columns</u>	<u>Format</u>	<u>Contents</u>
Card 1	1-2	I2	Constraint code, which is 5.
Card 2	1-5	I5	Station number of the station to be constrained.
Card 3	1-16	D16.8	Latitude φ , in degrees and decimal degrees.
	17-32	D16.8	Longitude λ , in degrees and decimal degrees.
	33-48	D16.8	Height h , in meters.
Card 4	1-16	D16.8	Standard deviation of φ , in seconds of arc.
	17-32	D16.8	Standard deviation of λ , in seconds of arc.
	33-48	D16.8	Standard deviation of h , in meters.

Any one, or all three of the coordinates can be constrained. If only one or two components are to be constrained, leave the field blank for the part not to be constrained. If the approximate coordinates are to be constrained, Card 3 should be blank.

4.2 Absolute Constraints

The five absolute constraints are listed in Table 2. Three of these constraints use the inner adjustment equations, and for this reason a more detailed description is necessary [Blaha, 1971].

Whenever an adjustment is to be performed on a network of observing stations, it is necessary to define an origin, establish some form of orientation, and set a scale. With optical observations, the orientation is determined from the observations themselves, and with range observations the scale is determined from the observations. The inner adjustment constraint package was developed for use when the origin, orientation or scale was not known. An example of its use could be on a net of observing stations, each station on an isolated island in the ocean. If the observing stations were cameras, the origin and scale would have to be determined before adjustment. By applying constraint codes 11 and 13, the program would use the inner adjustment equations to get the best origin and scale possible from the geometry of the network and the observations themselves.

Only one card is necessary to call any one of the inner adjustment constraints. This is the same as the first card of the weighted constraint package, which is the code number punched in columns 1 and 2 of the card. If the origin is to be defined, use code 11; for orientation, code 12; for scale, code 13. Codes 14 and 15 are not operational, but the same results can be obtained by using constraints 1 and 3, using very large weights.

4.3 Using Constraints Only in an Adjustment

In addition to the five different types of adjustment (the deck set-ups

of which are described in Figures 1 thru 5) it is possible to perform an adjustment without observations. This can be done by using constraints only. If there are enough constraints applied to tie all the stations together, this is equivalent to forming a set of normal equations. In this computer program, the constraints are always added to the existing normal equations (see [Mueller, et al., 1970], pp. 10-16 for a description of this). If the existing normal equations do not exist, the normal equations can be formed entirely from constraints. Care must be taken to insure that all stations are constrained properly.

The deck set-up for solution using constraints only is shown in Figure 8. This is identical to the deck set-up for a solution only (see Figure 4) except the degrees of freedom card and the punched normal equations are replaced by a blank card. As with the solution only run shown in Figure 4, PCODE (1) must be set equal to 3 on the problem codes card.

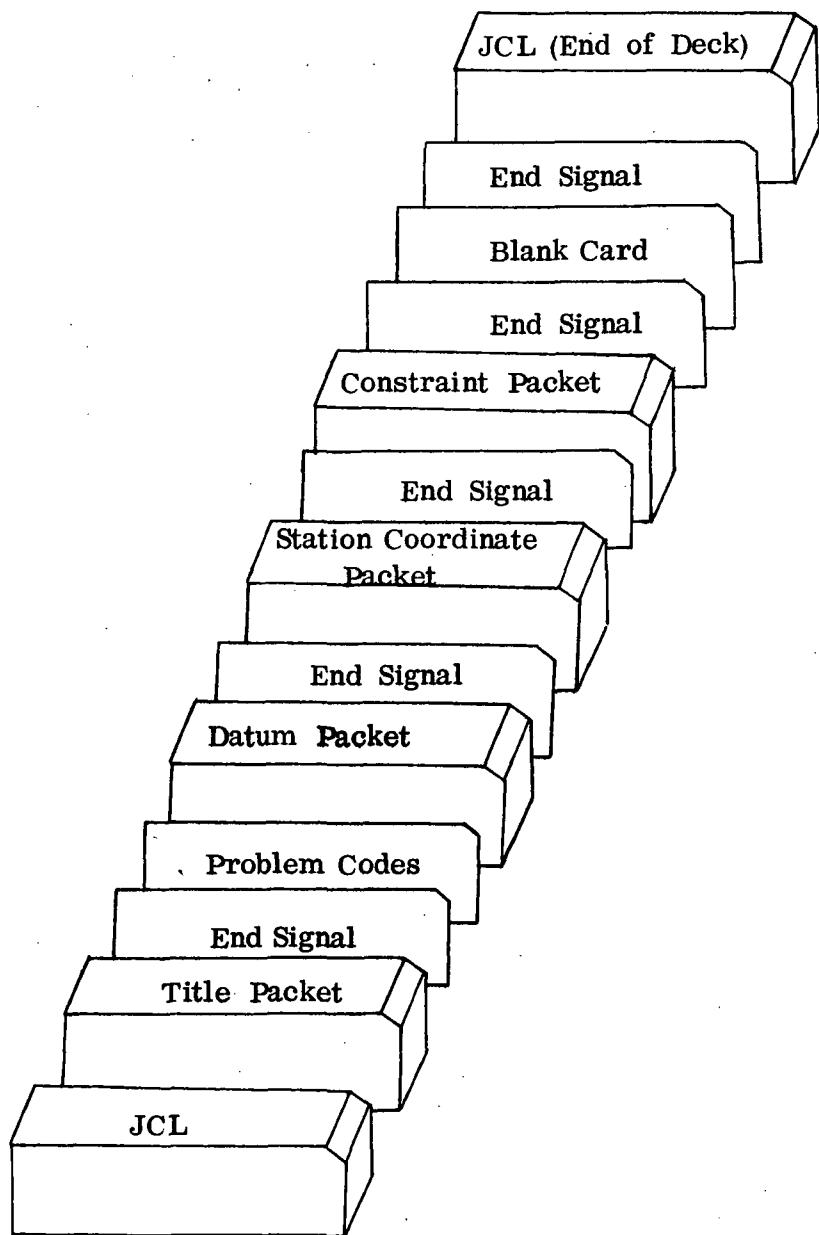


Figure 8. Deck Setup for Solution Using Constraints Only.

5. OUTPUT

As far as the geodetic analyst is concerned, the most important part of an adjustment is the final results which are given in the printed output from the computer. However, there is more information included in this printout than the solution vector itself, and for this reason an explanation may be necessary. In this section of the report, the entire printed output is explained in detail.

5.1 Output Common to All Adjustments.

The first few pages printed on any adjustment are identical in format regardless of the type of adjustment. These pages contain the information input to the program on punched cards giving the description of the job itself, the problem codes, the datums involved in the adjustment, and the input coordinates of stations. Figures 9, 10, and 11 are samples of the actual output for each of the items mentioned. It should be noted that at the bottom of Figure 11, there is written 'TEST DISTANCE = 5.00 SECONDS OF ARC'. This one line will differ for the different types of runs. The above is written for optical data. With range data in the geometric mode it will be 'TEST VARIANCE = some number', and for range observations in the orbital mode, it will be a printout of the coordinates of the center of mass and the uncertainties of these coordinates.

5.2 Output from Geometrical Adjustment.

5.2.1 Output of Optical Observations.

When running the optical adjustment, all events are printed out as shown in Figure 12. As can be seen, the iteration number is printed, then the test distance and then for each of the events, the adjustment information. This is referred to as an event adjustment because it is truly an adjustment for the satellite position. There will be a maximum of six lines of information for each 2 station event, as is illustrated in Figure 12.

SECUR SOUTHWEST PACIFIC DATA

THE GEOCENTRIC COORDINATES OF THE MAUI BAKER-NUNN IS CONSTRAINED AT 7 METERS IN EACH COMPONENT (X,Y,Z)

THE ELLIPSOIDAL HEIGHTS OF STATIONS 1 TO 7 ARE CONSTRAINED TO VALUES DETERMINED BY USING THE SAO GEOD MAP TOGETHER WITH THE LEVELED HEIGHTS AT EACH STATION

THE EGRS 6 DATA IN QUAD 3-8-10-11 FROM AMS IS INCLUDED

PROBLEM CODES

212011000000200010111

-38-

Figure 9.

DATUMS INVOLVED IN ADJUSTMENT

DATUM 1 SAC STANDARD EARTH 1969

A= 6378155.00 METERS

B= 6356770.10 METERS

Figure 10.

INPUT COORDINATES OF STATIONS

9001	URGAN PASS, N.M.	DATUM	1	SAC STANDARD EARTH 1969
	GEODETIC COORDINATES	32 25 24.7246	-	+253 26 48.3689 1611.5300
	CARTESIAN COORDINATES	-1535756.999	-5166995.995	3401042.003
9007	AR EQUIPA, PERU	DATUM	1	SAC STANDARD EARTH 1969
	GEODETIC COORDINATES	- 16 27 57.1295	-	+288 30 24.1819 2476.0400
	CARTESIAN COORDINATES	1942775.000	-5804081.001	-1796933.004
9009	CURACAO, ANTILLES	DATUM	1	SAC STANDARD EARTH 1969
	GEODETIC COORDINATES	12 5 24.7586	-	+291 9 44.1872 -27.9300
	CARTESIAN COORDINATES	2251828.999	-5816918.998	1327160.003

TEST DISTANCE = 5.000 SECONDS CF ARC

Figure 11.

BEGIN ITERATION 1

TEST DISTANCE = 5.00 SECONDS OF ARC

EVENT 1
9007 9 20 27.88550 9AUG65 17 3 15.9510 - 4 51 32.9900
9009 9 20 27.88550 9AUG65 16 46 40.8540 -39 42 38.7800
SATELLITE POSITION 909591.839 -9777429.886 -2145935.965
GEOD. COORD. OF SATELLITE -12.378295 275.314914 3674216.8
GQI = 0.32985 RMS MISCLOSURE IN METERS = 8.8

EVENT 2
9007 9 29 59.89010 9AUG65 16 49 43.6620 36 45 5.2200
9009 9 29 59.89010 9AUG65 16 34 17.2960 1 27 20.2100
SATELLITE POSITION 606301.075 -9926959.236 1439667.748
GEOD. COORD. OF SATELLITE 8.271373 273.495068 3671402.6
GQI = 0.33681 RMS MISCLOSURE IN METERS = 37.5

EVENT 3
9007 7 26 0.00370 12AUG65 21 21 29.6330 31 53 42.2100
9009 7 26 0.00370 12AUG65 21 19 44.5540 -6 37 8.1500
SATELLITE POSITION 5212034.937 -8549156.629 859722.350
GEOD. COORD. OF SATELLITE 4.928418 301.368777 3671504.9
GQI = 0.38921 RMS MISCLOSURE IN METERS = 1.6

Figure 12.

The first output line for each event is the number of the event. This numbering starts at 1 for the first event and continues on. The second line is the observational data from the first station in the event, plus the residual in seconds after the adjustment for the satellite position. Referring again to Figure 12, event 1, line 1, the information as one reads across the line is:

9007	Station Number.
9 20 27.8855	= $9^h 20^m 27.8855$ UT1
9 Aug 65	Date
17 3 15.9510	= $17^h 3^m 15.9510$ Right Ascension
-4 51 32.9900	= $-4^{\circ} 51' 32.9900$ Declination
2.00	Standard deviation in right ascension, multiplied by the cosine of the declination, in seconds of arc.
2.00	Standard deviation in declination, in seconds of arc.
0.00	Covariance between $\alpha \cos \delta$ and δ , in seconds of arc, squared.
0.4	The residual, in seconds of arc, after the ad- justment for the satellite position.

There will be one line of information for each observation. The information printed on lines 4 and 5 of event 1 give the satellite position in XYZ coordinates plus the geodetic coordinates φ , λ , h of the satellite. Either, or both, or neither of these two lines can be printed if the analyst so desires. This output is controlled by the value used for PCODE (11) (see Table 1).

The last line of each event gives a term referred to as GQI, and the RMS misclosure in meters. The term GQI stands for Geometric Quality Index, and is just the determinant of the matrix of normal equations used

in the event adjustment divided by the number of stations in the adjustment. It is used to give an idea of the conditioning of the matrix of normal equations; the smaller the GQI, the better the conditioning.

The events listed in Figure 12 are excellent examples of good data. However, not all data is good, and several examples of this are shown in Figure 13. As was mentioned earlier in this report, the optical program can reject observations and still give a satisfactory adjustment provided that after all rejections there are still good observations from at least two stations. In Figure 13, events 2279 and 2280 each have one observation rejected, which is denoted by the * printed at the end of the printed line. In both cases, the other two observations were good and the events were acceptable. At the bottom of Figure 13, events 2310 and 2311 were deleted due to insufficient number of good observations, which is the meaning of KODE = 2. If KODE = 3, it means that the deletion was due to insufficient geometrical separation between observations.

5.2.2 Output of Range Observations, Geometric Mode.

The output for the range observations, geometric mode, is almost identical in format to that of the optical observational data described earlier. A sample output is shown in Figure 14. The minimum number of stations required is 4. The adjustment for the satellite position is a least squares adjustment that iterates until convergence (maximum of 20 iterations). Referring to the second line of event 2, Figure 14, the information as one reads across the line is:

5401	Station Number.
66 July 3	Date (Notice that the order of the year and the day are the reverse of the optical printout.)
1 31 43.9990	1 ^h 31 ^m 43 ^s .9990 time
2164169.973	Range, in meters.
3.20	Standard deviation of the range measurement.
-1.12	The residual, in meters, of the range observation after the adjustment for the satellite position.

EVENT	2279	0 11 23.96570	17MAR66 14 43	34.4510	23 48 44.6300	2.00	2.00	0.0	0.9
8015	0 11 23.96570	17MAR66 14 20	59.5640	18 19 12.6700	2.00	2.00	0.0	0.9	
9066	0 11 23.96570	17MAR66 16 10	23.4330	29 55 54.4000	2.00	2.00	6.0	13.1*	
8004	0 11 23.96570	6328333.525	2143166.810	5475604.534					
SATELLITE POSITION									
GEOD. CONST. OF SATELLITE		39.474837	12.700203	22.68926.8					
GOI =	0.01753	RMS MISCLOSURE IN METERS =	11.3						
EVENT	2280	2 11 59.96550	17MAR66 11 1	14.3560	23 47 31.8500	2.00	2.00	0.0	0.7
8015	2 11 59.96550	17MAR66 12 0	35.4200	48 29 26.5100	2.00	2.00	0.0	18.6*	
9004	2 11 59.96550	17MAR66 11 7	57.7620	17 44 50.4200	2.00	2.00	0.0	0.6	
8066	2 11 59.96550	6503592.505	-1269492.659	5543847.906					
SATELLITE POSITION									
GEOD. CONST. OF SATELLITE		40.056820	348.954827	2270261.5					
GOI =	0.01162	RMS MISCLOSURE IN METERS =	9.7						
EVENT	2310	3 20 20.02100	17OCT66 10 24	42.5250	59 19 54.6300	2.00	2.00	0.0	11.3*
8015	3 20 20.02100	17OCT66 23 36	24.5500	70 5 58.6500	2.00	2.00	0.0	9.2*	
9051	3 20 20.02100	ENTIRE EVENT DELETED, KODE=	2						
GOI =	0.052580								
EVENT	2311	3 20 24.02100	17OCT66 10 30	9.5360	58 7 7.2500	2.00	2.00	0.0	11.4*
8015	3 20 24.02100	17OCT66 23 31	24.4550	70 43 59.3900	2.00	2.00	0.0	9.4*	
9051	3 20 24.02100	ENTIRE EVENT DELETED, KODE=	2						

Figure 13.

BEGIN ITERATION

1

TEST VARIANCE =

10.00

EVENT	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100	101	102	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122	123	124	125	126	127	128	129	130	131	132	133	134	135	136	137	138	139	140	141	142	143	144	145	146	147	148	149	150	151	152	153	154	155	156	157	158	159	160	161	162	163	164	165	166	167	168	169	170	171	172	173	174	175	176	177	178	179	180	181	182	183	184	185	186	187	188	189	190	191	192	193	194	195	196	197	198	199	200	201	202	203	204	205	206	207	208	209	210	211	212	213	214	215	216	217	218	219	220	221	222	223	224	225	226	227	228	229	230	231	232	233	234	235	236	237	238	239	240	241	242	243	244	245	246	247	248	249	250	251	252	253	254	255	256	257	258	259	260	261	262	263	264	265	266	267	268	269	270	271	272	273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294	295	296	297	298	299	300	301	302	303	304	305	306	307	308	309	310	311	312	313	314	315	316	317	318	319	320	321	322	323	324	325	326	327	328	329	330	331	332	333	334	335	336	337	338	339	340	341	342	343	344	345	346	347	348	349	350	351	352	353	354	355	356	357	358	359	360	361	362	363	364	365	366	367	368	369	370	371	372	373	374	375	376	377	378	379	380	381	382	383	384	385	386	387	388	389	390	391	392	393	394	395	396	397	398	399	400	401	402	403	404	405	406	407	408	409	410	411	412	413	414	415	416	417	418	419	420	421	422	423	424	425	426	427	428	429	430	431	432	433	434	435	436	437	438	439	440	441	442	443	444	445	446	447	448	449	450	451	452	453	454	455	456	457	458	459	460	461	462	463	464	465	466	467	468	469	470	471	472	473	474	475	476	477	478	479	480	481	482	483	484	485	486	487	488	489	490	491	492	493	494	495	496	497	498	499	500	501	502	503	504	505	506	507	508	509	510	511	512	513	514	515	516	517	518	519	520	521	522	523	524	525	526	527	528	529	530	531	532	533	534	535	536	537	538	539	540	541	542	543	544	545	546	547	548	549	550	551	552	553	554	555	556	557	558	559	560	561	562	563	564	565	566	567	568	569	570	571	572	573	574	575	576	577	578	579	580	581	582	583	584	585	586	587	588	589	590	591	592	593	594	595	596	597	598	599	600	601	602	603	604	605	606	607	608	609	610	611	612	613	614	615	616	617	618	619	620	621	622	623	624	625	626	627	628	629	630	631	632	633	634	635	636	637	638	639	640	641	642	643	644	645	646	647	648	649	650	651	652	653	654	655	656	657	658	659	660	661	662	663	664	665	666	667	668	669	670	671	672	673	674	675	676	677	678	679	680	681	682	683	684	685	686	687	688	689	690	691	692	693	694	695	696	697	698	699	700	701	702	703	704	705	706	707	708	709	710	711	712	713	714	715	716	717	718	719	720	721	722	723	724	725	726	727	728	729	730	731	732	733	734	735	736	737	738	739	740	741	742	743	744	745	746	747	748	749	750	751	752	753	754	755	756	757	758	759	760	761	762	763	764	765	766	767	768	769	770	771	772	773	774	775	776	777	778	779	780	781	782	783	784	785	786	787	788	789	790	791	792	793	794	795	796	797	798	799	800	801	802	803	804	805	806	807	808	809	810	811	812	813	814	815	816	817	818	819	820	821	822	823	824	825	826	827	828	829	830	831	832	833	834	835	836	837	838	839	840	841	842	843	844	845	846	847	848	849	850	851	852	853	854	855	856	857	858	859	860	861	862	863	864	865	866	867	868	869	870	871	872	873	874	875	876	877	878	879	880	881	882	883	884	885	886	887	888	889	890	891	892	893	894	895	896	897	898	899	900	901	902	903	904	905	906	907	908	909	910	911	912	913	914	915	916	917	918	919	920	921	922	923	924	925	926	927	928	929	930	931	932	933	934	935	936	937	938	939	940	941	942	943	944	945	946	947	948	949	950	951	952	953	954	955	956	957	958	959	960	961	962	963	964	965	966	967	968	969	970	971	972	973	974	975	976	977	978	979	980	981	982	983	984	985	986	987	988	989	990	991	992	993	994	995	996	997	998	999	1000
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Figure 14.

For the two events given in Figure 14, notice that PCODE (11) must have been zero (0) because of the fact that no satellite positions are printed.

The last line of each event gives the variance of the event adjustment σ_o^2 ($\sigma_o = V'PV/d.o.f.$), in meters squared. Also, the number of iterations of the least squares adjustment is given.

5.2.3 Output Common to All Geometric Adjustments.

After the events are printed in either the range or optical geometric adjustments, there are quite a few options, and all of these options are controlled by the PCODES. The basic options of course are; do you want to form normal equations? If you do form normal equations, do you want to perform a solution? Then there are the secondary options; do you want to print the normals? Do you want to punch the normals? Do you want to simulate the guide matrix? Do you want to perform a summary by observed lines? There are also several solution codes that are controlled by PCODE (16) thru PCODE (20). The analyst would be well advised to reread Table 1 to see just how all these options are initiated.

If there are no normal equations formed and no solution to be performed, the words 'NORMAL TERMINATION' are printed and the program stops. If normal equations are formed, and if a solution is to be performed, the first set of information that is printed is the analysis of misclosures by station and a summary of information as shown in Figure 15. It should be noted that these are the values prior to the addition of the constraints.

If PCODE (6) = 1, the guide matrix is printed. This is just a matrix to show at a glance what stations co-observed, and the arrangement of the matrix of normal equations. Figure 16 is a sample guide matrix. The number 999 printed at the end of each line in Figure 16 is just an indication of the end of a row. When the normal equations are generated in the computer, the number 999 is used to indicate the end of a row.

ANALYSIS OF MISCLOSURES BY STATION

STATION	NUMBER OF OBSERVATIONS	RMS MISCLOSURE
9001	8	C.0
9007	16	0.31
9009	16	C.31

TOTAL NUMBER OF GOOD OBSERVATIONS	32
TOTAL NUMBER OF GOOD EVENTS	8
CORRESPONDING DEGREES OF FREEDOM	8
TOTAL SUM OF SQUARES OF MISCLOSURES	3.06
CORRESPONDING STANDARD DEVIATION OF UNIT WEIGHT	0.62
WPN INCLUDING CONTRIBUTION FROM SATELLITE POSITION (I.E., VPV+UX)	2.96

Figure 15.

GUIDE MATRIX

5401	5401	5402	5403	5404	999				
5402	5402	5403	5404	5405	5406	5407	999		
5403	5403	5404	5405	5406	5407	5408	5410	5411	999
5404	5404	5405	5406	5407	999				
5405	5405	5406	5407	999					
5406	5406	5407	5408	999					
5407	5407	5408	5410	5411	999				
5408	5408	5410	5411	999					
5410	5410	5411	999						
5411	5411	999							

Figure 16.

If PCODE (7) = 1, the normal equations are printed. The description of the normal equations was described in Section 3.2.3. The printout would be in the format as shown in Figure 6.

If the inner adjustment constraints are used, the computer will print out one or more of the following messages:

"THE ORIGIN OF THE COORDINATE SYSTEM IS DEFINED BY
INNER ADJUSTMENT PROCEDURE."

"ORIENTATION OF THE COORDINATE SYSTEM DEFINED BY
INNER ADJUSTMENT PROCEDURE."

"SCALE OF THE COORDINATE SYSTEM DEFINED BY INNER
ADJUSTMENT PROCEDURE."

Another option available is the summary by observed lines, which can be had by setting PCODE (8) = 1. This is just a summation of the actual number of observations made by every two station combination in the network. Figure 17 is a sample listing.

If a solution is to be performed, a set of constraints must be included in the data deck as described in Section 4. When the constraints are included, they are printed out after the summary of observed lines.

Figure 18 is a sample of a partial listing. It is just a printout of the

OBSERVATIONS ON EACH LINE

2	6	27
2	7	14
2	8	23
2	9	19
2	10	22
6	7	31
6	8	23
6	9	19
6	10	21
7	8	14
7	9	27
7	10	21
8	9	14
8	10	23
9	10	27

Figure 17.

THE ELLIPSOIDAL COORDINATES (LAT., LNG., HEIGHT) OF STATION 36
ARE CONSTRAINED AT

38° 89' 33.5" 22.22 DEGREES 267.795033329 DEGREES 306.000 METERS

ON DATUM 1 NORTH AMERICAN

THE WEIGHTS FOR THESE CONSTRAINTS ARE COMPUTED FROM OBSERVATION STANDARD DEVIATIONS OF

0.100 SECONDS 0.100 SECONDS 5.000 METERS

CHORD DISTANCE CONSTRAINT IMPOSED BETWEEN STATION 27
AND STATION 29

CONSTRAINED DISTANCE = 1531562.90

THE WEIGHT IS COMPUTED FROM A RELATIVE UNCERTAINTY OF ONE PART IN 750000.00

Figure 18.

original constraints, but with enough printed titles and labels to make the printed output easily understood by anyone.

The description of the results after adjustment will be deferred until Section 5.4.

5.3 Output from Orbital Adjustment, Range Data.

The output from an orbital adjustment differs in every respect from the geometric adjustment, except for the final parameter output for each station. This is necessary because in essence each event of a geometric adjustment corresponds to one pass of orbital information, and naturally orbital passes contain many more observations, etc.

The first group of output common to all orbital adjustments is the input information itself as shown in Figure 19. This information is printed for every pass in the data deck.

The second group of output data is the results of the adjustment of each pass for the first iteration (see Figure 20). Notice that the orbital elements printed at the beginning of each pass are the Apparent Celestial Cartesian Coordinates. Regardless of the original set of orbital elements, the program converts to this system for the adjustment. If the original orbital elements are very close to the actual elements, the resulting misclosures will be very small. However, the usual case is that the approximate orbital elements will cause fairly large misclosures (Figure 20 is a typical example). However, if the observations are good, the second iteration will have very small residuals.

A good point to keep in mind when examining this particular portion of the printout is that the program has no way of rejecting bad observations in a pass, and even if an entire pass is bad, the program cannot reject it. It is up to the analyst to examine the misclosures for each observation as well as the RMS misclosures for the pass, and to physically remove from the data deck observations which are bad, or even entire

EARTH CONSTANTS FOR CRUTINIZATION

SEMI-MAJOR AXES

GM

ROTATION RATE

6378155.000

2.98601900000

14

0.72921151470-04

2 PASS 2

X= -0.583952160 07 Y= -0.328766450 07 Z= 0.407155310 07
 XDOT= 0.179101450 04 YDOT= 0.359149150 04 ZDOT= 0.581808990 04
 EPOCH= 0.402493990 05

VALUES STORED ON UNIT 3

X= -0.583952160 07 Y= -0.328766450 07 Z= 0.407155310 07
 XDOT= 0.179101450 04 YDOT= 0.359149150 04 ZDOT= 0.581808990 04
 EPOCH= 0.402493990 05
 28 JAN 69
 CAST= 0.472908960 01

STATION	DATE	TIME(UT)	OBSERVED	RANGE	UNCERTAINTY
4082	28 JAN 69	9 34 21.7000	2505094.64		2.00
4082	28 JAN 69	9 34 41.6998	2488615.81		2.00
4061	28 JAN 69	9 34 16.2000	2381823.98		2.00
4061	28 JAN 69	9 34 36.1998	2480479.17		2.00
4061	28 JAN 69	9 34 56.2000	2581702.52		2.00
4860	28 JAN 69	9 35 19.3793	1918628.62		2.00
4860	28 JAN 69	9 35 39.3791	1872374.74		2.00
4860	28 JAN 69	9 35 59.3791	1834132.55		2.00
4082	28 JAN 69	9 35 1.8000	2478700.29		2.00
4082	28 JAN 69	9 35 21.7998	2475526.67		2.00
4082	28 JAN 69	9 35 41.7999	247904		2.00
4061	28 JAN 69	9 35 16.1999	247904		2.00
4061	28 JAN 69	9 35 36.2000	356.26		2.00
4061	28 JAN 69	9 35 56.2000	383015.10		2.00
4860	28 JAN 69	9 36 1.8000	3470356.36		2.00
4860	28 JAN 69	9 36 18.82	3017246.32		2.00
4860	28 JAN 69	9 36 31.1888	3127873.84		2.00
		9 41 51.1987	3239424.59		2.00
		9 42 19.3769	2646703.51		2.00
4860	28 JAN 69	9 42 39.3762	2742840.17		2.00
4860	28 JAN 69	9 42 59.3759	2841280.66		2.00
4082	28 JAN 69	9 42 1.8000	3559865.74		2.00
4082	28 JAN 69	9 42 21.8000	3651357.54		2.00
4760	28 JAN 69	9 42 11.1885	3351769.98		2.00
4760	28 JAN 69	9 42 31.1877	3464788.52		2.00
4760	28 JAN 69	9 42 51.1876	3578380.17		2.00

Figure 19.

BEGIN ITERATION 1

2 PASS 2 EPOCH= 28 JAN 69 9H 33M 56.0000S UT=MJD 40249.398564815

CURRENT ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN COORDINATES
POSITION(METERS) -5839521.600 -3287664.500 4071553.100
VELOCITY(METERS/SEC) 1791.014500 3591.491500 5818.089900

STATION	DATE	TIME(UT)	CORRECTED RANGE	UNCERTAINTY	MISCLOSURE
4082	28 JAN 69	9 34 21.7000	2505094.64	2.00	-45.68
4082	28 JAN 69	9 34 41.6998	2488615.81	2.00	-53.92
4061	28 JAN 69	9 34 16.2000	2381823.98	2.00	-109.62
4061	28 JAN 69	9 34 36.1998	2480479.17	2.00	-113.16
4082	28 JAN 69	9 42 21.8000	3651357.54	2.00	-169.74
4760	28 JAN 69	9 42 11.1885	3351769.98	2.00	-282.57
4760	28 JAN 69	9 42 31.1877	3464788.52	2.00	-282.17
4760	28 JAN 69	9 42 51.1876	3578380.17	2.00	-283.08

PARTIAL UNCERTAINTIES OF ORBIT UNKNOWNs-FROM DDN(INV)
0.346249D 01 0.124982D 01 0.831489D 00 0.120211D-01 0.494939D-02 0.248116D-02

WEIGHTED SUM OF SQUARES OF MISCLOSURES = 414757.488
NUMBER OF OBSERVATIONS = 70
RMS MISCLOSURE = 76.975

Figure 20.

passes if the RMS misclosure is high. However, it is best to wait at least until the second iteration before removing data.

After the listing of a complete iteration, there are several items printed, these particular items being identical to that of a geometric adjustment. These are the Analysis of Misclosures by stations, guide matrix, normal equations, observations on each line, constraints, and the summary by observed lines. Samples of these can be seen in Figures 6, 15, 16, 17 and 18. As always, the analyst can repress the printing of some of these by using the proper PCODE.

After all of the above items have been printed (or repressed, as the case may be), there is a listing of corrections to orbit and error model unknowns for each pass (see Figure 21). This is extremely valuable, especially after the second iteration, to determine the quality of the orbits. This information should be used in conjunction with the listing shown in Figure 20, where at the bottom are given the uncertainties of these particular orbital elements.

After the corrections to orbit and error model unknowns are listed come the results of the adjustment, which will be described in Section 5.4. After that the next iteration begins (if there is a next iteration) and everything is repeated.

5.4 The Output of Adjusted Coordinated and Related Information.

The most important part of any adjustment is the adjusted coordinates of the parameters and the standard deviations of these adjusted coordinates. The printed output from this program gives all this information and more for each station in the network.

Prior to printing out the information for each station, a short tabulation is given listing degrees of freedom, $V'PV$, σ_o^2 , and σ_o . These values refer to the situation after the constraints have been added to the normal

CORRECTIONS TO ORBIT AND ERROR MODEL UNKNOWNs

2 PASS 2
CORRECTION VECTOR EPOCH= 28 JAN 69 9H 33M 56.0000S UT=MJD 40249.398564815
-143.67340819 -21.05024861 -166.81242405 -0.06976902 -0.03690959 -0.04738683

UPDATED ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN COORDINATES
POSITION(METERS) -5839665.273 -3287685.550 4071386.288
VELOCITY(METERS/SEC) 1790.955731 3591.454590 5818.042513

3 PASS 3
CORRECTION VECTOR EPOCH= 28 JAN 69 11H 23M 56.0000S UT=MJD 40249.474953704
-122.31373575 -88.52634812 -174.57801664 0.00583718 -0.04677484 -0.12510625

UPDATED ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN COORDINATES
POSITION(METERS) -6031421.514 -3762718.626 3248854.522
VELOCITY(METERS/SEC) 1111.462937 3194.186825 6233.512894

5 PASS 5
CORRECTION VECTOR EPOCH= 28 JAN 69 23H 18M 56.0000S UT=MJD 40249.971481481
99.79297248 98.72556723 -193.35861013 0.02146740 -0.07336664 0.14568621

UPDATED ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN COORDINATES
POSITION(METERS) 4614260.193 5197589.926 3550155.041
VELOCITY(METERS/SEC) 3545.638267 725.030533 -6111.083214

UPDATED COORDINATES OF THE CENTER OF MASS
-311.854 191.895 -563.219

Figure 21.

equations and the normal equations solved for the corrections to the parameters. A sample output is shown in Figure 22.

The output of the adjusted coordinates of the parameters is identical for every type of adjustment, a sample of which is shown in Figure 23. As can be seen, the results are self-explanatory. The adjusted coordinates and standard deviations are given in both the Cartesian coordinate system and in the φ , λ , h system with respect to the datum of adjustment. It should be mentioned here that the standard deviations are derived from the variance-covariance matrix multiplied by the value of σ_0^2 .

Additional information printed is the direction of eigenvectors and square roots of eigenvalues of variance-covariance matrix. This may or may not be useful to the analyst. Another group of information is the off-diagonal elements of the weight-coefficient matrix as well as the correlation coefficients.

A printout is made for each station in the network. After all stations are output, the program will return for another iteration, if there is to be another iteration. If this is the last iteration, the words 'NORMAL TERMINATION' will be printed as the last line of output.

NUMBER OF DEGREES OF FREEDOM =	271
QUADRATIC SUM OF ALL THE RESIDUALS (V _{PPV}) =	266.5961
VARIANCE OF UNIT WEIGHT =	0.9837
STANDARD DEVIATION OF UNIT WEIGHT =	0.9918

Figure 22.

STATION NUMBER -	5410	MIDWAY SEC GENC	ELLIPSOID 7	SAO 1969
PREL. COORD. -	X -5618729.5352	Y -258198.3722	Z 2997233.8079	LAT. 28 12 43.2370
CORRECTIONS -	-9.1203	-12.2942	11.6966	LONG. (+E) 182 37 51.8560
ADJ. COORD. -	-5618738.6555	-258210.6664	2997245.5046	ELL. HT. -9.2900

VARIANCE-COVARIANCE MATRIX OF THE STATION POSITION

91.331782	40.498980	-102.881694	0.052232	0.044362	2.647202
40.498980	65.061489	-68.062532	0.044362	0.083648	2.362280
-102.881694	-68.062532	150.026834	2.647202	2.362280	195.520361

STAND. DEV. - 9.5568 R.1155 12.2485

DIRECTIONS OF EIGENVECTORS AND SQUARE ROOTS OF EIGENVALUES OF VARIANCE-COVARIANCE MATRIX -

LATITUDE	LONGITUDE	ELEVATION	AZIMUTH	AXIS LENGTH
48 21 31.2589	-145 27 21.5953	58 14 44.4656	41 52 9.7999	16.1149
5 5 57.0727	118 47 2.7514	25 23 49.2719	- 98 13 26.8569	6.0566
41 11 6.4421	24 18 21.8179	- 17 45 28.8403	- 16 58 11.2579	3.2937

3X3 WEIGHT COEFFICIENT MATRICES

X	Y	Z	X	Y	Z	
STA. NO. 5410 WITH STA. NO. 5410	92.840499	41.167985	-104.581205	1.000000	0.522177	-0.878906
	41.167985	66.99460	-69.186862	0.522177	1.000000	-0.684712
	-104.581205	-69.186862	152.505139	-0.878906	-0.684712	1.000000

STA. NO. 5410 WITH STA. NO. 5411	-46.725036	-39.920902	-71.530350	-0.797291	-0.478285	-0.700819
	-0.624932	17.405108	-74.699018	-0.012557	0.245560	-0.861838
	35.241776	5.129640	114.152416	0.469193	0.047951	0.872624

6. ADDITIONAL FEATURES

The first five chapters of this report give the details of using the OSUGOP program. Although it has not been mentioned, it is possible to make changes to any or all of the subprograms that will not alter the subprograms on the disk. There have been many occasions where there was a need to modify certain parts of OSUGOP to perform a special type of adjustment. This is done by including the source version of the subprograms (with the appropriate changes made) as part of the deck setup. This requires a completely different set of JCL cards (see Appendix III). There have been other occasions where additional programs have been written to perform certain required tasks. These programs are run separately from OSUGOP, but the output can be used by OSUGOP.

Although there have been many modifications used, probably the most important is the ability to read more than one set of normal equations and to perform an adjustment using all normal equations. Another very real problem is the ability to input correlated observations. The following is a brief description of the ways to handle these problems.

6.1 Addition of Normal Equations.

It was mentioned in section 3.2.3.1 that different systems of normal equations can be combined, and a description was given as to how they should be combined. At OSU there are two different techniques for adding normal equations. The one is a modification to the subroutine RDSOLN and the other is a separate program.

The modification to RDSOLN is really just the addition of a DO LOOP that causes the program to keep reading normal equations. The deck setup is as shown in Figure 4 except that the degrees of freedom and $V'PV$, punched normal equations, and extra 999 card are repeated for each set of normal equations. At the end of the last set of normal equations the E card is inserted.

The separate program that adds normal equations has been called ADDITION. It adds normal equations together and then prints and punches the combined normal equations. In addition to this, different weights can be applied to the different sets of equations.

6.2 Forming Normal Equations Using Correlated Observations.

The OSUGOP program was written to form normal equations from uncorrelated observations. However, the observations from the NOAA BC-4 Worldwide Network are in the form of Greenwich Hour Angle and Declination for up to seven fictitious images from each camera plate. These observations are the result of a polynomial fit, and there is a full variance-covariance matrix for all of the observations. In order to use this data, which was recorded on 17 magnetic tapes, a special program was assembled to read these tapes and form the normal equations. The normal equations are compatible to those described earlier in this report and OSUGOP was used to perform the solution from the punched normal equations. This program is described in [Mueller, et al., in press].

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Krakiwsky, Edward J., J. Ferrier, and G. Blaha (1968). "Least Squares Adjustment of Satellite Observations for Simultaneous Direction or Ranges, Part 2 of 3: Computer Programs." Reports of the Department of Geodetic Science, No. 87, The Ohio State University, Columbus.

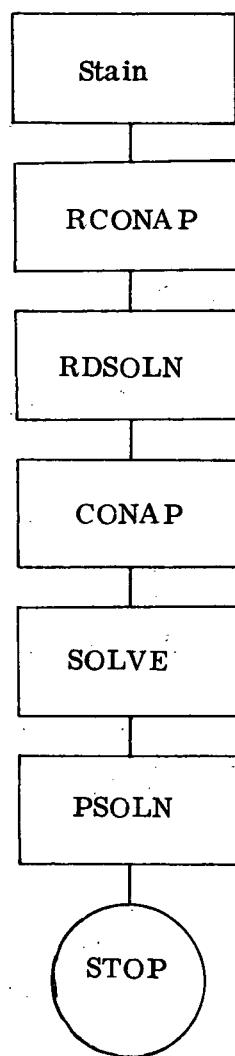
Krakiwsky, Edward J., J. Ferrier and J. P. Reilly (1967). "Least Squares Adjustment of Satellite Observations for Simultaneous Direction or Ranges, Part 3 of 3: Subroutines." Reports of the Department of Geodetic Science, No. 88, The Ohio State University, Columbus.

Mueller, Ivan I. (1969). Spherical and Practical Astronomy as Applied to Geodesy. Frederick Ungar Publishing Co., New York.

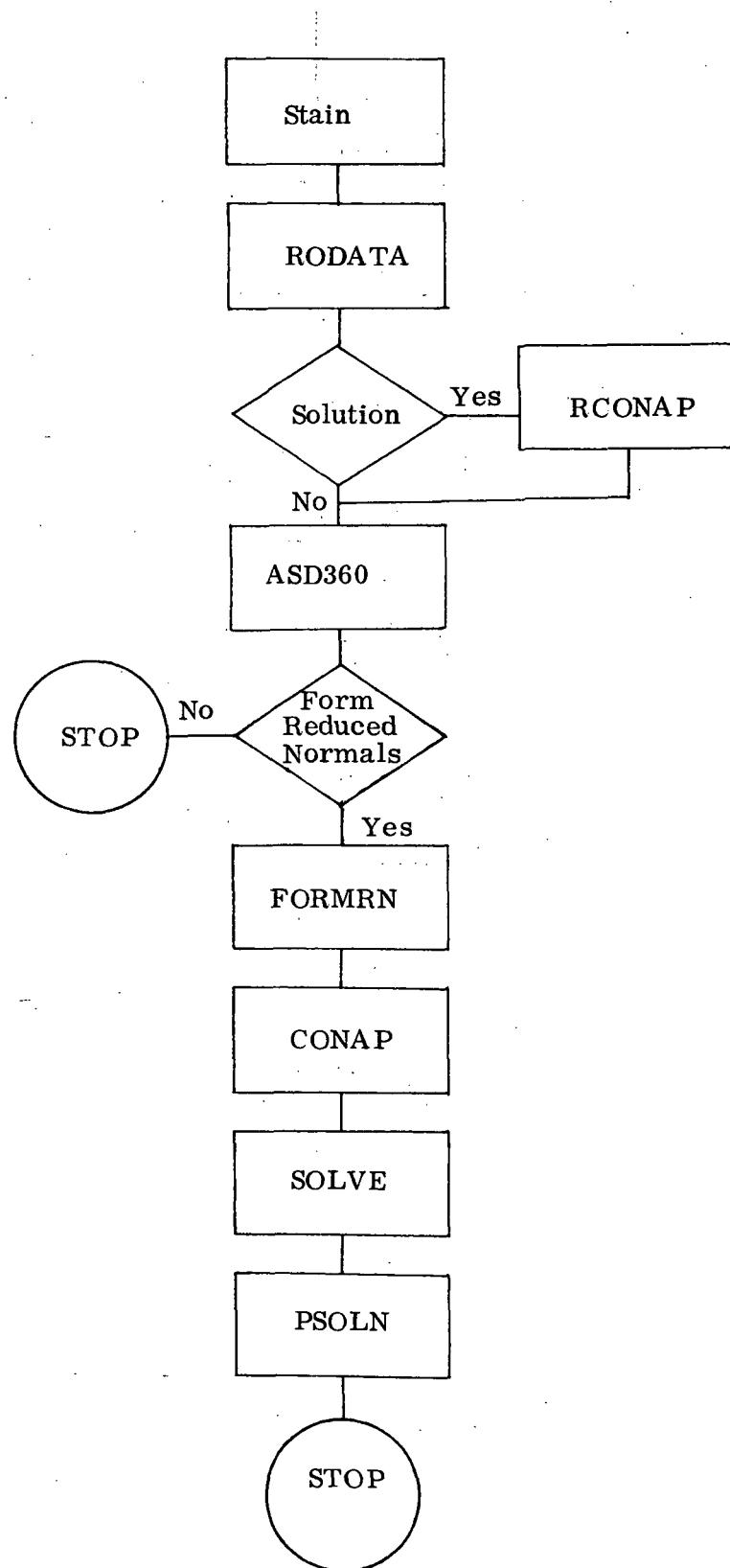
Mueller, Ivan I., J. P. Reilly, C. Schwarz and G. Blaha (1970). "SECOR Observations in the Pacific." Reports of the Department of Geodetic Science, No. 140, The Ohio State University, Columbus.

Mueller, Ivan I., M. Kumar, J. P. Reilly, and N. K. Saxena. (1973). "Analysis of the DOC/DOD Cooperative Worldwide Geodetic Satellite (BC-4) Network." Reports of the Department of Geodetic Science, No. 193, The Ohio State University, Columbus. (in press).

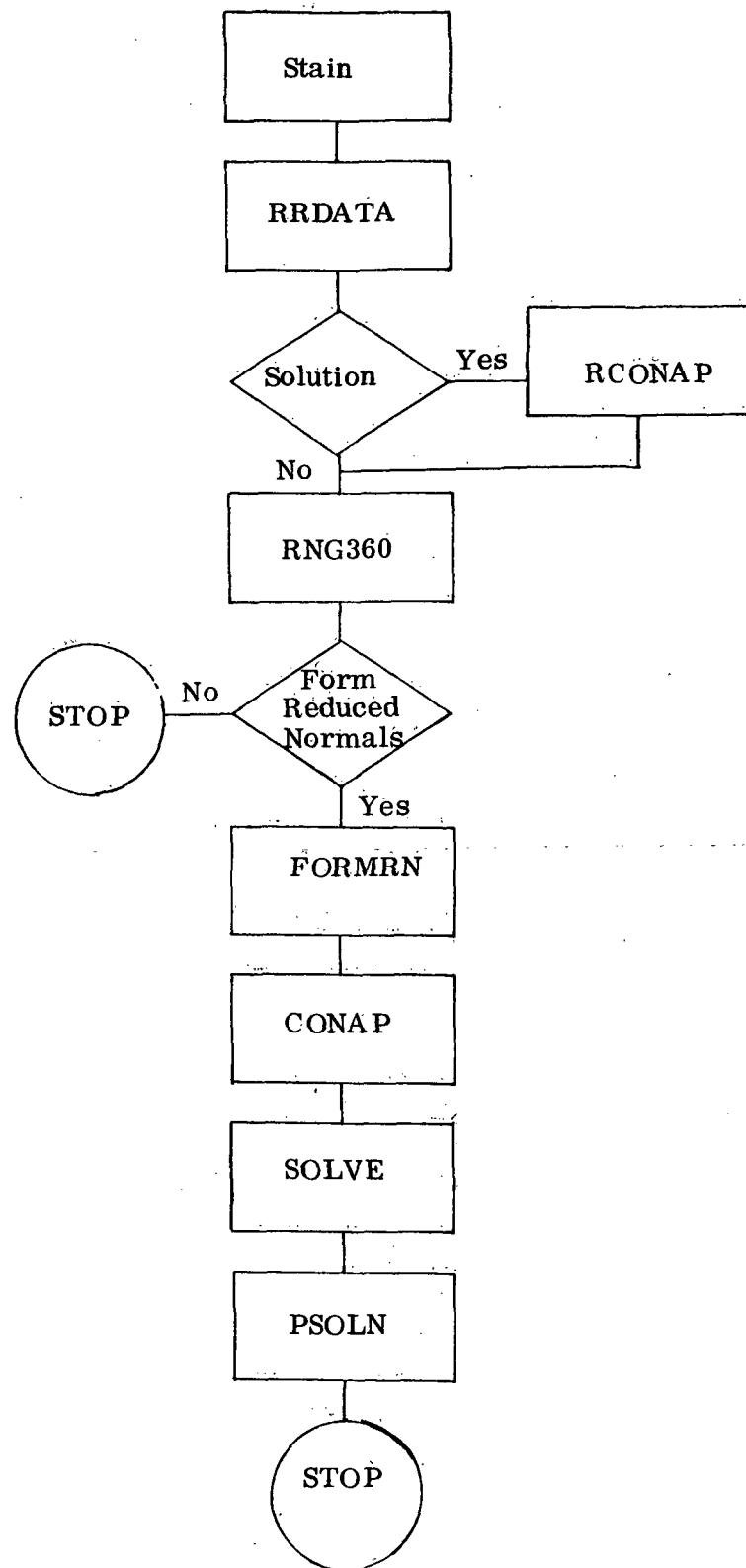
APPENDIX I
Flow Diagrams



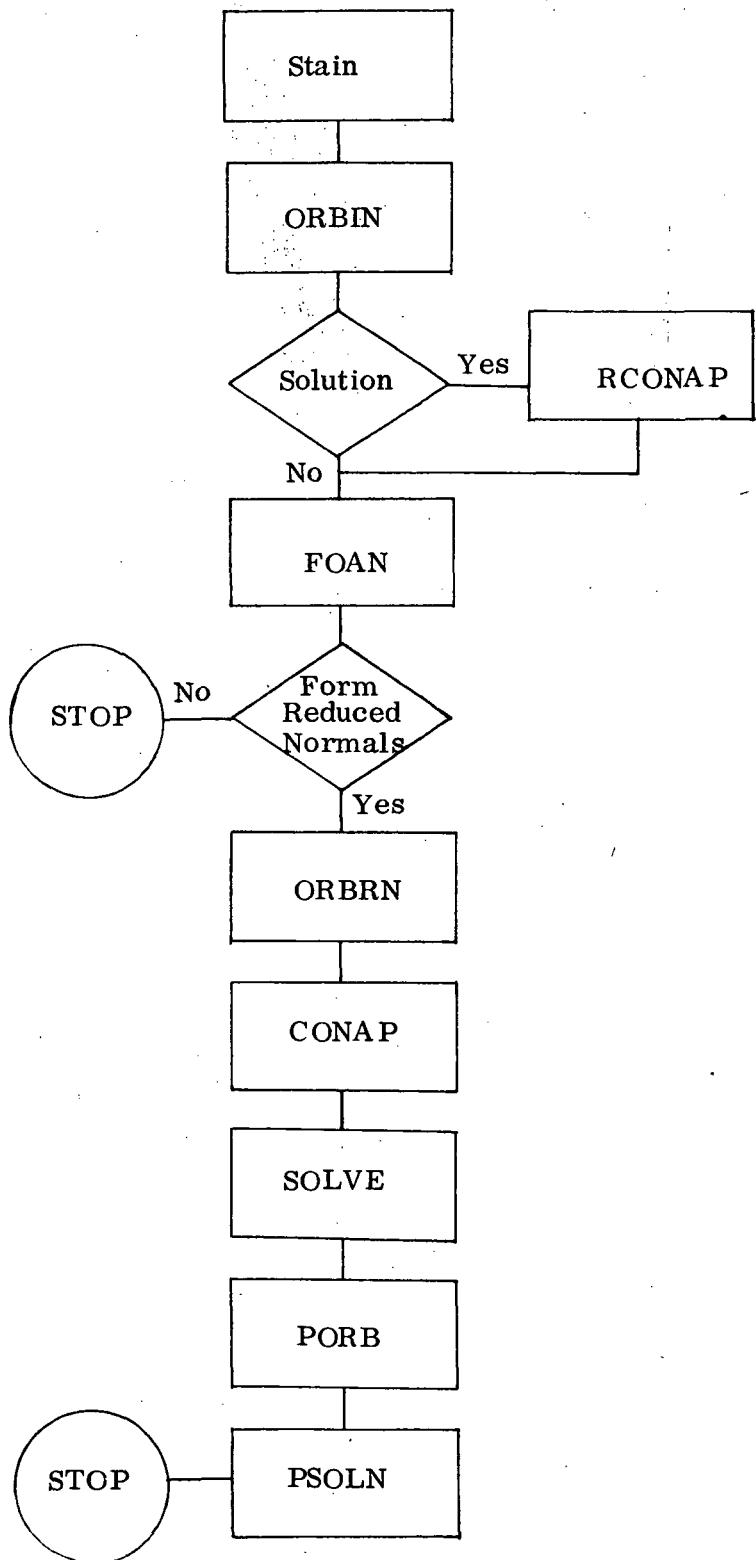
SOLUTION ONLY.



OPTICAL DATA, GEOMETRIC MODE.



RANGE DATA, GEOMETRIC MODE.



RANGE DATA, OPTICAL MODE.

APPENDIX II
Brief Description of Subprograms

Subprogram Listing

This is a listing of all subroutines and functions used in the OSUGOP program and a brief description of what is done by each. If the subprogram is a function, the letter F will appear in parenthesis after the name.

<u>Name</u>	<u>Purpose</u>
ANRADD (F)	Converts degrees, minutes, seconds into radians.
ASD 360	Processes optical directions. Reads data from disk, calls for the computation of satellite position, computes contribution to normal equations and contribution to $V'PV$.
CLEAR	Fills an array with Floating point zeros.
CONAP	Processes constraints on, and between, stations. Reads constraint codes, edits for wrong codes, calls CONAP1 and CONAP2.
CONAP1	Processes weighted constraints, adds contribution to normals, $V'PV$, and d.o.f.
CONAP2	Processes absolute constraints, adds contribution to normals, $V'PV$ and d.o.f.
DANG	Converts radians to degrees, minutes, seconds.
DEDIT	Edits optical data based on preliminary station positions and deletes bad observations and bad events, based on the test distance criteria.
DELL	Used to compute ΔX , ΔY , ΔZ . Also propagates the error from the cartesian coordinates to the Geodetic coordinates.
DPDOT (F)	Takes the dot product of two vectors.

DRIVER	Acts as the driving program for Orbit integration.
EXPAND	Expansion of Power Series coefficients, used for orbital runs.
FOAN	Forms normal equations for short arc mode processing.
FORMRN	Forms reduced normals for geometric mode processing.
GSTD (F)	Computation of Greenwich Sidereal Time
KEPEQ	Solves Kepler's Equation
ICEPTCE	Converts from Keplerian to Cartesian Orbit Elements.
KSID2 (F)	Stops the program if an observation is from a station not included in the list of input stations.
KSTAID (F)	Searches table of station identifiers for the internal number of a station.
MAIN	This is the driving program. Everything starts and stops here.
MATRUP	Updates the matrizant with respect to time. Used for orbital runs.
MJD(F)	Computed Modified Julian Day.
ORBIN	Orbit input subroutine.
ORBIT	Orbit integration controller.
ORBRN	Forms reduced normals for short arc mode processing.
POLE	Computes polar motions values x and y.
PORB	Prints updated orbital elements and Error Model Terms.
PRENUT	Computes precession and Nutation.
PSOLN	Prints the solution.
RCONAP	Reads the constraint cards and writes constraint information on a disk.
RDSOLN	Reads normal equations punched on cards and sets up storage for a solution.

RNG 360	Processes range measurements. Accepts or rejects events based on test variance. Computes contribution to normal equations, $V'PV$ and d. o. f.
RODATA	Reads the optical data cards, rotates into terrestrial coordinate system, puts all data onto a disk.
ROT3	Performs an R_3 rotation to a vector.
RRDATA	Reads Range data input cards, puts all data onto a disk.
SATXYZ	Computes the satellite position from approximate station coordinates and three or more range measurements.
SOLVE	Solves Normal Equations and computes inverse.
STAIN	Reads station coordinates and datum information from cards.
SWITCH	Switches rows and columns in a matrix.
UPDATE	Evaluates position and velocity at time $t + \Delta t$, given the position and velocity at time t . Used for Orbital Solutions.
UVWD	Converts geodetic to rectangular coordinates.
UVWTG	Converts rectangular to geodetic coordinates.
UVWTG2	Same as UVWTG. It is located in a different overlay.
UVWTG3	Same as UVWTG and UVWTG2. It is located in a different overlay.
VARIEQ	Generates the power series required to evaluate the matrix solution of variational equations. Used for orbital solutions.

APPENDIX III
JCL
(Job Control Cards)

JCL

```
// (10000,1000),CLASS=C
//STEP1 EXEC PROC=FORTRAN,TIME=(,20)
//CMP.SYSIN DD *
```

INSERT SOURCE SUBPROGRAMS HERE

```
/*
//STEP2 EXEC PGM=IEWL,PARM='MAP,LIST,OVLY,IDL',TIME=(0,20)
//MYLIB DD DSN=SCJ032.MUELLER,DISP=SHR
//SYSUT1 DD UNIT=SYSDA,DISP=(NEW,DELETE),SPACE=(CYL,(2,1))
//SYSLIB DD DSN=SYS1.FORTLIB,DISP=SHR
//          DD DSN=SYS2.FORTSSP,DISP=SHR
//SYSPRINT DD SYSOUT=A
//SYSLMOD DD DSNAME=&GD(MAIN),UNIT=SYSDA,SPACE=(CYL,(1,1,1)),
//          DISP=(NEW,PASS)
//SYSLIN DD DSNAME=*.STEP1.CMP.SYSLIN,DISP=(OLD,DELETE)
// DD * -
ALIAS      GEOMSG
INCLUDE    MYLIB(GEOMSG)
OVERLAY    ALPHA
INSERT     OBSD,MJD,PRENUT,POLE,GSTD,STAPLH,DPDOT,ANRADD,UVWTG2
INSERT     RCONAP,KSID2
INSERT     STAIN,UVND
OVERLAY    GAMMA
INSERT     RODATA,ASD360,DEDIT,DEDITC
OVERLAY    GAMMA
INSERT     RRDATA,RNG360,SATXYZ,RANGED
OVERLAY    GAMMA
INSERT     ERDCON,ORBCOM,ORRPAR,ORBIT,EXPAND,VARIO,UPDATE
INSERT     CLEAR
INSERT     ORBIN,KEPTCE,KEPEO,ROT3
INSERT     FOAN,DRIVER,MATRUP
OVERLAY    ALPHA
INSERT     NORMEO
OVERLAY    BETA
INSERT     DRBRN
OVERLAY    BETA
INSERT     FORMRN
OVERLAY    BETA
INSERT     RDSOLN
OVERLAY    BETA
```

JCL required when a subroutine (or subroutines) is different from that on the disk. This does not change the subroutines on the disk, but merely overrides them.

JCL (Continued)

```
INSERT      CONAP
OVERLAY    DELTA
INSERT      CONAP1,UVWTG3
OVERLAY    DELTA
INSERT      CONAP2
OVERLAY    BETA
INSERT      SOLVE,SWITCH
OVERLAY    ALPHA
INSERT      PORB
OVERLAY    ALPHA
INSERT      PSOLN,DANG,UVWTG,DELL,DEIGEN
/*
//GO EXEC PGM=*.STEP2.SYSLMOD,TIME=(04,20),REGION=252K
//FT01F001  DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//              DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT02F001  DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//              DCB=(LRECL=404,BLKSIZE=412,RECFM=VS)
//FT03F001  DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//              DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT04F001  DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//              DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT06F001  DD SYSOUT=A
//FT07F001  DD SYSOUT=B
//FT05F001  DD *
```

INSERT DATA HERE

```
/*
//
```

```
// (2000,200),CLASS=R
//JOBLIB DD DSNAME=SCJ032.MUELLER,DISP=SHR,PARM='1D'
//GO EXEC PGM=DSUG0P,TIME=2
//FT01F001 DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//           DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT02F001 DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//           DCB=(LRECL=408,BLKSIZE=412,RECFM=VS)
//FT03F001 DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//           DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT04F001 DD UNIT=SYSDA,SPACE=(CYL,(1,1)),DISP=(NEW,DELETE),
//           DCB=(LRECL=400,BLKSIZE=404,RECFM=VS)
//FT06F001 DD SYSOUT=A
//FT07F001 DD SYSOUT=B
//FT05F001 DD *
```

INSERT DATA HERE

```
/*
//
```

JCL required to use the standard program on the disk.

APPENDIX IV
Fortran IV Program with Subroutines

C
C PROBLEM CODE DEFINITIONS
C

C COLUMN MEANING

C 1. OVERALL PROBLEM CODE

C PCODE(1)=1 MEANS OPTICAL PROGRAM
C 2 MEANS RANGE
C 3 MEANS SOLUTION ONLY RUN
C 4 MEANS ORBITAL MODE, OPTICAL OBSERVATIONS
C 5 MEANS ORBITAL MODE, RANGE OBSERVATIONS
C 6 MEANS ORBITAL MODE, MIXED OBSERVATIONS
C PCODE(1)=7 MEANS OPTICAL PROGRAM, GEOMETRIC MODE(GEOS FORMAT)

C 2. PERFORM SOLUTION?

C PCODE(2)=1 MEANS YES
C 0 MEANS NO
C PCODE(1)=3 IMPLIES PCODE(2)=1

C 3. MAXIMUM NUMBER OF ITERATIONS?

C PCODE(1) MUST EQUAL 1 OR 2,
C PCODE(2) MUST EQUAL 1,
C PCODE(5) MUST EQUAL 1, FOR ONE OR MORE COMPLETE ITERATIONS

C 5. FORM NORMALS?

C PROCESSING CODES

C 1 MEANS YES, 0 MEANS NO

C 6. SIMULATE GUIDE MATRIX?

C 7. PRINT NORMALS?

C 8. PERFORM SUMMARY BY OBSERVED LINES?

C 9. PRINT NORMALS IN ASD FORMAT?

C 10. SUMMARIZE RESULTS

C PCODE(10)=0 DO NOT PRINT SUMMARY
C =1 PRINT THE OX'S AND STANDARD DEVIATIONS
C =2 PRINTS THE X,Y,Z'S AND STANDARD DEVIATIONS
C =3 PRINTS THE LATITUDE, LONGITUDE AND HEIGHT
C =4 PRINTS BOTH X,Y,Z & LAT., LONG, & H

C 11. PRINT SATELLITE POSITION FOR EACH EVENT?

C 0 MEANS NO
C 1 MEANS PRINT XYZ AND GEOMETRIC COORDINATES
C 2 MEANS PRINT XYZ ONLY
C 3 MEANS PRINT GEOMETRIC COORDINATES ONLY

C 12. THIS PARAMETER DESCRIBES WHERE THE STANDARD DEVIATIONS OF THE

C INDIVIDUAL OBSERVATIONS (USED TO FORM THE HEIGHTS) ARE TO BE FOUND
C PCODE(12)=0 MEANS TO READ THE OBSERVATIONAL STANDARD DEVIATION
C FROM THE CARD CONTAINING THE OBSERVATION.

C PCODE(12)=1 MEANS TO ASSOCIATE A SINGLE STANDARD DEVIATION WITH
C ALL OBSERVATIONS FROM A GIVEN STATION.** THE STANDARD DEVIATIONS
C TO BE ASSOCIATED WITH EACH STATION ARE GIVEN IN COLUMNS 73-79 OF
C THE CARD CONTAINING THE INPUT COORDINATES OF THE STATION.

C PCODE(12)=2 MEANS TO ASSOCIATE A SINGLE STANDARD DEVIATION WITH
C ALL OBSERVATIONS.** THIS NUMBER IS FOUND IN COLS. 21-30 OF THE
C CARD CONTAINING THE TEST DISTANCE (OPTICAL) OR TEST VARIANCE
C (RANGE).

C ** IN THE CASE OF OPTICAL OBSERVATIONS, THIS NUMBER IS INTERPRETED

```

C      AS THE STANDARD DEVIATION OF THE DECLINATION AND OF THE RIGHT
C      ASCENSION TIMES THE COSINE OF THE DECLINATION, AND THE
C      COVARIANCE IS SET TO ZERO.
C      CODES WHICH APPLY TO ORBITAL MODE PROCESSING ONLY
C      14. TREAT COORDINATES OF CENTER OF MASS AS UNKNOWN? (ORBITAL MODE 653Y)
C      15. PUNCH UPDATED ORBIT ELEMENTS? (ORBITAL MODE ONLY)
C
C      SOLUTION CODES
C
C      16. WRITE NORMALS AND INVERSE DURING SOLUTION PROCESSING?
C          0 MEANS PRINT NOTHING
C          1 MEANS PRINT PIVOT ELEMENTS
C          2 MEANS ALSO PRINT NORMALS AND INVERSE
C          3 MEANS ALSO PRINT REARRANGED NORMALS AND INVERSE
C      17. PUNCH ADJUSTED STATION XYZ AND VARIANCES FOR INPUT TO BADEKAS'
C          DATUM TRANSFORMATION PROGRAM?
C      18. PUNCH ADJUSTED STATION POSITIONS?
C      19. COMPUTE EIGENVECTORS OF VARIANCE-COVARIANCE MATRIX
C      20. COMPUTE CORRELATION COEFFICIENTS
C
C      COMMON/NINSTA/NSTA
C      INTEGER*2 ENDSIG/1HE/,CONTIN
C      INTEGER*2 PCODE(20)
C      COMMON/PCODES/PCODE
C      REAL*8 TITLE(10)
C      3 CONTINUE
C      WRITE(6,6001)
6001 FORMAT(1H1,20(/))
        4 READ(5,5001) TITLE,CONTIN
5001 FORMAT(9A8,A7,A1)
        IF(CONTIN.EQ.ENDSIG) GO TO 5
        WRITE(6,6012) TITLE
6012 FORMAT(30X,9A8,A7)
        GO TO 4
        5 CONTINUE
C
        READ(5,5050) PCODE
5050 FORMAT(80I1)
        WRITE(6,6050) PCODE
6050 FORMAT(///10X,'PROBLEM CODES',10X,20I1)
        CALL STAIN
        J=PCODE(1)
        GO TO (100,200,300,400,500,600,100),JCODE
100 CALL RODATA
        GO TO 105
200 CALL RRDATA
        GO TO 105
300 CALL ORBIF
        GO TO 105
105 IF(PCODE(2).EQ.1) CALL RCONMAP
        NITR=PCODE(3)

```

```
DO 110 I=1,MJTR
  WRITE(6,6010) I
6010 FORMAT(1H1//25X,'BEGIN ITERATION ',I5)
C      RUN OPTICAL
  IF(JCODE.EQ.1.OR.JCODE.EQ.7) CALL ASR360
C      RUN RANGE PROGRAM
  IF(JCODE.EQ.2) CALL RNC360
  IF(JCODE.EQ.5) CALL FOAM
  IF(PCODE(5).NE.1) GO TO 890
  IF(JCODE.LT.2.0R.JCODE.EQ.7) CALL FORTRAN
  IF(JCODE.GT.3.AND.JCODE.LT.7) CALL PDRPN
  ASSIGN 110 TO JRTN
  GO TO 800
110 CONTINUE
  GO TO 890
300 CONTINUE
  CALL RCONAP
  CALL RDSOLN
  ASSIGN 890 TO JRTN
  GO TO 800
800 CONTINUE
  CALL CONAP
  CALL SOLVE
  IF(JCODE.GT.3.AND.JCODE.LT.7) CALL PDRR
  CALL PSDLN
  GO TO JRTN,(110,890)
C
600 CONTINUE
400 CONTINUE
890 CONTINUE
  WRITE(6,6002)
6002 FORMAT(//1ONORIAL TERMINATION//1H1)
  STOP
  END
```

```
DOUBLE PRECISION FUNCTION DPDOT(X,Y,N)
DOUBLE PRECISION X(N),Y(N)
DPDOT=0.0
DO 10 I=1,N
10 DPDOT=DPDOT+X(I)*Y(I)
RETURN
END
```

```

SUBROUTINE FORMRN
IMPLICIT REAL*8(A-H,O-Z)
COMMON/NSTA/NSTA
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
COMMON/WPW/WPW,XPU,IDEGF,IFSTA
DIMENSION DDN(3,3),DDK(3),L1(3),L2(3),BNDDNI(3,3),TN(3,3),TK(3)
INTEGER*2 L,LSOLVE
INTEGER CONTIN,ENDSIG/1HE/
COMMON/STAORD/KORDER(150)
COMMON/NORMEQ/REDN(3,3,820),U(3,40),L(820),LSOLVE
DIMENSION BN(3,3,40),LG(40)
C FORM REDUCED NORMAL EQUATIONS FOR UP TO 40 STATIONS
DIMENSION KSTATE(40)
LOC(K)=(K*(K+1))/2
MAXSTA=40
IF(NSTA.GT.MAXSTA) GO TO 901
C
C THE REDUCED NORMAL EQUATIONS ARE STORED AS 3 X 3 BLOCKS IN THE ARRAY REDN.
C ONLY THE UPPER TRIANGULAR PART OF THE REDUCED NORMAL EQUATIONS IS STORED.
C THE BLOCKS OF THE REDUCED NORMAL EQUATIONS ARE NUMBERED
C ACCORDING TO THE FOLLOWING SCHEME:
C
C
C      1   2   4   7   11
C          3   5   8   12
C              6   9   13
C                  10  14
C                      15      ET CETERA
C
C L(820) IS THE GUIDE MATRIX
C L=1 SIGNIFIES A NON ZERO BLOCK
C L=0 SIGNIFIES A ZERO BLOCK
IB=LOC(NSTA)
DO 100 JB=1,IB
DO 99 I=1,3
DO 99 J=1,3
99 REDN(I,J,JB)=0.0
100 L(JB)=0
C
BACKSPACE 2
READ(2)      (((BN(I,J,KSTA),I=1,3),U(J,KSTA),J=1,3),
XKSTA=1,NSTA)
REWIND 2
C
C STASH DIAGONAL BLOCKS
DO 110 KSTA=1,NSTA
IB =LOC(KSTA)
DO 108 I=1,3
DO 108 J=1,3
108 REDN(I,J,IB)=BN(I,J,KSTA)
110 CONTINUE
C
FDEGF=IDEFG
IF(PCODE(9).EQ.1) WRITE(7,7010) FDEGF,WPW
7010 FORMAT(16X,2F16.6)
C READ BLOCKS FROM EACH EVENT AND REDUCE NORMAL EQUATIONS
C
150 READ(2)      NSTE,DDN,DDK,(((BN(I,J,IS),I=1,3),J=1,3),

```

```

1KSTATE(IS),IS=1,NSTE),CONTIN
C
DO 180 IS=1,NSTE
ISTA=KSTATE(IS)
IB=ISTA
CALL DGMPRD(BN(1,1,IS),DDN,BNDDNI,3,3,3)
CALL DGMPRD(BNDDNI,DDK,TK,3,3,1)
DO 155 I=1,3
155 U(I,ISTA)=U(I,ISTA)-TK(I)
DO 180 JS=1,NSTE
JSTA=KSTATE(JS)
JB=JSTA
C SKIP IF (ISTA.GT.JSTA), SINCE ONLY THE UPPER TRIANGULAR PART OF THE
C REDUCED NORMAL EQUATIONS IS BEING COMPUTED AND SAVED.
IF(ISTA.GT.JSTA) GO TO 180
C (IB,JB) GIVES THE ROW AND COLUMN NUMBER OF THE BLOCK IN THE REDUCED
C NORMAL EQUATIONS CURRENTLY BEING PROCESSED.
C
C SET INDICATOR
NB=LOC(JB-1)
NB=IB+NB
L(NB)=L(NB)+1
C PERFORM REDUCTION
CALL DGMPRD(BNDDNI,BN(I,1,JS),TN,3,3,3)
DO 130 I=1,3
DO 130 J=1,3
130 REDN(I,J,NB)=REDN(I,J,NB)-TN(I,J)
180 CONTINUE
C IF END OF DATA, GO OUT OF LOOP
IF(CONTIN.EQ.ENDSIG) GO TO 400
C IF NOT, RETURN TO PROCESS ANOTHER EVENT
GO TO 150
C
C ENTER HERE WHEN ALL EVENTS HAVE BEEN PROCESSED.
400 CONTINUE
C
C SIMULATE KRAKINSKI'S GUIDE MATRIX
IF(PCODE(6).NE.1) GO TO 441
C
WRITE(6,6001)
6001 FORMAT(1H1,10(/),20X,'GUIDE MATRIX')
DO 440 ISTA=1,NSTA
IB=0
LG(1)=1000
DO 435 JSTA=ISTA,NSTA
JB=LOC(JSTA-1)+ISTA
IF(L(JB).EQ.0) GO TO 435
IB=IB+1
LG(IB)=KORDER(JSTA)
435 CONTINUE
C
IB=IB+1
IF(IB.GT.1) LG(IB)=999
439 WRITE(6,6002) KORDER(ISTA),(LG(I),I=1,IB)
6002 FORMAT(20X,15,5X,18I5,200(/30X,18I5))
440 CONTINUE
441 CONTINUE
C

```

```

C PRINT NORMALS IN ASD FORMAT, AND PUNCH IF DESIRED.
  WRITE(6,6003)
6003 FORMAT(1H1//)
          NORMAL EQUATIONS (SEE GUIDE MATRIX) //)
  DO 450 ISTA=1,NSTA
  DO 442 I=1,3
  442 DDK(I)=-U(I,ISTA)
  IB=0
  JB=LOC(ISTA)
  IF(L(JB).GT.0) IB=1
C PUNCH NORMALS
  IF(PCODE(9).NE.1) GO TO 443
  WRITE(7,7001) KORDER(ISTA)
7001 FORMAT(14I5)
  WRITE(7,7006) DDK
7006 FORMAT(3(F16.10,5X))
  WRITE(7,7008) ((REDN(I,J,JB),J=1,3),I=1,3)
7008 FORMAT(3F16.10/3F16.10/3F16.10)
C
  443 CONTINUE
C PRINT DIAGONAL BLOCK
  IF(PCODE(7).NE.1) GO TO 444
  WRITE(6,6004) KORDER(ISTA)
6004 FORMAT(//15)
  WRITE(6,6006) DDK
6006 FORMAT(/3(F16.10,5X))
  WRITE(6,6008) ((REDN(I,J,JB),J=1,3),I=1,3)
6008 FORMAT(3F16.10)
  444 CONTINUE
C PRINT OFF-DIAGONAL BLOCKS
  KSTA=ISTA+1
  IF(ISTA.EQ.NSTA) GO TO 448
  DO 445 JSTA=KSTA,NSTA
  JB=LOC(JSTA-1)+ISTA
  IF(L(JB).EQ.0) GO TO 445
  IB=IB+1
  IF(PCODE(9).NE.1) GO TO 7445
  WRITE(7,7001) KORDER(JSTA)
  WRITE(7,7008) ((REDN(I,J,JB),J=1,3),I=1,3)
7445 CONTINUE
  IF(PCODE(7).NE.1) GO TO 445
  WRITE(6,6004) KORDER(JSTA)
  WRITE(6,6008) ((REDN(I,J,JB),J=1,3),I=1,3)
445 CONTINUE
448 I=1000
  IF(IB.GT.0) I=999
  IF(PCODE(7).EQ.1) WRITE(6,6004) I
  IF(PCODE(9).EQ.1) WRITE(7,7001) I
450 CONTINUE
  IF(PCODE(8).NE.1) GO TO 478
  WRITE(6,6010)
6010 FORMAT(10(/),20X,'OBSERVATIONS ON EACH LINE')
  IB=NSTA-1
  DO 475 ISTA=1,IB
  KSTA=ISTA+1
  DO 475 JSTA=KSTA,NSTA
  WRITE(6,6011) KORDER(ISTA),KORDER(JSTA),L(LOC(JSTA-1)+ISTA)
6011 FORMAT(8I10)
  475 CONTINUE

```

```
478 CONTINUE
RETURN
901 CONTINUE
WRITE(6,9001) MAXSTA,NSTA
9001 FORMAT(' FORMRN IS PRESENTLY DIMENSIONED TO HANDLE ONLY',I5,
10 ' UNKNOWN STATIONS.'/20X,' THIS PROBLEM HAS',I5,' UNKNOWN STATI
2IONS.'/10X,'EXECUTION IS TERMINATED BY PROGRAM.')
STOP
END
```

```

SUBROUTINE RDSOLN
C THIS SUBROUTINE READS THE NORMAL EQUATION FROM CARDS AND SETS UP
C STORAGE FOR A SOLUTION, SIMULATING STORAGE AFTER EXECUTION OF FORMRN.
C THIS SUBROUTINE IS CALLED ONLY FOR A SOLUTION-ONLY RUN.
C THE INPUT IS COMPATABLE WITH KRAKISKI'S SOLUTION PROGRAM.
C THE STATION COORDINATES MUST BE INITIALIZED BY CALLING STAIN.
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 L,LSOLVE
INTEGER ENDSIG/1HE/,CONTIN
COMMON/NSTA/NSTA/STAORD/KORDER(150)/WPW/WPW,XPU,IDEFG,IFSTA
COMMON/NORMEQ/REDN(3,3,820),U(3,40),L(820),LSOLVE
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
LOC(K)=(K*(K+1))/2
C
NB=LOC(NSTA)
DO 10 IB=1,NB
L(IB)=0
DO 10 I=1,3
DO 10 J=1,3
10 REDN(I,J,IB)=0.0
DO 12 IB=1,NSTA
DO 12 I=1,3
12 U(I,IB)=0.0
C
READ(5,5001) VUW,FDEGF,WPW
5001 FORMAT(3D16.8)
IDEFG=FDEGF
C
C READ NORMALS AND STORE
100 READ(5,5002) ID,CONTIN
5002 FORMAT(15,74X,A1)
IF(ID.EQ.999) CONTIN=ENDSIG
IF(CONTIN.EQ.ENDSIG) GO TO 600
ISTA=KSTAID(ID)
IF(ISTA.LE.0) GO TO 900
C
C READ CONSTANT COLUMN
READ(5,5003) (U(I,ISTA),I=1,3)
C
C SWITCH SIGN
DO 110 I=1,3
110 U(I,ISTA)=-U(I,ISTA)
C
C READ DIAGONAL BLOCK
NB=LOC(ISTA)
READ(5,5004) ((REDN(I,J,NB),J=1,3),I=1,3)
5003 FORMAT(3(D16.8,5X))
5004 FORMAT(3D16.8)
L(NB)=1
C
C READ OFF-DIAGONAL BLOCK
150 READ(5,5002) ID
IF(ID.EQ.999) GO TO 100
JSTA=KSTAID(ID)
IF(JSTA.LE.0) GO TO 900
C
C SWITCH SUBSCRIPTS IF NECESSARY SO THAT STORAGE IS MADE IN UPPER TRIANGULAR P
C PART OF REDUCED NORMAL EQUATIONS.
IF(JSTA.GE.ISTA) GO TO 160
NB=LOC(ISTA-1)+JSTA
READ(5,5004) ((REDN(J,I,NB),J=1,3),I=1,3)

```

```

GO TO 170
160 CONTINUE
NB=LOC(JSTA-1)+ISTA
READ(5,5004) ((REDN(I,J,NB),J=1,3),I=1,3)
170 CONTINUE
6001 FORMAT(3I5)
L{NB}=1
GO TO 150
C
900 WRITE(6,6000) ID
6000 FORMAT('0STATION NUMBER NOT FOUND IN INPUT LIST',16,' PROGRAM STOP
1S.')
STOP
C
600 CONTINUE
IF(PCODE(7).EQ.0) GO TO 620
WRITE(6,6003)
6003 FORMAT(//T30,'NORMAL EQUATIONS')
DO 615 ISTA=1,NSTA
WRITE(6,6002) KORDER(ISTA),ISTA
6002 FORMAT(//3I10)
WRITE(6,6004) (U(I,ISTA),I=1,3)
DO 615 JSTA=ISTA,NSTA
NB=LOC(JSTA-1)+ISTA
WRITE(6,6002) KORDER(ISTA),KORDER(JSTA),NB
WRITE(6,6004) ((REDN(I,J,NB),J=1,3),I=1,3)
6004 FORMAT(3F16.10)
615 CONTINUE
620 CONTINUE
RETURN
END

```

```

SUBROUTINE STAIN
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
INTEGER ENDSIG/1HE/,CONTIN
COMMON/NSTA/NSTA
COMMON/STAORD/KORDER(150)
INTEGER STANAM,IDS*2
INTEGER*2 PLUS/1H+/
INTEGER*2 ISGNP,IPHID,IPHIM,LONGD,LONGM,ISGNL
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1STANAM(5,150),IDS(150)
DIMENSION UNCE(3)
C UNCE FOR THE MOMENT IS A DUMMY ARRAY
COMMON/STAPLH/STAPLH(2,150)
COMMON/OBSD/OBSD(150),OVOBSD
MAXSTA=150
WRITE(6,6000)
6000 FORMAT(1H1)
6001 FORMAT(1H1,20(/))
WRITE(6,6002)
6002 FORMAT(////4X,29HDATUMS INVOLVED IN ADJUSTMENT,//)
C INPUT DATUMS
10 READ(5,5002) IDD,AE,BE,CONTIN
5002 FORMAT(12,2F12.3,53X,A1)
IF(CONTIN.EQ.ENDSIG) GO TO 30
DATPRM(1,IDD)=AE
DATPRM(2,IDD)=BE
READ(5,5003)(DATNAM(I,IDD),I=1,4)
5003 FORMAT(4A8)
WRITE(6,6003) IDD,(DATNAM(I,IDD),I=1,4),(DATPRM(I,IDD),I=1,2)
6003 FORMAT(6HODATUM,I3,3X,4A8,3HA= ,F10.2,12H METERS B= ,F10.2,
17H METERS)
GO TO 10
C
30 CONTINUE
C STATION INPUT
WRITE(6,6005)
6005 FORMAT(1H1//40X,29HINPUT COORDINATES OF STATIONS)
KSTA=0
35 KSTA=KSTA+1
READ(5,5005) IDD, IDTS, (STANAM(I,KSTA),I=1,5), ISGNP,IPHID,IPHIM,PHIS
1,LONGD,LONGM,FLONGS,H,UNCE,OBSD(KSTA),CONTIN
5005 FORMAT(14,I2,4A4,A2,A1,2(2I3,F8.4), F10.2,3F3.1,F7.2,A1)
IF(CONTIN.EQ.ENDSIG) GO TO 50
PHI=ANRADD(ISGNP,IPHID,IPHIM,PHIS)
ISGNL=PLUS
FLONG=ANRADD(ISGNL,LONGD,LONGM,FLONGS)
KORDER(KSTA)=IDD
IDS(KSTA)=IDTS
STAPLH(1,KSTA)=PHI
STAPLH(2,KSTA)=FLONG
CALL UVWD(DATPRM(1, IDTS),DATPRM(2, IDTS),PHI,FLONG,H,STAUVW(1,KSTA)
1,STAUVW(2,KSTA),STAUVW(3,KSTA))
WRITE(6,6006) IDD,(STANAM(I,KSTA),I=1,5),IDTS,(DATNAM(I, IDTS),I=1,4
1),ISGNP,IPHID,IPHIM,PHIS,ISGNL,LONGD,LONGM,FLONGS,H
6006 FORMAT(1H0,I4,8X,4A4,A2,10X,5HDATUM,I4,4X,4A8/10X,20HGEODETIC COOR
1DINATES,2(6X,A1,2I3,F8.4),F12.4)

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```
      WRITE (6,6007) (STAUVW(I,KSTA),I=1,3)
6007 FORMAT(10X,21HCARTESIAN COORDINATES,3F16.3)
      GO TO 35
50 CONTINUE
      NSTA=KSTA-1
      NSTAUN=3*NSTA
      REWIND 3
      RETURN
      END
```

```

SUBROUTINE ASD360
C S/360 VERSION OF ASD PROGRAM FOR OPTICAL SATELLITE DIRECTIONS
  IMPLICIT REAL*8(A-H,O-Z)
  INTEGER*2 PCODE(20)
  COMMON/PCODES/PCODE
  INTEGER*4 ENDSIG/1HE/,CONTIN,DELCOD(2)/1H .1H*/,ECODE
  INTEGER*2 PLUS/1H+/
  INTEGER*2 ISGNP,IPHID,IPHIM,LONGD,LONGM,ISGNL
  INTEGER*2 ID(50),KEY(50),IHR(50),MIN(50),IDAY(50),IYR(50),IRAH(50)
  1,IRAM(50),ISGND(50),IDEC(50),IDECM(50),IDAT(50,11)
  COMMON/DEDITC/ALFS(50),DEC(50),U(3,50),S(3),D(50),SDC(3,50),EVSUM,
  1GAST,STAXYZ(3,50),GQI,
  2TD,KSTATE(50),IPASS(50),NSTE,NSUSED,ECODE
  COMMON/NSTA/NSTA
  INTEGER STANAM,IDS*2
  DIMENSION MONTH(50)
  COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
  1STANAM(5,150),IDS(150)
  COMMON/STAORD/KORDER(150)
  EQUIVALENCE(ID(1),IDAT(1,1)),(KEY(1),IDAT(1,2)),(IHR(1),IDAT(1,3))
  1,(MIN(1),IDAT(1,4)),(IDAY(1),IDAT(1,5)),(IYR(1),IDAT(1,6)),(IRAH(1
  2),IDAT(1,7)),(IRAM(1),IDAT(1,8)),(ISGND(1),IDAT(1,9)),(IDEC(1),ID
  3AT(1,10)),(IDECM(1),IDAT(1,11))
  DIMENSION SEC(50),RAS(50),DECS(50),VARRA(50),VARDEC(50),COVRAD(50)
  1,DAT(50,6)
  EQUIVALENCE(SEC(1),DAT(1,1)),(RAS(1),DAT(1,2)),(DECS(1),DAT(1,3)),
  1(VARRA(1),DAT(1,4)),(VARDEC(1),DAT(1,5)),(COVRAD(1),DAT(1,6))
  DIMENSION DN(3,3,150),BN(3,3,50),DDN(3,3),DK(3,150),DDK(3),A(2,3),
  1W(2,2),DL(2)
  DIMENSION PM(3,3),AP(2,3)
  DIMENSION L1(3),L2(3),TA(3)
  COMMON/WPW/WPW,XPU,IDEKF,NFSTA
  DIMENSION NOBSTA(150)
  REAL*4 VPVSTA(150)
  MAXSTE=50
  SPR=206264.80625D0
  PI=3.14159265358D0
  PI2=2.0*PI
  RPD=180.0/PI
  WPWSP=0.0
C
  REWIND 2
  REWIND 3
  READ(3) TD
  WRITE(6,6004) TD
  6004 FORMAT(//20X,'TEST DISTANCE =',F20.2,'    SECONDS OF ARC')
C
C   START DATA INPUT
  DO 70 KSTA=1,NSTA
  NOBSTA(KSTA)=0
  VPVSTA(KSTA)=0.0
  DO 70 I=1,3
  DK(I,KSTA)=0.0
  DO 70 J=1,3
  DN(I,J,KSTA)=0.0
  70 CONTINUE
  KEVENT=0
  EPR=0.0

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```

210 CONTINUE
  READ (3) IEVENT,NSTE,GAST,PM,EPR,
  1((IDAT(IS,J),J=1,11),MONTH(IS),(DAT(IS,J),J=1,6),ALFS(IS),DEC(IS),
  2KSTATE(IS),IS=1,NSTE),CONTIN
  DO 270 IS=1,NSTE
  KSTA=KSTATE(IS)
  CALL DGMPRD(PM,STAUVW(1,KSTA),STAXYZ(1,IS),3,3,1)
270 CONTINUE
  WRITE(6,6008) IEVENT
6008 FORMAT(/ 1X,'EVENT',16)
C
  CALL DEDIT
C
  DO 280 IS=1,NSTE
280 WRITE(6,6010) ID(IS),KEY(IS),IHR(IS),MIN(IS),SEC(IS),IDAY(IS),
  1MONTH(IS),IYR(IS),IRAH(IS),IRAM(IS),RAS(IS),ISGND(IS),IDECD(IS),
  2IDECM(IS),DECS(IS),VARRA(IS),VARDEC(IS),COVRAD(IS),D(IS),
  3DELCOD(IPASS(IS))
  6010 FORMAT(I7,A1,2I3,F9.5,3X,I3,A3,I2,2I3,F8.4,3X,A1,I2,I3,F8.4,
  15X,3F6.2,F10.1,2X,A1)
C
  IF(ECODE.GT.1) GO TO 630
  IF(PCODE(11)) 610,630,610
  610 IF(PCODE(11)-3) 611,612,611
  611 WRITE(6,6024) S
  6024 FORMAT(' SATELLITE POSITION',3F15.3)
  IF(PCODE(11)-2) 612,630,612
  612 IDTS=IDS(KSTATE(1))
  CALL UVWTG2(S,DATPRM(1, IDTS),PHI,FLAM,H)
  PHI=PHI*RPD
  FLAM=FLAM*RPD
  WRITE(6,6023) PHI,FLAM,H
  6023 FORMAT(' GEOD. COORD. OF SATELLITE',2F14.6,F14.1)
  630 CONTINUE
  WRITE(6,6012) GQI
  6012 FORMAT(10X,'GQI=',F10.5)
  IF(ECODE.GT.1) GO TO 290
  IF(NSUSED.EQ.0) GO TO 290
  RMSMC=DSQRT(EVSUM/DFLOAT(NSUSED))
  WRITE(6,6011) RMSMC
  6011 FORMAT(1H+,27X,'RMS MISCLOSURE IN METERS=',F10.1)
  GO TO 300
  290 WRITE(6,6015) ECODE
  6015 FORMAT(1H+,27X,'ENTIRE EVENT DELETED, KODE=',I4)
  GO TO 600
C
C      SET UP OBSERVATION EQUATIONS FOR THIS EVENT AND COMPUTE CONTRIBUTIONS
C      TO THE NORMAL EQUATIONS
  300 CONTINUE
  IF(ECODE.GT.1) GO TO 600
  KEVENT=KEVENT+1
  DO 310 I=1,3
  DDK(I)=0.0
  DO 310 J=1,3
  DDN(I,J)=0.0
  310 CONTINUE
C
  JS=0

```

```

DO 390 IS=1,NSTE
IF(IPASS(IS).GT.1) GO TO 390
JS=JS+1
C      JS IS THE COUNTER FOR NON-DELETED STATIONS IN THE EVENT
RSQCSD=SDC(1,IS)**2+SDC(2,IS)**2
RSQ=RSQCSD+SDC(3,IS)**2
RCD=DSQRT(RSQ)
ASC=DATAN2(SDC(2,IS),SDC(1,IS))+GAST
DSC=DATAN(SDC(3,IS)/RCD)
DL(1)=ALFS(IS)-ASC
IF(DL(1).GT.PI) DL(1)=DL(1)-PI2
IF(DL(1).LT.(-PI)) DL(1)=DL(1)+PI2
DL(2)=DEC(IS)-DSC
C
COMPUTE WEIGHTS
VARRA(IS)=(VARRA(IS)/SPR)**2
VARRA(IS)=VARRA(IS)*RSQ/RSQCSD
VARDEC(IS)=(VARDEC(IS)/SPR)**2
COVRAD(IS)=COVRAD(IS)/SPR**2
DET=VARRA(IS)*VARDEC(IS)-COVRAD(IS)**2
W(1,1)=VARDEC(IS)/DET
W(2,2)=VARRA(IS)/DET
W(1,2)=-COVRAD(IS)/DET
W(2,1)=W(1,2)
C
COMPUTE OBSERVATION EQUATIONS
A(1,1)=-SDC(2,IS)/RSQCSD
A(1,2)=+SDC(1,IS)/RSQCSD
A(1,3)=0.0
FACTOR=SDC(3,IS)/(RSQ*RCD)
A(2,1)=-SDC(1,IS)*FACTOR
A(2,2)=-SDC(2,IS)*FACTOR
RANGE=DSQRT(RSQ)
A(2,3)=RCD/RSQ
CALL DGMprd(A,PM,AP,2,3,3)
6017 FORMAT(1X,7D17.9)
C
KSTA=KSTATE(IS)
C      ELIMINATE DELETED STATIONS FROM THE LIST OF STATIONS INVOLVED IN
C      THE EVENT.
KSTATE(JS)=KSTATE(IS)
C
COMPUTE VPV OF MISCLOSURES
VPVTO=0.0
DO 315 II=1,2
DO 315 JJ=1,2
315 VPVTO=DL(II)*W(II,JJ)*DL(JJ)+VPVTO
NOBSTA(KSTA)=NOBSTA(KSTA)+2
VPVSTA(KSTA)=VPVSTA(KSTA)+VPVTO
COMPUTE CONTRIBUTION TO NORMALS
DO 330 I=1,3
DO 325 J=1,3
TERM=0.0
DO 320 II=1,2
DO 320 JJ=1,2
320 TERM=TERM+AP(II,I)*W(II,JJ)*AP(JJ,J)
BN(I,J,JS)=-TERM
DN(I,J,KSTA)=DN(I,J,KSTA)+TERM

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      DDN(I,J)=DDN(I,J)+TERM
  325 CONTINUE
C
      TERM=0.0
      DO 328 II=1,2
      DO 328 JJ=1,2
  328 TERM=TERM+AP(II,I)*W(II,JJ)*DL(JJ)
      DK(I,KSTA)=DK(I,KSTA)-TERM
      DDK(I)=DDK(I)+TERM
  330 CONTINUE
C
  390 CONTINUE
C INVERT DDN
      DET=1.0
      CALL DMINV(DDN,3,DET,L1,L2)
      CALL DGMPRD(DDK,DDN,TA,1,3,3)
      CALL DGMPRD(TA,DDK,TB,1,3,1)
      WPWSP=WPWSP+TB
      NSUSED=JS
      WRITE(2) NSUSED,DDN,DDK,(((BN(I,J,JS),I=1,3),J=1,3),KSTATE(JS),
      1JS=1,NSUSED),CONTIN
  600 CONTINUE
C
C TEST FOR END OF INPUT
      IF(CONTIN.EQ.ENDSIG) GO TO 700
      GO TO 210
C
C
C
  700 CONTINUE
C
CHECK TO SEE IF END SIGNAL HAS BEEN WRITTEN ON DATA SET FT02
      IF(ECODE.EQ.1) GO TO 710
      BACKSPACE 2
C READ AND REWRITE LAST RECORD FROM LAST GOOD EVENT
      READ (2) NSUSED,DDN,DDK,(((BN(I,J,JS),I=1,3),J=1,3),KSTATE(JS),
      1JS=1,NSUSED)
      BACKSPACE 2
      WRITE(2) NSUSED,DDN,DDK,(((BN(I,J,JS),I=1,3),J=1,3),KSTATE(JS),
      1JS=1,NSUSED),CONTIN
  710 CONTINUE
      WRITE(2) (((DN(I,J,KSTA),I=1,3),DK(J,KSTA),J=1,3),
      XKSTA=1,NSTA)
C      WRITE(6,6018)(KORDER(KSTA),((DN(I,J,KSTA),J=1,3),I=1,3),
      1KSTA=1,NSTA)
  6018 FORMAT((I5/3(3D18.7/)))
      WPW=0.0
      NOBS=0
      WRITE(6,6019)
  6019 FORMAT(1H1,8(/),10X,'ANALYSIS OF MISCLOSURES BY STATION'//,
      1T10,'STATION',T20,'NUMBER OF OBSERVATIONS',T50,'RMS MISCLOSURE')
      DO 750 KSTA=1,NSTA
      NOBS=NOBS+NOBSTA(KSTA)
      WPW=WPW+VPVSTA(KSTA)
      RMSMC=0.0
      IF(NOBSTA(KSTA).GT.0) RMSMC=DSQRT(VPVSTA(KSTA)/DFLOAT(NOBSTA(KSTA))
      1)
      WRITE(6,6020) KORDER(KSTA),NOBSTA(KSTA),RMSMC

```

```
6020 FORMAT(T10,I7,T35,I7,T50,F14.2)
750 CONTINUE
  IDEGF=NOBS-3*KEVENT
  RMSMC=DSQRT(WPW/DFLOAT(IDEgf))
  WRITE(6,6021) NOBS,KEVENT,IDEgf,WPW,RMSMC
6021 FORMAT(///10X,'TOTAL NUMBER OF GOOD OBSERVATIONS',T60,I8//,
110X,'TOTAL NUMBER OF GOOD EVENTS',T60,I8//,
210X,'CORRESPONDING DEGREES OF FREEDOM',T60,I8//,
310X,'TOTAL SUM OF SQUARES OF MISCLOSURES',T60,F11.2//,
410X,'CORRESPONDING STANDARD DEVIATION OF UNIT WEIGHT',T60,F11.2)
  WPW=WPW-WPWSP
  WRITE(6,6022) WPW
6022 FORMAT(1H0,9X,'WPW INCLUDING CONTRIBUTION FROM SATELLITE POSITION',
1/15X,'(I.E., VPV+UX)',T60,F11.2)
  RETURN
  END
```

```

SUBROUTINE EXPAND (XPO,YPO,ZPO,CNM,SNM,LCT,ICT,UMT,VMT,CTB,CTT
1,ERD,XMU,ALF,OMG,ECC,NTE,KTR,KDR,NHT,CDC,CTH,KEY,DMT,KRG,CMC)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION LT(5),KT(5)
DIMENSION XPO(1),YPO(1),ZPO(1),CNM(1),SNM(1)
1, LCT(1),ICT(1),UMT(1),VMT(1),CTB(1),CTT(1)
DIMENSION DMT(1)
DIMENSION CMC(3)
COMMON INTERNAL WORKING ARRAYS
DIMENSION
1 XTL(8),CLB(8),CLT(8),RPT(8),RMO(8),RZR(8),VRB(8),
2 YTL(8),SLB(8),SLT(8),RMT(8),RMR(8)
COMMON /ORBCOM/
1CSQ(8),SSQ(8),SCS(8),BXB(8),BYB(8),BZB(8),XVR(8),YVR(8),ZVR(8),
2 QAV(8),QBV(8),QCV(8),AMT(72),FEE(72),CGB(8),SGB(8)
EQUIVALENCE
1 (CLB(1),CSQ(1)),(CLT(1),XVR(1)),(RPT(1),BXB(1)),(RZR(1),QAV(1))
2,(SLB(1),SSQ(1)),(SLT(1),YVR(1)),(RMT(1),BYB(1)),(RMR(1),QBV(1))
3,(XTL(1),SCS(1)),(YTL(1),ZVR(1)),(RMO(1),BZB(1)),(VRB(1),QCV(1))
COMMON DATA BLOCK END
EQUIVALENCE (LA,LT(1)),(LB,LT(2)),(LC,LT(3)),(K,LT(4)),(L,LT(5))
EQUIVALENCE (KA,KT(1)),(KB,KT(2)),(KC,KT(3)),(M,KT(4)),(N,KT(5))
C
CGB(1)=DCOS(ALF)
SGB(1)=DSIN(ALF)
CLB(1)=XPO(1)*CGB(1)+YPO(1)*SGB(1)
SLB(1)=YPO(1)*CGB(1)-XPO(1)*SGB(1)
RPT(1)=XPO(1)*XPO(1)+YPO(1)*YPO(1)+ZPO(1)*ZPO(1)
RMT(1)=1.0/RPT(1)
RMO(1)=DSQRT(RMT(1))
CLT(1)=ERD*RMT(1)*CLB(1)
SLT(1)=ERD*RMT(1)*SLB(1)
RZR(1)=ERD*RMT(1)*ZPO(1)
RMR(1)=ERD*ERD*RMT(1)
DRX=0.0
DRY=0.0
DRZ=0.0
NG = 0
1TR=3
KA=1
IF (KEY-2) 10,10,490
COMPUTE CTB,CTT ARRAYS TO MATCH TABLE ALLOCATION
10 CONTINUE
COMPUTATION OF CTB AND CTT ARRAYS MOVED TO DATA STATEMENT IN ORBIT
GO TO (60,150),KEY
COMPUTE ICT TABLE ENTRIES AND TOTAL ARRAY LENGTH
60 CONTINUE
KEY=3
NET=1
KTR=10
LNT=10
LA=LCT(2)
LB=LCT(1)
LC=LA
NA=NTE+1
DO 130 I=1,NA
KA=LA-1
KB=LB

```

```

KC=LC+1
  IF (KB-KA) 70,80,80
70  KB=KA
80  IF (LC) 110,110,90
90  IF (KB-KC) 100,110,110
100 KB=KC
110 LNTH=LNTH+KB+1
  ICT(I)=LNTH
  LA=LB
  LB=LC
  LC=LCT(I+2)
  IF (I-NTE) 130,120,120
120 LC=0
130 CONTINUE
  LNTH=LNTH-1
  KA=1
  GO TO 490
150 CONTINUE
  KEY=3
  NET=2
  LNTH=1
  DO 160 I=1,3
160 LT(I+2)=LCT(I)
  IF (NTE-2) 170,180,190
170 K=0
180 L=0
190 LB=K
  LA=L
  NA=NTE+2
  DO 280 J=1,NA
  DO 210 I=1,5
  KT(I)=0
  IF (LT(I)) 210,210,200
200 KT(I)=LT(I)+I-2
210 CONTINUE
  KK=0
  DO 230 I=1,5
  IF (KK-KT(I)) 220,230,230
220 KK=KT(I)
230 CONTINUE
  LNTH=LNTH+KK+1
  ICT(J)=LNTH
  DO 260 I=1,4
260 LT(I)=LT(I+1)
  L=LCT(J+3)
  N=J-NTE+2
  IF (N) 280,270,270
270 L=0
280 CONTINUE
  LNTH=LNTH-1
  KA=1
  GO TO 490
C BEGIN THE ITERATIVE INTEGRATION FOR THE SOLUTION.
C 300 SERIES - COMPUTE X,Y,Z COEFFICIENTS
300 KA=1+(ITR-3)*LNTH
  KB=1
  TMA=0.0
  TMB=0.0

```

```

TMC=0.0
TMD = 0.5
KG = 0
NA = ICT(1)-2
NB = 1
NC = NA+2
DO 380 I=1,NTE
KC = LCT(I)
IF (KC)380,380,310
310 IF (I-2)330,315,320
315 NA = 1
GO TO 325
320 NA = ICT(I-2)
325 NB = ICT(I-1)
NC = ICT(I)
330 LA = KA+NA+1
LB = KA+NB
LC = KA+NC-1
DO 375 J=1,KC
TME = CTB(KB)
TMF = TME
IF (KG)345,345,350
345 TMF = TMD*TME
TME = -TMF
TMI = J
TMJ = J+2
TMD = TMD*TMI/TMJ
350 TME = TME*UMT(LA)
TMF = TMF*VMT(LA)
TMG = UMT(LC)
TMH = VMT(LC)
TMI = CNM(KB)
TMJ = SNM(KB)
TMK = TME-TMG
TML = TMF+TMH
TMM = TMF-TMH
TMN = TME+TMG
TMO = UMT(LB)
TMP = VMT(LB)
IF (KRG)370,370,365
365 DMT(NG+1) = TMK
DMT(NG+2) = TML
DMT(NG+3) = TMO
DMT(NG+4) = TMM
DMT(NG+5) = TMN
DMT(NG+6) = TMP
NG = NG+6
370 TMA = TMA+TMI*TMK+TMJ*TMM
TMB = TMB-TMI*TML+TMJ*TMN
TMC = TMC+CTT(KB)*(TMI*TMO+TMJ*TMP)
LA = LA+1
LB = LB+1
LC = LC+1
KB = KB+1
375 CONTINUE
380 KG = 1
TMD = 0.5/ERD
KK=ITR-2

```

```

XTL (KK)=TMA*TMD
YTL (KK)=TMB*TMD
TMC=TMC*TMD
TMD=KK*(KK+1)
KA=KK
TMA=DRX
TMB=DRY
TMC=TMC+DRZ
DO 390 I=1,KK
TME=CGB(I)
TMF=SGB(I)
TMG=XTL(KA)
TMH=YTL(KA)
TMA=TMA+TME*TMG-TMF*TMH
TMB=TMB+TMF*TMG+TME*TMH
390 KA=KA-1
TMA=TMA-OMG*OMG*(CMC(1)*CGB(KK)-CMC(2)*SGB(KK))
TMB=TMB-OMG*OMG*(CMC(1)*SGB(KK)+CMC(2)*CGB(KK))
XPO(ITR)=TMA/TMD
YPO(ITR)=TMB/TMD
ZPO(ITR)=TMC/TMD
ITR=ITR+1
IF (KTR-ITR)400,410,410
400 RETURN
C 400 SERIES - COMPUTE EXTENSION DERIVATIVES
410 TMD=KK
KA=KK+1
CGB(KA)=-OMG*SGB(KK)/TMD
SGB(KA)=OMG*CGB(KK)/TMD
KK=KA
KB=KK
TMA=0.0
TMB=0.0
TMC=0.0
DO 420 I=1,KK
TME=XPO(KB)
TMF=YPO(KB)
TMG=CGB(I)
TMH=SGB(I)
TMA=TMA+TME*TMG+TMF*TMH
TMB=TMB+TMF*TMG-TME*TMH
TMC=TMC+TME*XPO(I)+TMF*YPO(I)+ZPO(KB)*ZPO(I)
420 KB=KB-1
CLB(KA)=TMA
SLB(KA)=TMB
RPT(KA)=TMC
KB=KK-1
RMT(KA)=0.0
TMA=0.0
DO 430 I=2,KK
TMA=TMA-RMT(KB)*RPT(I)
430 KB=KB-1
RMT(KA)=RMT(1)*TMA
KB=KK-1
TMA=RMT(KA)
KC=KB
IF (KC-1)460,460,440
440 DO 450 I=2,KC

```

```

        TMA=TMA-RMO(KB)*RMO(I)
450 KB=KB-1
460 RMO(KA)=TMA/(2.0*RMO(1))
        KB=KK
        TMA=0.0
        TMB=0.0
        TMC=0.0
        DO 470 I=1,KK
        TMI=RMT(KB)
        TMA=TMA+TMI*CLB(I)
        TMB=TMB+TMI*SLB(I)
        TMC=TMC+TMI*ZPO(I)
470 KB=KB-1
        CLT(KA)=TMA*ERD
        SLT(KA)=TMB*ERD
        RMR(KA)=RMT(KA)*ERD*ERD
        RZR(KA)=TMC*ERD
C 500 SERIES - COMPUTE U,V ARRAY EXTENSION COEFFICIENTS
490 CONTINUE
        NA=LNTH*(ITR-3)+1
        UMT(NA)=XMU*RMO(KA)
        VMT(NA)=0.0
        LA=1
        LB=1
        KB=2
        NB=NTE+NET
        DO 580 I=1,NB
        KC=ICT(I)-1
        M=I-1
        N=M
        IF (KC-KB)550,500,500
500 KK=2
        DO 540 J=KB,KC
        LC=LB
        NC=KA
        TMA=0.0
        TMB=0.0
        TMC=0.0
        TMD=0.0
        TME=N+N+1
        TMF=N+M
        TMG=N-M+1
        DO 530 K=1,KA
        GO TO (510,520),KK
510 TMI=RMR(NC)
        TMC=TMC+TMI*UMT(LC-1)
        TMD=TMD+TMI*VMT(LC-1)
520 TMJ=RZR(NC)
        TMA=TMA+TMJ*UMT(LC)
        TMB=TMB+TMJ*VMT(LC)
        NC=NC-1
530 LC=LC+LNTH
        NA=NA+1
        UMT(NA)=(TME*TMA-TMF*TMC)/TMG
        VMT(NA)=(TME*TMB-TMF*TMD)/TMG
        LB=LB+1
        N=N+1
540 KK=1

```

```
550 IF (I-NB)560,580,580
560 NA=NA+1
    TMC=2*M+1
    NC=KA
    TMA=0.0
    TMB=0.0
    DO 570 K=1,KA
    TMG=CLT(NC)
    TMH=SLT(NC)
    TMI=UMT(LA)
    TMJ=VMT(LA)
    TMA=TMA+TMG*TMI-TMH*TMJ
    TMB=TMB+TMG*TMJ+TMH*TMI
    LA=LA+LNTH
570 NC=NC-1
    UMT(NA)=TMA*TMC
    VMT(NA)=TMB*TMC
    KB=KC+2
    LB=LB+1
    LA=LB
580 CONTINUE
    IF (KDR) 300,300,600
  CODING FOR DRAG COMPUTATIONS ARE OMITED
  600 GO TO 300
  END
```

```

SUBROUTINE VARIEQ(XPO, YPO, ZPO, CNM, SNM, LCT, ICT, UMT, VMT, UVM, ERD, ALF,
*OMG, CDC, CTW, NTE, KTR, KDR)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION TMT(9)
DIMENSION XPO(1), YPO(1), ZPO(1), CNM(1), SNM(1), LCT(1), ICT(1), UMT(1),
*VMT(1), UVM(1)
COMMON INTERNAL WORKING ARRAYS
  DIMENSION XTL( 8),CLB( 8),CLT( 8),RPT( 8),RMO( 8),RZR( 8),VRB( 8),
* YTL( 8),SLB( 8),SLT( 8),RMT( 8),RMR( 8)
  COMMON/ORBCOM/CSQ( 8),SSQ( 8),SCS( 8),BXB( 8),BYB( 8),BZB( 8),
1 XVR( 8),YVR( 8),ZVR( 8),QAV( 8),QBV( 8),QCV( 8),
2 AMT( 72),FEE( 72),CGB( 8),SGB( 8)
  EQUIVALENCE(CLB(1),CSQ(1)),(CLT(1),XVR(1)),(RPT(1),BXB(1)),(RZR(1
*),QAV(1)),(SLB(1),SSQ(1)),(SLT(1),YVR(1)),(RMT(1),BYB(1)),(RMR(1),
*QBV(1)),(XTL(1),SCS(1)),(YTL(1),ZVR(1)),(RMO(1),BZB(1)),(VRB(1),QC
*V(1))
COMMON DATA BLOCK END
  EQUIVALENCE(TMT(1),TMA),(TMT(2),TMB),(TMT(3),TMC)
  EQUIVALENCE(TMT(4),TMD),(TMT(5),TME),(TMT(6),TMF)
  EQUIVALENCE(TMT(7),TMG),(TMT(8),TMH),(TMT(9),TMI)
C
C INITIALIZE ALL OF THE EXTERNAL ARRAYS.
C
  CALLCLEAR(TMT,9)
  IF(KTR-2)90,90,20
20 KA=KTR-2
  DO40KC=1,KA
    KB=KC
    DO30I=1,KC
      AU=CGB(KB)
      BU=SGB(I)
      TMA=TMA+CGB(I)*AU
      TMB=TMB+BU*SGB(KB)
      TMC=TMC+BU*AU
30 KB=KB-1
  CSQ(KC)=TMA
  SSQ(KC)=TMB
  SCS(KC)=TMC
  TMA=0.0
  TMB=0.0
  TMC=0.0
40 CONTINUE
  RDS=ERD*ERD
  LNC=ICT(NTE+2)-1
  ITR=2
  L=9*KTR-18
  CALLCLEAR(AMT,L)
  KDG=KDR-2
C  DRAG COMPUTATIONS DELETED. INITIALIZATION OF HNV AND VRN DELETED.
C
C  TEST FOR PROCESS COMPLETE - BEGIN THE ITERATIVE LOOP
C
  80 ITR=ITR+1
    IF(KTR-ITR)90,100,100
90 RETURN
100 KA=ITR-3
  I=I+KA*LNC
  LA=I+ICT(2)-1

```

```

LB=I+ICT(1)
LC=I+2
LD=LB
LE=LA
M=0
KB=1
TMQ=0.0
TMR=1.0/24.0
TMS=1.0/6.0
ICT2=ICT(2)
DO270I=1,NTE
N=M
KC=LCT(I)
IF(KC)210,210,120
120 CONTINUE
DO200J=1,KC
L=N-M+2
CZZ=L*(L-1)
CPZ=L*(L+1)
CPM=(L+1)*(L+2)
AU=UMT(LA)
BU=UMT(LB)
CU=UMT(LC)
DU=UMT(LD)
EU=UMT(LE)
AV=VMT(LA)
BV=VMT(LB)
CV=VMT(LC)
DV=VMT(LD)
EV=VMT(LE)
CC=CNM(KB)
SC=SNM(KB)
TMI=TMQ+CZZ*(CC*CU+SC*CV)
IF(M-1)140,160,180
140 AU=AU*TMR
BU=-BU*TMS
AV=-AV*TMR
BV=BV*TMS
TMQ=TMQ+1.0
TMR=TMR*TMQ/(TMQ+4.0)
TMS=TMS*(TMQ+1.0)/(TMQ+3.0)
GOTD180
160 AU=-AU*TMS
AV=AV*TMS
TMQ=TMQ+1.0
TMS=TMS*TMQ/(TMQ+2.0)
180 T=L-1
TMA=TMA+CC*(EU+CZZ*(CPM*AU-CU-CU))+SC*(EV+CZZ*(CPM*AV-CV-CV))
TMB=TMB+CC*(EV-CZZ*CPM*AV)+SC*(CZZ*CPM*AU-EU)
TMC=TMC-T*(CC*(CPZ*BU-DU)+SC*(CPZ*BV-DV))
TMF=TMF+T*(CC*(CPZ*BV+DV)-SC*(CPZ*BU+DU))
KB=KB+1
LA=LA+1
LB=LB+1
LC=LC+1
LD=LD+1
LE=LE+1
200 N=N+1

```

```

210 L=KA*LNC
    IF(M-1)220,230,240
220 TMQ=2.0
    TMS=1.0/12.0
    LA=L+ICT2+2
    LB=L+4
    GOT0260
230 LA=L+5
    GOT0250
240 LA=L+4+ICT(I-1)
250 LB=L+3+ICT(I)
260 LC=L+2+ICT(I+1)
    LD=L+1+ICT(I+2)
    LE=L+ICT(I+3)
270 M=M+1
    TMR=2.0*RDS
    TMC=TMC/TMR
    TMF=TMF/TMR
    TMR=2.0*TMR
    TMA=TMA/TMR
    TMB=TMB/TMR
    TMD=TMB
    TMG=TMC
    TMH=TMF
    TMI=TMI/RDS
    TME=-TMA-TMI
    L=9*KA+1
    DO280I=1,9
    FEE(L)=TMT(I)
280 L=L+1
    CALLCLEAR(TMT,9)
C
C   EVALUATE THE K-TH TERM OF THE MATRIX A(3,3)
C
C   DRAG COMPUTATIONS DELETED.
500 L=0
    KB=KA+1
    N=KB
    DO510I=1,KB
    AU=FEE(L+1)
    BU=FEE(L+2)
    CU=FEE(L+3)
    DU=FEE(L+5)
    EU=FEE(L+6)
    AV=CSQ(N)
    BV=SSQ(N)
    CV=SCS(N)
    DV=CGB(N)
    EV=SGB(N)
    TMA=TMA+AU*AV-2.0*BU*CV+DU*BV
    TMB=TMB+CV*(AU-DU)+BU*(AV-BV)
    TMC=TMC+CU*DV-EU*EV
    TMF=TMF+CU*EV+EU*DV
    N=N-1
510 L=L+9
    L=9*KA+9
    TMI=FEE(L)
    TMD=TMB

```

```

TME=-TMA-TMI
TMG=TMC
TMH=TMF
L=L-8
D0520I=1,9
AMT(L)=AMT(L)+TMT(I)
520 L=L+1
CALLCLEAR(TMT,9)
C EVALUATE U AND V MATRICES FOR (K+2)TH TERM IN SERIES
C
KB=KA+1
TMQ=KB*(KB+1)
TMQ=1.0/TMQ
TMR=1.0
KC=18*KA-2
D0600LA=1,2
LB=KC+39
LC=0
D0580LD=1,KB
LE=1
D0570J=1,3
I=3*J+KC
AU=UVM(I)
BU=UVM(I+1)
CU=UVM(I+2)
C DRAG COMPUTATIONS DELETED.
D0570M=1,3
I=LC+M
TMS=AMT(I)*AU+AMT(I+3)*BU+AMT(I+6)*CU
C DRAG COMPUTATIONS DELETED.
560 TMT(LE)=TMT(LE)+TMS
570 LE=LE+1
KC=KC-18
TMR=TMR+1.0
580 LC=LC+9
D0590I=1,9
UVM(LB)=TMQ*TMT(I)
590 LB=LB+1
CALLCLEAR(TMT,9)
TMR=1.0
600 KC=18*KA+7
GOTO80
END

```

```

SUBROUTINE ORBIN
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION EMODEL(10,2)
DIMENSION XIN(6),ORBEL(6),RO(3),VO(3),ORBNAME(6),ORBUNK(16),S(3,3)
EQUIVALENCE (ORBEL(1),ORBA),(ORBEL(2),ORBECC),(ORBEL(3),ORBINC),
1(ORBEL(4),RANODE),(ORBEL(5),ARGPGE),(ORBEL(6),ORBIM)
EQUIVALENCE (XIN(1),RO(1)),(XIN(4),VO(1)),(THEDOT,OMGI)
INTEGER*2 LFLG(40),KSTATO(15),MODEL(10,2)
INTEGER*2 GUIDE(10,2),IDAY,IYR
INTEGER*4 ENDSIG/1HE/,CONTIN,BLANK/1H/,ALFA/1HA/,ALFR/1HR/,ZCODE
DATA MAXSTO/15/,MAXSTA/40/,MAXEMU/10/
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
COMMON/OBSD/OBSD(150),OVOBSD
COMMON/ERDCON/ERDI,XMUI,OMGI,XCM,YCM,ZCM
COMMON/NSTA/NSTA
COMMON/STAORD/KORDER(150)
INTEGER STANAM,IDS*2
COMMON/STALOC/STAUVW(3,150),DATDUM(6,15),STANAM(5,150),IDS(150)
DIMENSION CM(3)
EQUIVALENCE(CM(1),XCM)
C READ EARTH CONSTANTS
READ(5,1065) ERDI,XMUI,OMGI,OVOBSD
1065 FORMAT(4D20.8)
C READ THE COORDINATES OF THE CENTER OF MASS IN THE COORDINATE SYSTEM IN WHICH
C THE STATION COORDINATES ARE GIVEN.
READ(5,1052) RO
C READ UNCERTAINTIES OF CENTER OF MASS
READ(5,1052) VO
WRITE(6,1063) RO,VO
1063 FORMAT(//,*,* COORDINATES OF THE CENTER OF MASS*,3F20.3/
1      *,* UNCERTAINTIES      *,3F20.3)
WRITE(6,1064) ERDI,XMUI,OMGI
1064 FORMAT(1H1//,*,* EARTH CONSTANTS FOR ORBIT INTEGRATION*/
1      T7,*SEMI-MAJOR AXIS*,T30,*GM*,T48,*ROTATION RATE*/
2      F20.3,1PD20.10,0PD20.10)
WRITE(3) ERDI,XMUI,OMGI,RO,VO
CHANGE THE DATUM ID NUMBER IF RESULTS WILL BE IN AN EARTH CENTERED DATUM
IF(RO(1).EQ.0) GO TO 135
IERD=ERDI
I=0
132 I=I+1
IDAT=DATPRM(1,I)
IF(IDAT.EQ.IERD) GO TO 133
IF(I.LT.15) GO TO 132
WRITE(6,1001)
1001 FORMAT(//10X,*EARTH DIAMETER DOES NOT CORRESPOND TO ANY KNOWN DATU
1M*/)
GO TO 135
133 DO 134 J=1,150
134 IDS(J)=I
135 CONTINUE
DO 130 I=1,3
130 CM(I)=RO(I)
IF(PCODE(14).EQ.0) GO TO 140
DO 139 I=1,3
139 STAUVW(I,NSTA+1)=RO(I)
KORDER(NSTA+1)=0

```

```

140 CONTINUE
  SPR=206264.80625
  NOBS=0
  NUNK=0
  DO 160 I=7,16
160  ORBUNK(I)=0
C      READ FIRST CARD
  165 READ(5,1051) NORB,ORBNAM,IOCODE,CONTIN
  1051 FORMAT(A4,6A8,I1,26X,A1)
C      TEST FOR END OF ALL PASSES
  IF(CONTIN.EQ.ENDSIG) GO TO210
  WRITE(6,1062) NORB,ORBNAM
  1062 FORMAT(7(/),1X,A4,5X,6A8)
C
C      IOCODE=0 MEANS RECTANGULAR ELEMENTS ARE GIVEN IN TRUE SIDEREAL SYSTEM
C      IOCODE=1 MEANS RECTANGULAR ELEMENTS ARE GIVEN IN MODIFIED SIDEREAL
C      SYSTEM.
C      IOCODE=2 MEANS RECTANGULAR ELEMENTS ARE GIVEN IN EARTH FIXED SYSTEM
C      IOCODE=3 MEANS KEPLERIAN ELEMENTS ARE GIVEN, REFERED TO TRUE EQUINOX.
C      IOCODE=4 MEANS KEPLERIAN ELEMENTS ARE GIVEN, REFERED TO TRUE EQUATOR
C      AND 1950 EQUINOX (I.E., THE SOA ORBITAL SYSTEM).
C
C      DO 166 I=1,40
166  LFLG(I)=0
  DO 167 I=1,10
  DO 167 J=1,2
  GUIDE(I,J)=00.0
167  CONTINUE
  NUTORB=6
  NEMUNK=0
  IF(IOCODE.LT.3)GO TO 170
  READ(5,1052) ORBA,ORBECC,ORBINC
  READ(5,1052) RANODE,ARGPGE,ORBm
  1052 FORMAT(3D15.8)
  WRITE(6,1055) ORBEL
  1055 FORMAT(1X,'A=',D18.8,'ECC=',D18.8,'INC=',D18.8,'RA OF NODE=',D15.8,
  'ARG OF PERIGEE=',D15.8,'MEAN ANOMALY=',D15.8)
  CALL KEPTCE(ORBEL,XIN)
  IOCODE=IOCODE-3
  GO TO 171
170  READ(5,1052) RD,VO
C      THERE ARE THREE WAYS THE EPOCH CAN APPEAR, AS JULIAN DAYS IN COLUMNS
C      0 THRU 15, AS A DATE IN COLUMNS 16 THRU 35, OR IF A DATE OTHER THAN THAT
C      IS TO BE USED THE DATE OF THE ORBITAL ELEMENTS GIVEN WILL BE IN
C      COLUMNS 16 THRU 35 WITH THE DESIRED EPOCH TIME IN COLUMNS 51 THRU 71
C      ZCODE COLUMNS 36 THRU 41 MUST HAVE SOMETHING PUNCHED IN IT IN THE LAST CASE
  171  READ(5,1053) EPOCH,1DAY,MONTH,IYR,IH,MIN,ESEC,ZCDE,IHR,IMIN,SEC
  1053 FORMAT(D15.8,I5,2X,A3,3I5,D10.5,3X,A2,T51,2I5,D10.5)
  IF(EPOCH.EQ.0) GO TO 175
  ITEM=EPOCH
  TEM=ITEM
  TE=(EPOCH-TEM)*86400
  TEMP=TE/3600.0
  IH=TEMP
  TEMP=(TEMP-IH)*60.0
  MIN=TEMP
  ESEC=(TEMP-MIN)*60.0
  TO=IHR*3600+IMIN*60+SEC

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TIME=EPOCH
GO TO 180
175 TEM=MJD(IDAY,MONTH,IYR)
TE=IH*3600+MIN*60+ESEC
TO=IHR*3600+IMIN*60+SEC
EPOCH=TEM+TE/86400.0
TIME=TEM+TO/86400.0
IF(ZCODE.EQ.BLANK) TIME=EPOCH
180 CONTINUE
WRITE(6,1054) RO,VO,EPOCH
IF(IOCODE.NE.1) GO TO 185
CALL PRENUT(TIME,PAN)
CALL ROT3(-PAN,RO)
CALL ROT3(-PAN,VO)
185 GASTO=GSTD(TIME)
CALL PRENUT(TIME,PAN)
GASTO=GASTO+PAN
1054 FORMAT(/1X,'X=',D18.8,1X,'Y=',D18.8,1X,'Z=',D18.8,/1X,'XDOT=',D15.8,1X,'YDOT=',D15.8,1X,'XDOT=',D15.8,/1X,'EPOCH=',D15.8,1X,'215,2X,A3,I5,/22X,'GAST=',D15.8//)
IF(IOCODE.NE.2) GO TO 190
R01=RO(1)
R02=RO(2)
CALL ROT3(-GASTO,RO)
VO(1)=VO(1)-THEDOT*R01
VO(2)=VO(2)+THEDOT*R02
CALL ROT3(-GASTO,VO)
190 IF(TO-TE) 195,200,196
195 IF(ZCODE.EQ.BLANK) GO TO 200
196 IF(TE.NE.0) GO TO 198
EPOCH=TIME
GO TO 200
198 CALL ORBIT(0,TO,TE,GASTO,XIN)
200 CALL POLE(EPOCH,XPM,YPM)
XPM=XPM/SPR
YPM=YPM/SPR
XP=XPM
YP=YPM
GAST=GSTD(EPOCH)
CALL PRENUT(EPOCH,PAN)
GAST=GAST+PAN
WRITE(6,1154)
1154 FORMAT(/1X,'VALUES STORED ON UNIT 3')
WRITE(6,1054) RO,VO,EPOCH,IDAD,MONTH,IYR,GAST
210 WRITE(3) IORB,NORB,DRBNAM,EPOCH,IDAD,MONTH,IYR,GAST,RO,VO,TE,IH,
IMIN,ESEC,XPM,YPM,CONTIN
IF(CONTIN.EQ.ENDSIG) RETURN
NOBSTD=0
WRITE(6,6005)
6005 FORMAT(' STATION DATE',T25,'TIME(UT)',3X,'OBSERVED RANGE',2X,
1 'UNCERTAINTY')
C      READ OBSERVATION CARD
220 READ(5,1061) ID,IYR,MONTH,IDAD,IH,IMIN,SEC,SEC1,RS0,RS01,VARRA,
1CONTIN
1061 FORMAT(14X,I4,5I2,F2.0,F4.0,F16.0,F3.0,11X,F6.3,9X,A1)
C      TEST FOR END OF PASS
IF(CONTIN.EQ.ENDSIG) GO TO 240
KSTA=KSTAID(ID)

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IF(CONTIN.EQ.ALFA) GO TO 290
IF(CONTIN.EQ.ALFR) GO TO 300
NOBSTO=NOBSTO+1
NOBS=NOBS+1
IF(PCODE(12).EQ.1) VARRA=OBSD(KSTA)
IF(PCODE(12).EQ.2) VARRA=OVOBSD
RSO=RSO+RSO1/1000.0
IF(SEC1.LT.1) GO TO 230
SEC=SEC+SEC1/10000.0
GO TO 231
230 SEC=SEC+SEC1
231 CONTINUE
LFLG(KSTA)=1
TEMO=MJD(IDAY,MONTH,IYR)
TOBS=IH*3600+IMIN*60+SEC+(TEMO-TEM)*86400.0
TO=TOBS-TE
IF(DABS(TO).GT.10000) GO TO 260
GASTO=GAST+TO*THEDOT
SINST=DSIN(GASTO)
COSST=DCOS(GASTO)
C FILL IN "S" ARRAY ACCORDING TO SAO SPECIAL REPORT 123, "PRECISE ASPECTS
C OF TERRESTRIAL AND CELESTIAL REFERENCE FRAMES", PAGE 8
S(1,1)=COSST
S(2,1)=-SINST
S(3,1)=-XP*COSST-YP*SINST
S(1,2)=SINST
S(2,2)=COSST
S(3,2)=-XP*SINST+YP*COSST
S(1,3)=XP
S(2,3)=-YP
S(3,3)=1.0
WRITE(6,1056) ID, IDAY, MONTH, IYR, IH, IMIN, SEC, RSO, VARRA
1056 FORMAT(17,14,1X,A3,I3,15,I3,F8.4,F13.2,2F15.2)
240 WRITE(3) TOBS, TO, KSTA, GASTO, SINST, COSST, RSO, VARRA, ID, IDAY, MONTH,
1 IYR, IH, IMIN, SEC, S, CONTIN
IF(CONTIN.EQ.ENDSIG) GO TO 270
GO TO 220
260 WRITE(6,1057) TOBS, TE
1057 FORMAT(1X,'TIME OF OBSERVATION=',F20.10,/1X,'TIME OF EPOCH=',
1F20.10)
STOP
270 NSTATO=0
DO 280 KSTA=1,NSTA
IF(LFLG(KSTA).EQ.0) GO TO 280
NSTATO=NSTATO+1
IF(NSTATO.LE.MAXSTO) GO TO 277
WRITE(6,1058) MAXSTO
1058 FORMAT(1X,'MORE THAN',I2,' STATIONS OBSERVED')
STOP
277 KSTATO(NSTATO)=KSTA
LFLG(KSTA)=NSTATO
280 CONTINUE
IF(PCODE(14).EQ.0) GO TO 2800
NSTATO=NSTATO+1
KSTATO(NSTATO)=NSTATO+1
2800 CONTINUE
DO 281 I=1,6
281 ORBUNK(I)=XIN(I)

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IF(NEMUNK.EQ.0) GO TO 285
DO 283 I=1,NEMUNK
C           GET THE ORDER # OF THE STATION TO WHICH THIS ERROR MODEL
C TERM PERTAINS
KSTA=MODEL(I,2)
ISTA=LFLG(KSTA)
IF(ISTA.NE.0) GO TO 282
WRITE(6,1059) KORDER(KSTA)
1059 FORMAT(1X,'ERROR MODEL UNKNOWN FOR STATION ',I3,' IS MEANINGLESS
1SINCE STATION DOES NOT OBSERVE THIS PASS.')
282 MODCOD=MODEL(I,1)
GUIDE(ISTA,MODCOD)=I+6
283 CONTINUE
285 WRITE(4) ORBUNK,NSTATO,LFLG,KSTATO,NEMUNK,EMODEL,MODEL,GUIDE
1 , NUTORB
GO TO 165
290 MODCOD=1
GO TO 310
300 MODCOD=2
GO TO 310
310 NEMUNK=NEMUNK+1
IF(NEMUNK.LE.MAXEMU) GO TO 311
WRITE(6,1060) MAXEMU
1060 FORMAT(1X,'ERROR-NUMBER OF ERROR MODEL UNKNOWN'S EXCEEDS ',I5)
STOP
311 NUTORB=NUTORB+1
MODEL(NEMUNK,1)=MODCOD
MODEL(NEMUNK,2)=KSTA
EMODEL(NEMUNK,1)=RS0
EMODEL(NEMUNK,2)=VARRA
ORBUNK(NUTORB)=RS0
GO TO 220
END
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SUBROUTINE FOAN
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 IDAY,1YR
INTEGER*4 CONTIN,ENDSIG/1HE/
COMMON/NSTA/NSTA
INTEGER STANAM,IDS*2
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1 STANAM(5,150),IDS(150)
COMMON/STAORD/KORDER(150)
COMMON/WPW/WPW,XPU,IDEKF,NFSTA
DIMENSION NOBSTA(40)
REAL*4 VPVSTA(40)
DIMENSION RO(3),VO(3),XT(6),UVWT(3,9),EXT(3),EUVWT(3,9),XS(3),
1 DCS(3),S(3,3)
COMMON/ERDCON/ERDI,XMUI,OMGI,CM(3)
DIMENSION C(3),CC(3),D(9),DNO(3,3,40)
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE

C THIS SUBROUTINE IS DIMENSIONED FOR A MAXIMUM OF
C 40 OBSERVING STATIONS(MAXIMUM)
C 15 STATIONS OBSERVING ANY GIVEN PASS
C 10 ERROR MODEL UNKNOWNs FOR ANY PASS (FOR A TOTAL OF 16 UNKNOWNs
C FOR ANY PASS).
DIMENSION ORBNAM(6),ORBUNK(16),DB(16)
DIMENSION A(3),B(16)
DIMENSION DN(3,3,40),BN(3,16,15),DK(3,40),DDN(16,16),DDK(16)
INTEGER*2 LFLG(40),KSTATO(15),MODEL(10,2),IGUIDE(10,2)
DIMENSION EMODEL(10,2)
EQUIVALENCE(D(1),C(1)),(D(4),B(1))
INTEGER*4 MODALF(3,2)
DATA MODALF/"ZERO","SET","  ","REFR","ACTE","ON  "/
MAXSTA=40
MAXSTO=15
MAXEMU=10
MAXUNK=16
REWIND 1
REWIND 3
REWIND 4
REWIND 2
IF(PCODE(14).NE.0) NSTA=NSTA+1
C INITIAL ACCUMULATING ARRAYS
DO 60 KSTA=1,NSTA
DO 50 J=1,3
DO 40 I=1,3
DNO(I,J,KSTA)=0.0
40 DN(I,J,KSTA)=0.0
50 DK(J,KSTA)=0.0
VPVSTA(KSTA)=0.0
60 NOBSTA(KSTA)=0
WPW=0.0
WPWSP=0.0
NOBS=0
NUNK=0
READ(3) ERDI,XMUI,OMGI,RO,VO
C RO HOLDS A PRIORI COORDINATES OF CENTER OF MASS
C VO HOLDS A PRIORI UNCERTAINTIES OF COORDINATES OF CENTER OF MASS
DO 65 I=1,3
65 CM(I)=RO(I)

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      IF(PCODE(14).EQ.0) GO TO 75
      DO 70 I=1,3
C   UPDATE COORDINATES OF CENTER OF MASS
      CM(I)=STAUVW(I,NSTA)
C   IMPOSE A PRIORI CONSTRAINTS ON COORDINATES OF CENTER OF MASS
      IF(VO(I).EQ.0.0) GO TO 70
      DN(I,I,NSTA)=1.0/VO(I)**2
      DK(I,NSTA)=(RO(I)-STAUVW(I,NSTA))/VO(I)**2
      70 CONTINUE
      75 CONTINUE
C
      150 CONTINUE
C   PROCESS ANOTHER PASS
C   READ ORBIT HEADER
      READ(3) IORB,NORB,ORBNAM,EPOCH,IDAY,MONTH,IYR,GAST,RO,VO,TE,IH,MIN
      1,ESEC,XPM,YPM,CONTIN
      WRITE(1) IORB,NORB,ORBNAM,1DAY,MONTH,IYR,IH,MIN,ESEC,EPOCH,CONTIN
C   TEST FOR END OF DATA
      IF(CONTIN.EQ.ENDSIG)GO TO 700
      READ(4) ORBUNK,NSTATO,LFLG,KSTATO,NEMUNK,EMODEL,MODEL,IGUIDE
      1 , NUTORB
C   INITIALIZE ACCUMULATION ARRAYS
      WPWT0=0.0
      NOBST0=0
      DO 170 KSTA=1,NSTATO
      DO 170 J=1,NUTORB
      DO 170 I=1,3
      170 BN(I,J,KSTA)=0.0
      DO 180 I=1,MAXUNK
      DDK(I)=0.0
      DO 180 J=1,MAXUNK
      180 DDN(I,J)=0.0
      WRITE(6,6001) NORB,ORBNAM,1DAY,MONTH,IYR,IH,MIN,ESEC,EPOCH
6001 FORMAT(/////////2X,A4,3X,6A8,' EPOCH=',I3,A3,1X,2I3,'H',I3,'M',
      1 F8.4,'S UT=MJD',F17.9)
      WRITE(6,6002) (ORBUNK(I),I=1,6)
6002 FORMAT('0 CURRENT ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN
      1 COORDINATES'/5X,'POSITION(METERS)',4X,3F16.3/
      1 5X,'VELOCITY(METERS/SEC)',3F16.6)
      IF(NEMUNK.EQ.0) GO TO 185
      WRITE(6,6003)
6003 FORMAT('0 CURRENT VALUE OF ERROR MODEL UNKNOWN$',)
      DO 184 I=1,NEMUNK
      MODCOD=MODEL(I,1)
      KSTA=MODEL(I,2)
      184 WRITE(6,6004) (MODALF(J,MODCOD),J=1,3),KORDER(KSTA),(STANAM(J,KSTA
      1),J=1,5),ORBUNK(I+6)
6004 FORMAT(5X,3A4,'FOR STATION',I6,2X,5A4,2X,'=',F15.3)
      185 CONTINUE
      WRITE(6,6005)
6005 FORMAT('0STATION DATE',T25,'TIME(UT)',3X,'CORRECTED RANGE',2X,
      1 'UNCERTAINTY',4X,'MISCLUSION')
C   INITIALIZE ORBIT
      NFLG=1
      CALL ORBIT(NFLG,TE,TE,GAST,ORBUNK)
C
C   READ OBSERVATION RECORD
200 READ (3) TOBS,TO,KSTA,GASTO,SINST,COSST,RSO,VARRA,1D,1DAY,MONTH,

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1 IYR,IH,MIN,SEC,S,CONTIN
IF(CONTIN.EQ.ENDSIG) GO TO 350
NN=1
CALL DRIVER(NN,TOBS,GH,DT,XT,UVWT)
C GET INTERNAL NUMBER OF THIS STATION IN THIS PASS
ISTA=LFLG(KSTA)
C TRANSFORM SATELLITE POSITION AND TRANSITION MATRIX TO TERRESTRIAL COORDINATES
CALL DGMPRD(S,XT,EXT,3,3,1)
CALL DGMPRD(S,UVWT,EUVWT,3,3,9)
COMPUTE STATION TO SATELLITE VECTOR
DO 220 I=1,3
220 XS(I)=EXT(I)-STAUVW(I,KSTA)
C GET COMPUTED RANGE
RSC=0.0
DO 225 I=1,3
225 RSC=RSC+XS(I)**2
RSC=DSQRT(RSC)
C GET STATION TO SATELLITE DIRECTION COSINES AND COMPUTE PARTIALS
DO 230 I=1,3
DCS(I)=XS(I)/RSC
A(I)=-DCS(I)
230 CONTINUE
CALL DGMPRD(DCS,EUVWT,D,1,3,9)
C PREPARE PARTIALS WITH RESPECT TO THE CENTER OF MASS.
IF(PCODE(14).NE.0) CALL DGMPRD(S,C,CC,3,3,1)
C FILL IN B WITH ZEROS
IF(NUTORB.LE.6) GO TO 250
DO 234 I=7,NUTORB
234 B(I)=0.0
235 CONTINUE
CORRECT OBSERVATION FOR ERROR MODEL
IUNK=IGUIDE(ISTA,1)
IF(IUNK.EQ.0) GO TO 245
RSO=RSO-ORBUNK(IUNK)
B(IUNK)=1.0
245 IUNK=IGUIDE(ISTA,2)
IF(IUNK.EQ.0) GO TO 250
COEFFICIENT OF REFRACTION TERM IS 1.0/(SIN OF ELEVATION ANGLE)
C GET SIN OF ELEVATION ANGLE (SE)
C GEODETIC LATITUDE IS APPROXIMATED BY SPHERICAL LATITUDE IN COMPUTATION
C OF SIN OF ELEVATION ANGLE
SE=0.0
RSTA=0.0
DO 246 I=1,3
SE=SE+STAUVW(I,KSTA)*DCS(I)
246 RSTA=RSTA+STAUVW(I,KSTA)**2
SE=SE/DSQRT(RSTA)
RSO=RSO-ORBUNK(IUNK)/SE
B(IUNK)=1.0/SE
250 CONTINUE
COMPUTE MISCLOSURE
DL=RSO-RSC
WRITE(6,6006) KORDER(KSTA),IDAY,MONTH,IYR,IH,MIN,SEC,RSO,VARRA,DL
6006 FORMAT(I7,I4,1X,A3,I3,I5,I3,F8.4,F13.2,2F15.2)
WT=1.0/VARRA**2
WPW OBS=DL*WT*DL
VPVSTA(KSTA)=VPVSTA(KSTA)+WPW OBS
WPW TO=WPW TO+WPW OBS

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WPH=WPH+WPHOBS
NOBSTA(KSTA)=NOBSTA(KSTA)+1
NOBSTO=NOBSTO+1
NOBS=NOBS+1

C ACCUMULATE NORMAL EQUATIONS
  DO 270 I=1,3
    DK(I,KSTA)=DK(I,KSTA)+A(I)*WT*DL
    DO 260 J=1,3
      DO 260 DN(I,J,KSTA)=DN(I,J,KSTA)+A(I)*WT*A(J)
      DO 265 J=1,NUTORB
        265 BN(I,J,ISTA)=BN(I,J,ISTA)+A(I)*WT*B(J)
      270 CONTINUE
      DO 280 I=1,NUTORB
        DDK(I)=DDK(I)+B(I)*WT*DL
        DO 280 J=1,NUTORB
          280 DDN(I,J)=DDN(I,J)+B(I)*WT*B(J)
          IF(PCODE(14).EQ.0) GO TO 200
C ACCUMULATE NORMALS PERTAINING TO THE CENTER OF MASS
  DO 295 I=1,3
    DK(I,NSTA)=DK(I,NSTA)+CC(I)*WT*DL
    DO 291 J=1,3
      DN(I,J,NSTA)=DN(I,J,NSTA)+CC(I)*WT*CC(J)
    291 DNO(I,J,KSTA)=DNO(I,J,KSTA)+A(I)*WT*CC(J)
    DO 292 J=1,NUTORB
      292 BN(I,J,NSTATO)=BN(I,J,NSTATO)+CC(I)*WT*B(J)
    295 CONTINUE
C RETURN TO PROCESS ANOTHER OBSERVATION
  GO TO 200
C
  350 CONTINUE
C ENTER ON END OF PASS
  NUNK=NUNK+NUTORB
C ADD A PRIORI CONSTRAINTS ON ORBIT UNKNOWNNS TO DDN AND DDK AT THIS POINT
C ADD A PRIORI CONSTRAINTS ON ERROR MODEL UNKNOWNNS
  IF(NEMUNK.EQ.0) GO TO 370
  DO 369 I=1,NEMUNK
    IF(EMODEL(I,2).LE.0.0) GO TO 369
    DL=EMODEL(I,1)-ORBUNK(I+6)
    WT=1.0/EMODEL(I,2)**2
    DDN(I+6,I+6)=DDN(I+6,I+6)+WT
    DDK(I)=DDK(I)+WT*DL
    WPHOBS=DL*WT*DL
    NOBS=NOBS+1
    WPH=WPH+WPHOBS
  369 CONTINUE
  370 CONTINUE
  IF(NUTORB.EQ.MAXUNK)GO TO 380
C PAD OUT DDN
  NN=NUTORB+1
  DO 379 I=NN,MAXUNK
    DDK(I)=0.0
  379 DDN(I,I)=1.0
C INVERT DDN
  DET=1.0
  380 CALL DMINV(DDN,MAXUNK,DET,UVWT,B)
COMPUTE PARTIAL UNCERTAINTIES OF ORBIT UNKNOWNNS
  DO 390 I=1,NUTORB

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390 B(I)=DSQRT(DDN(I,I))
      WRITE(6,6008) (B(I),I=1,NUTORB)
6008 FORMAT(//' PARTIAL UNCERTAINTIES OF ORBIT UNKNOWNNS-FROM DDN(INV) '
      1 ,10(/8D15.6))
C
      DO 395 I=1,NUTORB
      DO 395 J=1,NUTORB
395 WPWSP=WPWSP+DDK(I)*DDN(I,J)*DDK(J)
      WRITE(1) DDN,DDK,BN
      WRITE(1) NUTORB,NSTATO,NEMUNK,KSTATO,LFLG,ORBUNK,EMODEL,MODEL,
      1 IGUIDE
      RMSTO=DSQRT(WPWTO/DFLOAT(NOBSTO))
      WRITE(6,6007) WPWTO,NOBSTO,RMSTO
6007 FORMAT(//5X,'WEIGHTED SUM OF SQUARES OF MISCLOSURES =',F15.3/
      1 5X,'NUMBER OF OBSERVATIONS =',18/5X,'RMS MISCLOSURE =',F15.3)
C RETURN TO PROCESS ANOTHER PASS
      GO TO 150
C
C ENTER HERE AT THE END OF ALL PASSES
700 CONTINUE
      WRITE(2) DN,DK
      IF(PCODE(14).NE.0) WRITE(2) DNO
      REWIND 1
      REWIND 2
      REWIND 4
C PERFORM ANALYSIS OF MISCLOSURES BY STATION
      WRITE(6,6019)
6019 FORMAT(1H1,8(/),10X,'ANALYSIS OF MISCLOSURES BY STATION'//
      1T10,'STATION',T40,'NUMBER OF OBSERVATIONS',T70,'RMS MISCLOSURE')
      DO 750 KSTA=1,INSTA
      IF(NOBSTA(KSTA).GT.0) RMSMC=DSQRT(VPVSTA(KSTA)/DFLOAT(NOBSTA(KSTA)
      1))
      WRITE(6,6020) KORDER(KSTA),(STANAM(I,KSTA),I=1,5),NOBSTA(KSTA),
      1 RMSMC
6020 FORMAT(1T10,17,1X,5A4,T55,17,T70,F14.2)
750 CONTINUE
COMPUTE DEGREES OF FREEDOM
      IDEGF=NOBS-NUNK
      WPW=WPW-WPWSP
      RMSMC=DSQRT(WPW/DFLOAT(IDEgf))
      WRITE(6,6021) NOBS,NUNK,IDEgf,WPW,RMSMC
6021 FORMAT(///10X,'TOTAL NUMBER OF OBSERVATIONS',T60,I8//,
      110X,'TOTAL NUMBER OF ORBIT AND ERROR MODEL UNKNOWNNS',T60,I8//,
      210X,'CORRESPONDING DEGREES OF FREEDOM',T60,I8//,
      310X,'TOTAL SUM OF SQUARES OF MISCLOSURES, I.E.,VPV+XU',T60,F11.2//,
      410X,'CORRESPONDING STANDARD DEVIATION OF UNIT WEIGHT',T60,F11.2)
      RETURN
      END

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SUBROUTINE POLE(DT, XPM, YPM)
DOUBLE PRECISION DT,XPM, YPM
DIMENSION PM(286,2),PMT(2)
DIMENSION PMX1(90),PMY1(90),PMX2(90),PMY2(90),PMX3(15),PMY3(15)
EQUIVALENCE (PM,PMX1),(PM(1,2),PMY1),(PM(91,1),PMX2),
1(PM(91,2),PMY2),(PM(181,1),PMX3),(PM(181,2),PMY3)
DIMENSION PMX4(5),PMY4(5)
EQUIVALENCE (PM(196,1),PMX4),(PM(196,2),PMY4)
DIMENSION PMX5(86),PMY5(86)
EQUIVALENCE (PM(201,1),PMX5),(PM(201,2),PMY5)
C POLAR MOTION TABLES FURNISHED BY TOMLINSON (TAKEN FROM IPMS)
DATA PMX1/
1-0.173, -0.215, -0.235, -0.237, -0.218, -0.162, -0.097, -0.032, 0.036,
2 0.111, 0.188, 0.237, 0.348, 0.398, 0.398, 0.368, 0.330, 0.280,
3 0.218, 0.144, 0.069, 0.000, -0.062, -0.112, -0.140, -0.153, -0.151,
4-0.126, -0.086, -0.037, 0.026, 0.092, 0.161, 0.223, 0.272, 0.299,
5 0.308, 0.296, 0.261, 0.202, 0.135, 0.073, 0.046, 0.035, 0.013,
6-0.026, -0.072, -0.096, -0.107, -0.103, -0.087, -0.039, 0.004, 0.040,
7 0.070, 0.080, 0.109, 0.117, 0.117, 0.109, 0.092, 0.074, 0.065,
8 0.064, 0.062, 0.057, 0.046, 0.034, 0.030, 0.032, 0.040, 0.043,
9 0.042, 0.041, 0.039, 0.028, 0.019, -0.010, -0.027, -0.021, -0.009,
A 0.008, 0.027, 0.047, 0.071, 0.095, 0.120, 0.144, 0.162, 0.173 /
DATA PMX2/
1 0.171, 0.157, 0.128, 0.094, 0.056, 0.017, -0.019, -0.054, -0.086,
2-0.110, -0.121, -0.119, -0.105, -0.076, -0.038, 0.009, 0.070, 0.134,
3 0.191, 0.239, 0.274, 0.301, 0.281, 0.237, 0.176, 0.112, 0.048,
4-0.011, -0.069, -0.122, -0.171, -0.206, -0.194, -0.169, -0.139, -0.101,
5-0.055, 0.004, 0.074, 0.164, 0.214, 0.240, 0.241, 0.239, 0.255,
6 0.250, 0.219, 0.161, 0.099, 0.042, -0.012, -0.067, -0.120, -0.160,
7-0.185, -0.196, -0.194, -0.174, -0.130, -0.072, -0.003, 0.071, 0.127,
8 0.168, 0.201, 0.221, 0.227, 0.220, 0.194, 0.138, 0.075, 0.033,
9 0.000, -0.029, -0.058, -0.086, -0.105, -0.116, -0.119, -0.115, -0.104,
A-0.086, -0.057, -0.010, 0.052, 0.096, 0.117, 0.125, 0.123, 0.115/
DATA PMY1/
1 0.022, 0.098, 0.187, 0.265, 0.328, 0.389, 0.443, 0.478, 0.493,
2 0.478, 0.447, 0.411, 0.365, 0.307, 0.235, 0.165, 0.097, 0.043,
3-0.007, -0.038, -0.057, -0.064, -0.057, -0.025, 0.032, 0.120, 0.211,
4 0.285, 0.340, 0.372, 0.393, 0.406, 0.410, 0.401, 0.370, 0.320,
5 0.260, 0.201, 0.143, 0.090, 0.043, 0.007, -0.012, -0.007, 0.025,
6 0.059, 0.094, 0.123, 0.153, 0.182, 0.209, 0.238, 0.263, 0.288,
7 0.300, 0.306, 0.301, 0.288, 0.271, 0.249, 0.220, 0.189, 0.161,
8 0.150, 0.151, 0.158, 0.161, 0.160, 0.155, 0.153, 0.150, 0.151,
9 0.154, 0.157, 0.165, 0.174, 0.191, 0.212, 0.242, 0.276, 0.297,
A 0.309, 0.314, 0.312, 0.304, 0.290, 0.271, 0.246, 0.214, 0.175/
DATA PMY2/
1 0.132, 0.092, 0.068, 0.060, 0.067, 0.083, 0.104, 0.128, 0.160,
2 0.200, 0.248, 0.295, 0.329, 0.356, 0.376, 0.388, 0.387, 0.375,
3 0.349, 0.307, 0.251, 0.193, 0.139, 0.091, 0.046, 0.008, -0.020,
4 0.005, 0.041, 0.078, 0.120, 0.168, 0.230, 0.294, 0.353, 0.412,
5 0.455, 0.467, 0.459, 0.436, 0.394, 0.339, 0.275, 0.219, 0.168,
6 0.123, 0.085, 0.060, 0.046, 0.043, 0.049, 0.069, 0.103, 0.153,
7 0.226, 0.286, 0.334, 0.374, 0.408, 0.434, 0.444, 0.433, 0.399,
8 0.349, 0.303, 0.259, 0.221, 0.186, 0.156, 0.131, 0.114, 0.103,
9 0.098, 0.100, 0.108, 0.124, 0.149, 0.181, 0.215, 0.255, 0.298,
A 0.330, 0.344, 0.345, 0.337, 0.324, 0.308, 0.291, 0.273, 0.253/
DATA PMX3/
1 0.099, 0.079, 0.056, 0.031, 0.012, -0.001, -0.006, -0.008, -0.002,
2 0.012, 0.035, 0.055, 0.046, 0.027, 0.008 /

```

```

DATA PMX4           /-0.010,-0.029,-0.049,-0.063,-0.066/
DATA PMY3/
1 0.233, 0.213, 0.194, 0.177, 0.165, 0.157, 0.155, 0.154, 0.152,
2 0.156, 0.163, 0.172, 0.183, 0.195, 0.208 /
DATA PMY4           / 0.220, 0.234, 0.249, 0.269, 0.289/
DATA PMX5/
1-.056,-.037,-.014,.008,.031,.051,.064,.067,.064,.060,.088,.119,.10
26 ,.054,.008,-.027,-.056,      -.084,-.109,-.123,-.127,-.120,-.102
3,-.073,-.033,.010,.052,.091,.125,.154,.174,.185,.184,.168,.127,.07
47 ,.029,-.021,-.071,-.115,-.157,-.184,-.184,-.166,-.135,-.100,-.06
53 ,-.025,.017,.083,.154,.210,.241,.250,.238,.207,.167,.119,.068,
6.019,-.045,-.127,-.211,-.242,-.225,-.184,-.135,-.078,-.027,.026,
7.086,.150,.214,.261,.270,.256,.220,.177,.143,.114,.071,.015,-.015,
8-.042,-.081,-.125/
DATA PMY5/
1.302,.308,.308,.302,.290,.276,.260,.245,.231,.216,.202,.183,.166,
2.157,.156,.161,.172,      .197,.233,.265,.289,.310,.330,.350,.370,
3.386,.392,.386,.367,.337,.302,.260,.212,.167,.134,.115,.105,.104,
4.114,.134,.168,.216,.273,.333,.384,.419,.449,.465,.463,.436,.391,
5.347,.303,.252,.196,.148,.112,.080,.045,.017,.020,.052,.097,.149,
6.204,.270,.340,.389,.443,.478,.482,.468,.444,.409,.337,.276,.236,
7.160,.166,.122,.070,.033,.030,.053,.109,.174/
A=(DT-0.36203861D5)*0.54758185D-1
L=A+1.0
IF(L.LT.2)GOTO901
IF(L.GT.284) GO TO 901
TL=L
AN=A+1.0-TL
B=AN*(AN-1.0)/4.0
DO10I=1,2
DELO=PM(L,I)-PM(L-1,I)
DEL1=PM(L+1,I)-PM(L,I)
DEL2=PM(L+2,I)-PM(L+1,I)
PMT(I)=PM(L,I)+AN*DEL1+B*(DEL2-DELO)
10 CONTINUE
XPM=PMT(1)
YPM=PMT(2)
RETURN
901 WRITE(6,9001)
9001 FORMAT(72HOTABLES OF POLAR MOTION COVER ONLY FROM 1958.0 TO 1972.2
15, PLEASE EXTEND)
STOP
END

```

```

SUBROUTINE KEPTCE(ORBEL,Z)
CONVERT FROM KEPLERIAN TO CARTESIAN ORBIT ELEMENTS
C (EPOCH POSITION AND VELOCITY VECTORS)
C
IMPLICIT REAL*8(A-Z)
COMMON/ERDCON/AE,GM,OMGI,XCM,YCM,ZCM
DIMENSION ORBEL(6),Z(6),RXQ(3,3),Q(3,2)
DATA RHOD/ 57.295779513082/
A=ORBEL(1)
ECC=ORBEL(2)
INC=ORBEL(3)/RHOD
NODE=ORBEL(4)/RHOD
ARGP=ORBEL(5)/RHOD
M=ORBEL(6)/RHOD
C
CALL KEPEQ(M,ECC,1.0D-12,E)
CE=DCOS(E)
SE=DSIN(E)
Q(1,1)=A*(CE-ECC)
Q(2,1)=A*DSQRT(1.0-ECC**2)*SE
Q(3,1)=0.0
N=DSQRT(GM/A**3)
FACTOR=N*A/(1.0-ECC*CE)
Q(1,2)=-SE*FACTOR
Q(2,2)=DSQRT(1.0-ECC**2)*CE*FACTOR
Q(3,2)=0.0
CN=DCOS(NODE)
SN=DSIN(NODE)
CW=DCOS(ARGP)
SW=DSIN(ARGP)
CI=DCOS(INC)
SI=DSIN(INC)
RXQ(1,1)=CN*CW-SN*CI*SW
RXQ(1,2)=-CN*SW-SN*CI*CW
RXQ(1,3)=SN*SI
RXQ(2,1)= SN*CW+SW*CN*CI
RXQ(2,2)=-SN*SW+CW*CN*CI
RXQ(2,3)=-CN*SI
RXQ(3,2)=CW*SI
RXQ(3,1)=SW*SI
RXQ(3,3)=CI
CALL DGMPRD(RXQ,Q(1,1),Z(1),3,3,1)
CALL DGMPRD(RXQ,Q(1,2),Z(4),3,3,1)
RETURN
END

```

```
SUBROUTINE KEPEQ(M,EC,CONV,E)
IMPLICIT REAL*8(A-H,K-Z)
E=M+EC*DSIN(M)
DO 10 I=1,50
DELE=(M-E+EC*DSIN(E))/(1.0-EC*DCOS(E))
E=E+DELE
IF(DABS(DELE).LT.CONV) GO TO20
10 CONTINUE
WRITE(6,100)
100 FORMAT(52HKEPLERS EQUATION FAILS TO CONVERGE IN 50 ITERATIONS )
20 RETURN
END
```

```

SUBROUTINE ORBIT(NFLG,TON,TEN,ALFI,XIN)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION
1 XPO(10),YPO(10),ZPO(10),UMT(224),VMT(224),UVM(180),XIN(6)
2,CNM(15),SNM(15),CTT(15),CTB(15),LCT(6),ICT(8),XM(6)
DIMENSION CMC(3),CMI(3)
COMMON/ORBPAR/TZE,STEP,ALF,OMG,EPS,CUD,CUV,CUT,XM,XPO,YPO,ZPO,UVM
1,CMC
COMMON/ERDCON/ERDI,XMUI,OMGI,XCM,YCM,ZCM
EQUIVALENCE(CMI(1),XCM)
C VALUES OF SPHERICAL HARMONIC COEFFICIENTS FROM SAO 1969 STANDARD EARTH
C 20.5 FIELD SAO 1969
DATA CNM/1.000,0.000,-.1082628D-02,0.25380D-05,0.15930D-05,2*0.00,
1 0.21276D-5,-.50270D-06,0.15575D-05,.30469D-06,.73844D-07,
2 .95700D-07,.59130D-07,-.16838D-08/
DATA SNM/7*0.000,.28099D-06,-.46263D-06,-.88052D-06,-.21678D-06,
1 .15794D-06,.19946D-06,-.92433D-08,.71686D-08/
C ARRAYS OF CONSTANT DEPENDING ON DEGREE AT WHICH GRAVITY FIELD IS TRUNCATED
DATA LCT/5,4,3,2,1,0/
DATA CTT/-2.000,-4.000,-6.000,-8.000,-10.000,-2.000,-4.000,
1 -6.000,-8.000,-2.000,-4.000,-6.000,-2.000,-4.000,-2.000/
DATA CTB/2.000,6.000,12.000,20.000,30.000,2.000,6.000,12.000,
1 20.000,2.000,6.000,12.000,2.000,6.000,2.000/
CNM(2)=ZCM/ERDI
CNM(6)=XCM/ERDI
SNM(6)=YCM/ERDI
CCC USE THIS VALUE FOR COMPARISON PURPOSES ONLY
CNM(5)=0.15903E-05
KFLG=1
NTE=5
NHT=0
KRG=0
KDR=0
KV=0
KT=10
KTR=KT
C NFLG=0 MEANS TO COMPUTE THE CONDITIONS AT TEN FROM TON AND
C RETURN THE NEW POSITION AND VELOCITY ELEMENTS IN XIN
C NFLG=1 MEANS TO UPDATE THE ORBIT FROM TON TO TEN IF NECESSARY AND TO
C INITIALIZE THE SERIES COEFFICIENTS AT TON
C NFLG=2 INDICATES THAT THE EXPANSIONS ALREADY EXIST FOR THIS ORBIT,
C BUT THE ORBIT IS TO BE INTEGRATED UP TO TEN AND NEW EXPANSIONS
C FORMED ABOUT THAT POINT
IF(NFLG-1) 150,120,240
120 DO 130 I=1,36
130 UVM(I)=0.0
DO 140 I=1,9,4
UVM(I+27)=1.0
140 UVM(I)=1.0
150 CONTINUE
10 CONTINUE
C COMPUTE CONONICAL UNITS AND INITIALIZE CONSTANTS
CUM=XIN(1)*XIN(1)+XIN(2)*XIN(2)+XIN(3)*XIN(3)
CUD=DSQRT(CUM)
CUM=CUM*CUD
CUT=DSQRT(CUM/XMUI)
CUV=CUD/CUT

```

```

ERD=ERDI/CUD
XMU=1.0
ALF=ALFI
OMG=OMGI*CUT
EPS=0.04/CUD
XPO(1)=XIN(1)/CUD
YPO(1)=XIN(2)/CUD
ZPO(1)=XIN(3)/CUD
XPO(2)=XIN(4)/CUV
YPO(2)=XIN(5)/CUV
ZPO(2)=XIN(6)/CUV
TIN=TON/CUT
TFI=TEN/CUT
DO 180 I=1,3
180 CMC(I)=CMH(I)/CUD
200 CONTINUE
TZE=TIN
DELT=TFI-TZE
IF(DABS(DELT).GT.0.0) GO TO 220
KFLG=2
IF(NFLG.NE.0) GO TO 220
C      SET UP RETURN FOR NFLG=0
XIN(1)=XPO(1)*CUD
XIN(2)=YPO(1)*CUD
XIN(3)=ZPO(1)*CUD
XIN(4)=XPO(2)*CUV
XIN(5)=YPO(2)*CUV
XIN(6)=ZPO(2)*CUV
RETURN
220 CONTINUE
KEY=KFLG
CALL EXPAND (XPO,YPO,ZPO,CNM,SNM,LCT,ICT,UMT,VMT,CTB,CTT
1,ERD,XMU,ALF,OMG,ECC,NTE,KTR,KDR,NHT,CDC,CTW,KEY,DMT,KRG,CMC)
ICN=1
CALL UPDATE (ICN,KTR,EPS,STEP,XPO,YPO,ZPO,XM(1),XM(3),XM(5))
IF(KFLG.NE.2) GO TO 250
CALL VARIEQ(XPO,YPO,ZPO,CNM,SNM,LCT,ICT,UMT,VMT,UVH,ERD,ALF,OMG,CD
1C,CTW,NTE,KTR,KDR)
C      WRITE(6,356) TFI,STEP,ALF,OMG,EPS,CUD,CUV,CUT,XPO,YPO,ZPO,UVH
356 FORMAT(2(4E20.12/),6(5E20.12/),30(6E20.12/))
RETURN
C      ENTER HERE WITH NFLG=2
240 TIN=TON
TFI=TEN
TZE=TIN
DELT=TFI-TZE
C      INTEGRATE ORBIT
250 DEL=DELT
IF(DABS(DELT).GT.STEP) DEL=DSIGN(STEP,DELT)
ICN=2
CALL UPDATE(ICN,KTR,EPS,DEL,XPO,YPO,ZPO,XM(1),XM(3),XM(5))
TIN=TIN+DEL
ALF=ALF+DEL*OMG
C      WRITE(6,356) TIN,DEL,ALF,OMG,EPS,CUD,CUV,CUT,XPO,YPO,ZPO,XM
XPO(1)=XM(1)
YPO(1)=XM(3)
ZPO(1)=XM(5)

```

XPO(2)=XM(2)
YPO(2)=XM(4)
ZPO(2)=XM(6)
GO TO 200
END

```
SUBROUTINECLEAR(A,N)
IMPLICITREAL*8(A-H,O-Z)
C
C      FILL A  ARRAY WITH FLOATING POINT ZEROES.
C
C
C      DIMENSIONA(1)
C
      DO10J=1,N
10      A(J)=0.
      RETURN
      END
```

```

SUBROUTINE DRIVER (NN,TM,GH,DT,XOT,VEM)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION XOT(1),VEM(1)
DIMENSION XM(6),XPO(10),YPO(10),ZPO(10),UVM(180)
DIMENSION AMT(36)
COMMON /ORBPAR/TEP,DLT,ALF,OMG,EPS,CUD,CUV,CUT,XM,XPO,YPO,ZPO,UVM,
1 CMC(3)
50 TOT=TM/CUT
51 DEL=TOT-TEP
NFLG=1
IF(DABS(DEL).LT.DLT) GO TO 67
C TIME FROM T ZERO IS TOO LARGE FOR CONVERGENCE OF SERIES
C INCREMENT T ZERO BY ONE STEP
DEL=DSIGN(DLT,DEL)
NFLG=2
67 CONTINUE
KTR=9
CALL MATRUP(KTR,DEL,UVM,AMT)
IF(NFLG.EQ.1) GO TO 69
DO 68 I=1,36
68 UVM(I)=AMT(I)
TON=TEP
TEN=TEP+DEL
ALFI=ALF
CALL ORBIT(NFLG,TON,TEN,ALFI,XOT)
GO TO 51
69 CONTINUE
GH=ALF+DEL*OMG
KTR=10
ICN=2
100 FORMAT(6E18.5)
CALL UPDATE (ICN,KTR,EPS,DEL,XPO,YPO,ZPO,XM(1),XM(3),XM(5))
XOT(1)=XM(1)*CUD
XOT(2)=XM(3)*CUD
XOT(3)=XM(5)*CUD
XOT(4)=XM(2)*CUV
XOT(5)=XM(4)*CUV
XOT(6)=XM(6)*CUV
J=10
K=19
DO 70 I=1,9
VEM(J)=AMT(I)
VEM(K)=AMT(J)*CUT
VEM(I)=-VEM(J)
J=J+1
K=K+1
70 CONTINUE
DO 80 I=1,9,4
80 VEM(I)=VEM(I)+1.0
RETURN
END

```

```

SUBROUTINE UPDATE(ICN,KTR,EPS,DEL,XPO,YPO,ZPO,XOT,YOT,ZOT)
IMPLICITREAL*8(A-H,O-Z)
DIMENSIONXPO(1),YPO(1),ZPO(1),XOT(1),YOT(1),ZOT(1)
C ICN=1 - COMPUTE STEP (DEL)      KTR - NUMBER OF TERMS TO USE
C ICN=2 - INTERGRATE POS.,VEL.    EPS - TRUNCATION ERROR LIMIT
C ICN=3 - DO BOTH TOGETHER      DEL - VALID STEP SIZE
C      INITIAL AND FINAL ADDRESSES FOR X,Y,Z MAY BE SYNONYMOUS.
C
      KA=KTR
      KB=ICN-2
      IF(KB)10,20,10
10 CONTINUE
C SCALE DATA TO REMAIN IN COMPUTER RANGE
      B=1.0
      A=DSQRT((B*XPO(KA))**2+(B*YPO(KA))**2+(B*ZPO(KA))**2)
      A=DLOG((B*EPS)/A)
      B=KA-1
      DEL=DEXP(A/B)
      IF(KB)40,20,20
20 XPI=XPO(KA)
      YPI=YPO(KA)
      ZPI=ZPO(KA)
      KA=KA-1
      KB=KA
      A=KA
      XVI=XPI*A
      YVI=YPI*A
      ZVI=ZPI*A
      DT=DEL
      DO30I=2,KB
      XPI=XPI*DT+XPO(KA)
      YPI=YPI*DT+YPO(KA)
      ZPI=ZPI*DT+ZPO(KA)
      A=KA-1
      XVI=XVI*DT+A*XPO(KA)
      YVI=YVI*DT+A*YPO(KA)
      ZVI=ZVI*DT+A*ZPO(KA)
      KA=KA-1
30 CONTINUE
      XOT(1)=XPI*DT+XPO(1)
      YOT(1)=YPI*DT+YPO(1)
      ZOT(1)=ZPI*DT+ZPO(1)
      XOT(2)=XVI
      YOT(2)=YVI
      ZOT(2)=ZVI
40 RETURN
      END

```

```
SUBROUTINE MATRUP(KTR,DEL,UVM,TAR)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION TAR(36),UVM(1),UVO(1)
C SUBROUTINE TO UPDATE THE MATRIZANT WITH RESPECT TO TIME (DEL).
C
C
L=KTR
N=18*(L-1)+1
DO10 I=1,18
TAR(I)=UVM(N)
TAR(I+18)=0.0
10 N=N+1
DO20 I=2,L
N=18*(L-I)+1
M=N+18
TM=L-I+1
DO20 K=1,18
TAR(K)=DEL*TAR(K)+UVM(N)
TAR(K+18)=DEL*TAR(K+18)+TM*UVM(M)
N=N+1
M=M+1
20 CONTINUE
RETURN
END
```

```
SUBROUTINE ROT3(ANG,X)
IMPLICIT REAL*8(A-H,D-Z)
DIMENSION X(3),XX(3)
COS=DCOS(ANG)
SIN=DSIN(ANG)
XX(1)=COS*X(1)+SIN*X(2)
XX(2)=-SIN*X(1)+COS*X(2)
DO 100 I=1,2
100 X(I)=XX(I)
RETURN
END
```

```

SUBROUTINE ORBRN
C FORM REDUCED NORMAL EQUATIONS FOR UP TO 40 STATIONS
C FOR SHORT ARC MODE PROCESSING
  IMPLICIT REAL*8(A-H,O-Z)
  COMMON/NSTA/NSTA
  INTEGER*2 PCODE(20)
  COMMON/PCODES/PCODE
  COMMON/WPW/WPW,XPU,IDEKF,IFSTA
  INTEGER CONTIN,ENDSIG/IHE/
  COMMON/STAORD/KORDER(150)
  COMMON/NORMEQ/REDN(3,3,820),U(3,40),L(820),LSOLVE
  DIMENSION BN(3,16,15),DUMMY(3,3,40),DDN(16,16),DDK(16),BNDDNI(3,
1 16),DDK3(3)
  INTEGER*2 L,LSOLVE,LG(40),IDAY,IYR,KSTATO(15)
  DIMENSION ORBNAM(6)
  EQUIVALENCE(BN(1,1,1),DUMMY(1,1,1)),(DDK3(1),DDK(1))
  LOC(K)=(K*(K+1))/2
  MAXSTA=40
  IF(NSTA.GT.MAXSTA) GO TO 901
C
C THE REDUCED NORMAL EQUATIONS ARE STORED AS 3 X 3 BLOCKS IN THE ARRAY REDN.
C ONLY THE UPPER TRIANGULAR PART OF THE REDUCED NORMAL EQUATIONS IS STORED.
C THE BLOCKS OF THE REDUCED NORMAL EQUATIONS ARE NUMBERED
C ACCORDING TO THE FOLLOWING SCHEME:
C
C
C      1   2   4   7   11
C          3   5   8   12
C              6   9   13
C                  10  14
C                      15      ET CETERA
C
C L(820) IS THE GUIDE MATRIX
C L=1 SIGNIFIES A NON ZERO BLOCK
C L=0 SIGNIFIES A ZERO BLOCK
  REWIND 1
  REWIND 2
  IB=LOC(NSTA)
  DO 100 JB=1,IB
  DO 99 I=1,3
  DO 99 J=1,3
  99 REDN(I,J,JB)=0.0
  100 L(JB)=0
C
C     READ (2) DUMMY,U
C
C STASH DIAGONAL BLOCKS
  DO 110 KSTA=1,NSTA
  IB =LOC(KSTA)
  DO 108 I=1,3
  DO 108 J=1,3
  108 REDN(I,J,IB)=DUMMY(I,J,KSTA)
  110 CONTINUE
C
  IF(PCODE(14).EQ.0) GO TO 130
  READ(2) DUMMY
  NN=NSTA-1
  IB=LOC(NN)
  DO 120 KSTA=1,NN

```

```

NB=IB+KSTA
DO 120 I=1,3
DO 120 J=1,3
120 REDN(I,J,NB)=DUMMY(I,J,KSTA)
130 CONTINUE
C
FDEGF=IDEFG
IF(PCODE(9).EQ.1) WRITE(7,7010) FDEGF,WPH
7010 FORMAT(16X,2F16.6)
C   READ BLOCKS FROM EACH ORBIT AND REDUCE NORMAL EQUATIONS.
C
150 READ (1) IORB,NORB,ORBNAME,IDAY,MONTH,IYR,IH,MIN,ESEC,EPOCH,CONTIN
C   IF END OF DATA, GO OUT OF LOOP
IF(CONTIN.EQ.ENDSIG) GO TO 400
READ (1) DDN,DDK,BN
READ (1) NUTORB,NSTATO,NEMUNK,KSTATO
DO 180 IS=1,NSTATO
ISTA=KSTATO(IS)
C   GET BN * DDN(INVERSE)
DO 160 I=1,3
DO 160 J=1,NUTORB
BNDDNI(I,J)=0.0
DO 160 K=1,NUTORB
160 BNDDNI(I,J)=BNDDNI(I,J)+BN(I,K,IS)*DDN(K,J)
DO 165 I=1,3
DO 165 K=1,NUTORB
165 U(I,ISTA)=U(I,ISTA)-BNDDNI(I,K)*DDK(K)
DO 180 JS=1,NSTATO
JSTA=KSTATO(JS)
C   SKIP IF (ISTA.GT.JSTA), SINCE ONLY THE UPPER TRIANGULAR PART OF THE
C   REDUCED NORMAL EQUATIONS IS BEING COMPUTED AND SAVED.
IF(ISTA.GT.JSTA) GO TO 180
NB=LOC(JSTA-1)+ISTA
DO 170 I=1,3
DO 170 J=1,3
DO 170 K=1,NUTORB
170 REDN(I,J,NB)=REDN(I,J,NB)-BNDDNI(I,K)*BN(J,K,JS)
L(NB)=L(NB)+1
180 CONTINUE
C   RETURN TO PROCESS ANOTHER PASS
GO TO 150
C
C   ENTER HERE WHEN ALL PASSES HAVE BEEN PROCESSED
400 CONTINUE
C
C   SIMULATE KRAKOWSKI'S GUIDE MATRIX
IF(PCODE(6).NE.1) GO TO 441
C
WRITE(6,6001)
6001 FORMAT(1H1,10(/),20X,'GUIDE MATRIX')
DO 440 ISTA=1,NSTA
IB=0
LG(1)=1000
DO 435 JSTA=ISTA,NSTA
JB=LOC(JSTA-1)+ISTA
IF(L(JB).EQ.0) GO TO 435
IB=IB+1
LG(IB)=KORDER(JSTA)

```

```

435 CONTINUE
C
  IB=IB+1
  IF(IB.GT.1) LG(IB)=999
  439 WRITE(6,6002) KORDER(ISTA),(LG(I),I=1,IB)
  6002 FORMAT(20X,I5,5X,18I5,200(/30X,18I5))
  440 CONTINUE
  441 CONTINUE
C
C  PRINT NORMALS IN ASD FORMAT, AND PUNCH IF DESIRED.
  WRITE(6,6003)
  6003 FORMAT(1H1//          NORMAL EQUATIONS (SEE GUIDE MATRIX) '///)
  DO 450 ISTA=1,NSTA
  DO 442 I=1,3
  442 DDK(I)=-U(I,ISTA)
  IB=0
  JB=LOC(ISTA)
  IF(L(JB).GT.0) IB=1
C  PUNCH NORMALS
  IF(PCODE(9).NE.1) GO TO 443
  WRITE(7,7001) KORDER(ISTA)
  7001 FORMAT(14I5)
  WRITE(7,7006) DDK3
  7006 FORMAT(3(F16.10,5X))
  WRITE(7,7008) ((REDN(I,J,JB),J=1,3),I=1,3)
  7008 FORMAT(3F16.10/3F16.10/3F16.10)
C
  443 CONTINUE
C  PRINT DIAGONAL BLOCK
  IF(PCODE(7).NE.1) GO TO 444
  WRITE(6,6004) KORDER(ISTA)
  6004 FORMAT(//I5)
  WRITE(6,6006) DDK3
  6006 FORMAT(/3(F16.10,5X))
  WRITE(6,6008) ((REDN(I,J,JB),J=1,3),I=1,3)
  6008 FORMAT(3F16.10)
  444 CONTINUE
C  PRINT OFF-DIAGONAL BLOCKS
  KSTA=ISTA+1
  IF(ISTA.EQ.NSTA) GO TO 448
  DO 445 JSTA=KSTA,NSTA
  JB=LOC(JSTA-1)+ISTA
  IF(L(JB).EQ.0) GO TO 445
  IB=IB+1
  IF(PCODE(9).NE.1) GO TO 7445
  WRITE(7,7001) KORDER(JSTA)
  WRITE(7,7008) ((REDN(I,J,JB),J=1,3),I=1,3)
  7445 CONTINUE
  IF(PCODE(7).NE.1) GO TO 445
  WRITE(6,6004) KORDER(JSTA)
  WRITE(6,6008) ((REDN(I,J,JB),J=1,3),I=1,3)
  445 CONTINUE
  448 I=1000
  IF(IB.GT.0) I=999
  IF(PCODE(7).EQ.1) WRITE(6,6004) I
  IF(PCODE(9).EQ.1) WRITE(7,7001) I
  450 CONTINUE
  IF(PCODE(8).NE.1) GO TO 478

```

```
      WRITE(6,6010)
6010 FORMAT(10(/),20X,'OBSERVATIONS ON EACH LINE')
      IB=NSTA-1
      DO 475 ISTA=1,IB
      KSTA=ISTA+1
      DO 475 JSTA=KSTA,NSTA
      WRITE(6,6011) KORDER(ISTA),KORDER(JSTA),L(LOC(JSTA-1)+ISTA)
6011 FORMAT(8I10)
475 CONTINUE
478 CONTINUE
      RETURN
901 CONTINUE
      WRITE(6,9001) MAXSTA,NSTA
9001 FORMAT(' FORMRN IS PRESENTLY DIMENSIONED TO HANDLE ONLY',15,
1      ' UNKNOWN STATIONS.',/20X,' THIS PROBLEM HAS',15,' UNKNOWN STAT
2 IONS.',/10X,'EXECUTION IS TERMINATED BY PROGRAM.')
      STOP
      END
```

```

SUBROUTINE PORB
C   PRINT UPDATED ORBIT ELEMENTS AND ERROR MODEL TERMS
IMPLICIT REAL*8(A-H,O-Z)
COMMON/NSTA/NSTA
COMMON/WPW/WPW,XPU,IDEKF,NFSTA
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
INTEGER STANAM,IDS*2
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1STANAM(5,150),IDS(150)
COMMON/STAORD/KORDER(150)
DIMENSION DX(3,40), DDN(16,16),DDK(16),BN(3,16,15),DDKM(16)
INTEGER CONTIN,ENDSIG/1HE/
INTEGER*2 LFLG(40),KSTATO(15),MODEL(10,2),IGUIDE(10,2)
INTEGER*2 IDAY,IYR
DIMENSION EMODEL(10,2),ORBNAME(6),ORBUNK(16),DB(16)
DIMENSION COVX(3,3)
INTEGER*4 MODALF(3,2)
DATA MODALF/'ZERO','SET','REFR','ACTI','ON'/
REWIND 1
REWIND 2
REWIND 4
MAXUNK=16
WRITE(6,6001)
6001 FORMAT(1H1,////10X,'CORRECTIONS TO ORBIT AND ERROR MODEL UNKNOWNNS
1')
C   SKIP HEADER RECORD ON 2
READ(2)
C   GET CORRECTIONS AT EACH STATION
DO 120 ISTA=1,NSTA
120 READ(2)(DX(I,ISTA),I=1,3),COVX
REWIND 2
C   BEGIN PROCESSING PASSES
150 CONTINUE
C READ ORBIT HEADER
READ (1) NORB,NORB,ORBNAME,IDAY,MONTH,IYR,IH,MIN,ESEC,EPOCH,CONTIN
IF(CONTIN.EQ.ENDSIG) GO TO 700
READ (1) DDN,DDK,BN
C   GET PREVIOUS SET OF UNKNOWNNS
READ (1) NUTORB,NSTATO,NEMUNK,KSTATO,LFLG,ORBUNK,EMODEL,MODEL,
1 IGUIDE
WRITE(6,6002) NORB,ORBNAME,IDAY,MONTH,IYR,IH,MIN,ESEC,EPOCH
6002 FORMAT(/////////2X,A4,3X,6A8,' EPOCH=',I3,A3,1X,2I3,'H',I3,'M',
1 F8.4,'S UT=MJD',F17.9)
COPY DDK INTO DDKM
DO 170 I=1,MAXUNK
170 DDKM(I)=DDK(I)
DO 180 IS=1,NSTATO
ISTA=KSTATO(IS)
DO 180 I=1,NUTORB
DO 180 J=1,3
180 DDKM(I)=DDKM(I)-BN(J,I,IS)*DX(J,ISTA)
CALL DGMPRD(DDN,DDKM,DB,MAXUNK,MAXUNK,1)
WRITE(6,6003)(DB(I),I=1,NUTORB)
6003 FORMAT(' CORRECTION VECTOR',100(/6F20.8))
C   UPDATE ORBIT AND ERROR MODEL UNKNOWNNS
DO 210 I=1,NUTORB
210 ORBUNK(I)=ORBUNK(I)+DB(I)

```

```

C PRINT UPDATED ELEMENTS
  WRITE(6,6004) (ORBUNK(I),I=1,6)
6004 FORMAT('0 UPDATED ORBIT ELEMENTS IN APPARENT SIDEREAL CARTESIAN
1 COORDINATES'/5X,'POSITION(METERS)',4X,3F16.3'
1 5X,'VELOCITY(METERS/SEC)',3F16.6)
  IF(PCODE(15).EQ.0) GO TO 220
C PUNCH UPDATED ORBIT ELEMENTS.
  WRITE(7,7001) NORB,ORBNAM
7001 FORMAT(A4,6A8)
  WRITE(7,7002) (ORBUNK(I),I=1,6)
7002 FORMAT(3F15.3/3F15.6)
220 CONTINUE
  IF(NEMUNK.EQ.0) GO TO 230
  WRITE(6,6005)
6005 FORMAT('0 UPDATED VALUE OF ERROR MODEL UNKNOWN'S)
  DO 229 I=1,NEMUNK
    MODCOD=MODEL(I,1)
    KSTA=MODEL(I,2)
229  WRITE(6,6006) (MODALF(J,MODCOD),J=1,3),KORDER(KSTA),(STANAM(J,KSTA
1),J=1,5),ORBUNK(I+6)
6006 FORMAT(5X,3A4,'FOR STATION',I6,2X,5A4,2X,'=',F15.3)
230 CONTINUE
C UPDATE PASS RECORD ON UNIT 4
  WRITE(4)ORBUNK,NSTATO,LFLG,KSTATO,NEMUNK,EMODEL,MODEL,IGUIDE
  1 , NUTORB
C RETURN TO PROCESS ANOTHER PASS
  GO TO 150
C
C ENTER HERE WHEN ALL PASSES HAVE BEEN PROCESSED
700 CONTINUE
  REWIND 1
  REWIND 4
  IF(PCODE(14).EQ.0) GO TO 730
C UPDATE COORDINATES OF CENTER OF MASS
  DO 710 I=1,3
710 STAUVW(I,NSTA)=STAUVW(I,NSTA)+DX(I,NSTA)
  WRITE(6,6007) (STAUVW(I,NSTA),I=1,3),COVX
6007 FORMAT(///5X,'UPDATED COORDINATES OF THE CENTER OF MASS'/
1 5X,3F15.3//5X,'WEIGHT COEFFICIENT MATRIX'/3(5X,3F15.3/))
C RESET NSTA TO ACTUAL NUMBER OF GROUND OBSERVING STATIONS
  NSTA=NSTA-1
730 CONTINUE
  RETURN
  END

```

```
FUNCTIONMJD(DATE,MONTH,YEAR)
COMPUTATION OF MODIFIED JULIAN DAY
INTEGER*2 DATE,YEAR
DIMENSIONMONTHS(2,12)
DATAMONTHS/3HJAN,0,3HFEB,31,3HMAR,59,3HAPR,90,3HMAY,120,3HJUN,151,
13HJUL,181,3HAUG,212,3HSEP,243,3HOCT,273,3HNOV,304,3HDEC,334/
ID=365*(YEAR-50)+(YEAR-49)/4
DO20I=1,12
IF(MONTH.EQ.MONTHS(1,I))GOTO25
20 CONTINUE
IF(MONTH.LE.12) GO TO 21
WRITE(6,6001) MONTH
6001 FORMAT(3X,22HMONTH NAME MISPELLED ,A3)
STOP
21 I=MONTH
MONTH=MONTHS(1,I)
25 CONTINUE
ID=ID+MONTHS(2,I)
IF(MOD(YEAR*1,4).EQ.0.AND.I.GT.2) ID=ID+1
MJD=ID+DATE+33281
RETURN
END
```

DOUBLE PRECISION FUNCTION GSTD(1)
COMPUTATION OF GREENWICH SIDEREAL TIME
DOUBLE PRECISION, DGST
REAL*8 PI2/6.28318530717958/
DGST=0.277987616D0+1.002737811910*(T-0.33282D5)
C LINEAR TERM COEFFICIENT SHOULD BE 1.002737811906 BY A.E. SUPP.
DGST=DGST-DBLE(IFLOAT(1DINT(DGST)))
GSTD=DGST*PI2
RETURN
END

```

SUBROUTINE PRENUT(DT,PAN)
COMPUTATION OF PRECESSION AND NUTATION SINCE 1950.0
C   THE DIFFERENCE BETWEEN MODIFIED AND TRUE SIDEREAL
C   TIME
C   DT IS MODIFIED JULIAN DATE OF EPOCH
    IMPLICIT REAL*8(A-H,O-Z)
    DIMENSION FUNARG(5)
    REAL*4 COEFF(5,13)           /4*0.0,1.0,4*0.0,2.0,2*0.0,2.0,-2.0
1,2.0,0.0,1.0,4*0.0,1.0,2.0,-2.0,2.0,0.0,-1.0,2.0,-2.0,2.0,2*0.0,2.
20,-2.0,1.0,2*0.0,2.0,0.0,2.0,1.0,6*0.0,2.0,0.0,2*1.0,0.0,2.0,0.0,2
3.0,1.0,2*0.0,-2.0,0.0,-1.0,0.0,2.0,0.0,2.0/
    REAL*4 TCOEF(2,13)          /-172327.0,-173.7,2088.0,0.2,-12729
1.0,-1.3,1261.0,-3.1,-497.0,1.2,214.0,-0.5,124.0,0.1,-2037.0,-0.2,6
275.0,0.1,-342.0,-0.4,-261.0,0.0,-149.0,0.0,114.0,0.0/
    REAL*8 FUNCOF(3,5)          /.82251280093,.362916456847160-1,19
113865.0D-20,.99576620370,.2737778519279D-2,-31233.0D-20,.031252469
214,.036748195691688,-668609.0D-20,.97427079475,.033863192198393,-2
399023.0D-20,.71995354167,-0.147094228332D-3,432630.00-20/
    DATARPS/4.8481368D-6/,TPI/6.28318530717958/
    COSE=0.91739033
    NTERMS=13
    BT=DT-15019.5
    BT2=BT*BT
    TT=BT/36525.0
    DO50I=1,5
    FARG=FUNCOF(1,I)+FUNCOF(2,I)*BT+FUNCOF(3,I)*BT2
50  FUNARG(I)=(FARG-DBLE(FLOAT(IDINT(FARG))))*TPI
C
    DLONG=0.0
    DO80I=1,NTERMS
    ARG=0.0
    DO60J=1,5
60  ARG=ARG+COEFF(J,I)*FUNARG(J)
    TERM=(TCOEF(1,I)+TCOEF(2,I)*TT)*DSIN(ARG)
    DLONG=DLONG+TERM
80  CONTINUE
    DMU=DLONG*COSE*0.0001
C   COMPUTE PRECESSION SINCE 1950.0 * KAPPA & OMEGA
    BY=(DT-0.33281923D5)/365.2422D0
    PRECES=(46.0990+1.39E-4*BY)*BY
    PAN=(PRECES+DMU)*RPS
    RETURN
    END

```

C SUBROUTINE UVWD(A,B,PHI,LAMDA,H,U,V,W)
DOUBLE PRECISION VERSION OF UVW JAN 5, 1968
DOUBLEPRECISIONPHI,LAMDA,N,E2,FAC,U,V,W,SP
REAL*8 A,B,H
E2=1.0-(B/A)**2
SP=DSIN(PHI)
N=A/DSQRT(1.0-E2*SP*SP)
FAC=(N+H)*DCOS(PHI)
U=FAC*DCOS(LAMDA)
V=FAC*DSIN(LAMDA)
W=(N*(1.0-E2)+H)*SP
RETURN
END

```

SUBROUTINE DEDIT
IMPLICIT REAL*8(A-H,O-Z)
COMMON/DEDITC/ALFS(50),DEC(50),U(3,50),S(3),D(50),SDC(3,50),SUM,
1GAST,STAXYZ(3,50),GQI,
2TD,KSTATE(50),IPASS(50),NSTE,NSUSED,KODE
C EDIT DATA BASED ON PRELIMINARY STATION POSITIONS AND DELETE BAD
C OBSERVATIONS AND BAD EVENTS, BASED ON THE DISTANCE CRITERION TD
C THIS SUBROUTINE IS DIMENSION FOR A MAXIMUM OF MAXSTE=50 STATIONS
C PARTICIPATING IN ANY ONE EVENT. ALL AFFECTED ARRAYS ARE IN
C COMMON BLOCK /DEDITC/.
C
C THE NUMBER OF STATIONS PARTICIPATING IN THE EVENT IS NSTE.
C THE NUMBER OF STATIONS NOT DELETED IS NSUSED.
C
COMMON/STALOC/STAUVH(3,150)
DIMENSION Q(3,3),RHS(3),QI(3,3),VI(3)
C
MAXSTE=50
C INITIALIZE
KODE=1
DO 110 IS=1,NSTE
110 IPASS(IS)=1
C IPASS=1 MEANS THIS DIRECTION OK
C IPASS=2 MEANS THIS DIRECTION DELETED FROM EVENT
C
C FORM UNIT VECTORS FOR ALL DIRECTIONS IN THIS EVENT
DO 125 IS=1,NSTE
STS=ALFS(IS)-GAST
CA=DCOS(STS)
SA=DSIN(STS)
CD=DCOS(DEC(IS))
SD=DSIN(DEC(IS))
U(1,IS)=CA*CD
U(2,IS)=SA*CD
U(3,IS)=SD
125 CONTINUE
C
C INITIALIZE ARRAYS FOR THIS ITERATION
130 CONTINUE
NSUSED=0
DO 140 I=1,3
RHS(I)=0.0
S(I)=0.0
DO 140 J=1,3
Q(I,J)=0.0
140 CONTINUE
C
C ACCUMULATE EQUATIONS
DO 190 IS=1,NSTE
IF(IPASS(IS).EQ.2) GO TO 190
NSUSED=NSUSED+1
DO 170 I=1,3
DO 169 J=1,3
169 QI(I,J)=U(I,IS)*U(J,IS)
170 QI(I,I)=QI(I,I)-1.0
DO 175 I=1,3
DO 175 J=1,3
Q(I,J)=Q(I,J)+QI(I,J)

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RHS(I)=RHS(I)+QI(I,J)*STAXYZ(J,IS)
175 CONTINUE
190 CONTINUE
C
C TEST FOR DELETION OF WHOLE EVENT
IF(NSUSED.LT.2) GO TO 420
C
C INVERT AND SOLVE
C THE SATELLITE POSITION S IS SELECTED IN SUCH A WAY THAT THE SUM OF
C THE SQUARES OF THE DISTANCES FROM S OF THE NON-DELETED RAYS IS MINIMIZED.
DET=1.0
CALL DMINV(Q,3,DET,QI(1,1),QI(1,2))
GQI=DABS(DET/DFLOAT(NSUSED))
IF(GQI.LT.1.0D-4) GO TO 430
CALL DGMPRD(Q,RHS,S,3,3,1)
C
C COMPUTE DISTANCE FROM S FOR EACH RAY
ISMAX=0
DMAX=0.0
SUM=0.0
DO 280 IS=1,NSTE
DO 270 I=1,3
DO 269 J=1,3
269 QI(I,J)=U(I,IS)*U(J,IS)
QI(I,I)=QI(I,I)-1.0
VI(I)=S(I)-STAXYZ(I,IS)
270 CONTINUE
DDI=DPDOT(VI,U(1,IS),3)
DDI=DABS(DDI)
DI=0.0
DO 275 I=1,3
DI=DI+(VI(I)-DDI*U(I,IS))**2
SDC(I,IS)=VI(I)
275 CONTINUE
D(IS)=DSQRT(DI)/DDI*206264.80625
IF(IPASS(IS).EQ.2) GO TO 280
SUM=SUM+DI
C TEST D AGAINST TD AND DELETE IF NECESSARY
IF(D(IS).LT.DMAX) GO TO 280
DMAX=D(IS)
ISMAX=IS
280 CONTINUE
IF(DMAX.LT.TD) RETURN
IPASS(ISMAX)=2
C
C GO BACK AND MAKE ANOTHER PASS THROUGH THE DATA
GO TO 130
400 CONTINUE
C DELETE WHOLE EVENT
DO 410 IS=1,NSTE
410 IPASS(IS)=2
NSUSED=0
RET RN
420 CONTINUE
C DELETE FOR INSUFFICIENT GOOD OBSERVATIONS
KODE=2
GO TO 400
C DELETE FOR INSUFFICIENT GEOMETRICAL SEPARATION BETWEEN OBSERVATIONS

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SUBROUTINE RODATA
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 PCODE(120)
COMMON/PCODES/PCODE
INTEGER*4 ENDSIG/1HE/,CONTIN,DELCOD(2)/1H ,1H*/,ECODE
INTEGER*2 PLUS/1H+/
INTEGER*2 ISGNP,IPHID,IPHIM,LONGD,LONGM,ISGNL
INTEGER*2 ID(50),KEY(50),IHR(50),MIN(50),IDAY(50),IYR(50),IRAH(50)
1,IRAM(50),ISGND(50),IDEC(50),IDECM(50),IDAT(50,11)
COMMON/DEDITC/ALFS(50),DEC(50),U(3,50),S(3),D(50),SDC(3,50),EVSUM,
1GAST,STAXYZ(3,50),GQI,
2TD,KSTATE(50),IPASS(50),NSTE,NSUSED,ECODE
COMMON/NSTA/NSTA
COMMON/STAORD/KORDER(150)
INTEGER STANAM,IDS*2
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1STANAM(5,150),IDS(150)
DIMENSION PM(3,3),AP(2,3)
DIMENSION MONTH(50)
EQUIVALENCE(ID(1),IDAT(1,1)),(KEY(1),IDAT(1,2)),(IHR(1),IDAT(1,3))
1,(MIN(1),IDAT(1,4)),(IDAY(1),IDAT(1,5)),(IYR(1),IDAT(1,6)),(IRAH(1
2),IDAT(1,7)),(IRAM(1),IDAT(1,8)),(ISGND(1),IDAT(1,9)),(IDEC(1),ID
3AT(1,10)),(IDECM(1),IDAT(1,11))
DIMENSION SEC(50),RAS(50),DECS(50),VARRA(50),VARDEC(50),COVRAD(50)
1,DAT(50,6)
EQUIVALENCE(SEC(1),DAT(1,1)),(RAS(1),DAT(1,2)),(DECS(1),DAT(1,3)),
1(VARRA(1),DAT(1,4)),(VARDEC(1),DAT(1,5)),(COVRAD(1),DAT(1,6))
COMMON/OBSD/OBSD(150),OVOBSD
C
IF(PCODE(1).EQ.1) GO TO 3
IF(PCODE(1).EQ.7) GO TO 3
RETURN
C
3 MAXSTE=50
PI=3.14159265358D0
SPR=206264.80625D0
PI2=2.0*PI
WPWSP=0.0
C
READ(5,5004) TD,OVOBSD
WRITE(6,6004) TD
5004 FORMAT(F20.2,F10.2)
6004 FORMAT(//20X,'TEST DISTANCE =',F20.2,'    SECONDS OF ARC')
WRITE(3) TD
C START DATA INPUT
IEVENT=0
KEVENT=0
EPR=0.0
IS=0
C
C ENTER HERE FOR A NEW OBSERVATION
C
200 IS=IS+1
IF(PCODE(1).EQ.7) GO TO 211
C ENTER HERE IF THE OPTICAL DATA IS IN THE KRAKISKY FORMAT
205 CONTINUE
READ(5,1021,END=901) ID(IS),KEY(IS),IHR(IS),MIN(IS),SEC(IS),

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XIDAY(IS),
1MONTH(IS),IYR(IS),IRAH(IS),IRAM(IS),RAS(IS),ISGND(IS),IDECD(IS),
2IDECM(IS),DECS(IS),VARRA(IS),VARDEC(IS),COVRAD(IS),CONTIN
1021 FORMAT(I3,A1,1X,I2,I3,F9.4,I2,A3,I2,2I3,F9.4,A1,I2,I3,F8.4,2X,
13F5.2,7X,A1)
1022 FORMAT(14X,I4,5I2,F6.4,1X,2I2,F5.3,A1,2I2,F4.2,26X,A1)
IF(CONTIN.EQ.ENDSIG) GO TO 250
DDT=DFLOAT(MJD(IDAY(IS),MONTH(IS),IYR(IS)))
DDT=DDT+(DFLOAT((IHR(IS)*60+MIN(IS))*60)+SEC(IS))/864.0D2
IF(IS.LE.1) GO TO 210
C THIS TEST SHOULD BE TRUE ONLY FOR THE FIRST CARD OF THE FIRST EVENT.
C
CHECK FOR END OF EVENT, ALLOWING 0.5 MS DISCREPANCY
IF(DABS(DDT-EPR).GT.0.58D-8) GO TO 250
C
C ENTER HERE TO BEGIN A NEW EVENT
C THE FIRST ENTRY OF THE EVENT SHOULD ALWAYS BE MADE WITH IS=1
210 CONTINUE
  IDD=ID(IS)
  KSTA=KSTAID( IDD)
  IF(KSTA.GT.0) GO TO 220
  WRITE(6,6042) ID(IS),IHR(IS),MIN(IS),SEC(IS),IDAY(IS),MONTH(IS),
  1IYR(IS)
6042 FORMAT(5X,'STATION NUMBER NOT FOUND IN INPUT LIST',15,3X,2I3,
  1F8.4,3X,I3,A3,I2,'OBSERVATION IGNORED')
  IF(PCODE(1).EQ.1) GO TO 205
C ENTER HERE IF THE OPTICAL DATA IS IN THE GEOS FORMAT
211 CONTINUE
  READ(5,5000,END=901) ID(IS),IYR(IS),MONTH(IS),IDAY(IS),IHR(IS),
  1MIN(IS),SEC(IS),IRAH(IS),IRAM(IS),RAS(IS),ISGND(IS),IDECD(IS),
  2IDECM(IS),DECS(IS),CONTIN,VARRA(IS),VARDEC(IS),COVRAD(IS)
5000 FORMAT(14X,I4,5I2,F6.4,I3,I2,F5.3,A1,2I2,F4.2,17X,A1,2F3.2,F3.1)
  IF(CONTIN.EQ.ENDSIG) GO TO 250
  DDT=DFLOAT(MJD(IDAY(IS),MONTH(IS),IYR(IS)))
  DDT=DDT+(DFLOAT((IHR(IS)*60+MIN(IS))*60)+SEC(IS))/864.0D2
  IF(IS.LE.1) GO TO 212
C THIS TEST SHOULD BE TRUE ONLY FOR THE FIRST CARD OF THE FIRST EVENT.
C
CHECK FOR END OF EVENT, ALLOWING 0.5 MS DISCREPANCY
IF(DABS(DDT-EPR).GT.0.58D-8) GO TO 250
C
C ENTER HERE TO BEGIN A NEW EVENT
C THE FIRST ENTRY OF THE EVENT SHOULD ALWAYS BE MADE WITH IS=1
212 CONTINUE
  IDD=ID(IS)
  KSTA=KSTAID( IDD)
  IF(KSTA.GT.0) GO TO 220
  WRITE(6,6042) ID(IS),IHR(IS),MIN(IS),SEC(IS),IDAY(IS),MONTH(IS),
  1IYR(IS)
  GO TO 221
220 CONTINUE
219 IF(PCODE(12).EQ.1) GO TO 221
  IF(PCODE(12).EQ.2) GO TO 222
  GO TO 230
221 VARRA(IS)=OBSD(KSTA)
  VARDEC(IS)=OBSD(KSTA)
  COVRAD(IS)=0.0
  GO TO 230

```

```

222 VARRA (IS)=OVOBSD
  VARDEC(IS)=OVOBSD
  COVRAD(IS)=0.0
  GO TO 230
230 CONTINUE
  KSTATE(IS)=KSTA
  EPR=DDT
  GO TO 200
C
C END OF INPUT FOR THIS EVENT. BEGIN PROCESSING
250 CONTINUE
  NSTE=IS-1
  IEVENT=IEVENT+1
  GAST=GSTD(EPR)
  CALL PRENUT(EPR,PAN)
  GAST=GAST+PAN
  CALL POLE(EPR,XP,YP)
  XP=XP/SPR
  YP=YP/SPR
  XP=0.000
  YP=0.000
COMPUTE STATION POSITION IN INSTANTANEOUS TERRESTRIAL SYSTEM
  PM(1,1)=1.
  PM(1,2)=0.0
  PM(1,3)=-XP
  PM(2,1)=0.0
  PM(2,2)=1.0
  PM(2,3)=YP
  PM(3,1)=XP
  PM(3,2)=-YP
  PM(3,3)=1.0
C
  DO 270 IS=1,NSTE
  ISGNL=PLUS
  RA=ANRADD(ISGNL,IRAH(IS),IRAM(IS),RAS(IS))*15.0
  ALFS(IS)=RA
  DEC(IS)=ANRADD(ISGND(IS),IDEC0(IS),IDECM(IS),DECS(IS))
270 CONTINUE
  WRITE(3) IEVENT,NSTE,GAST,PM,EPR,
  1((IDAT(IS,J),J=1,11),MONTH(IS),(DAT(IS,J),J=1,6),ALFS(IS),DEC(IS),
  2KSTATE(IS),IS=1,NSTE),CONTIN
C TEST FOR END OF INPUT
  IF(CONTIN.EQ.ENDSIG) GO TO 700
C PREPARE FOR NEXT EVENT
  DO 610 I=1,6
  610 DAT(1,I)=DAT(NSTE+1,I)
  MONTH(1)=MONTH(NSTE+1)
  DO 611 I=1,11
  611 IDAT(1,I)=IDAT(NSTE+1,I)
C RETURN TO START A NEW EVENT
  IS=1
  GO TO 210
C
  700 RETURN
C
C ERROR EXITS.
  901 CONTINUE
C ENTER HERE IF END SIGNAL CARD IS MISSING FROM INPUT DATA DECK

```

SUBROUTINE RCONAP

CONSTRAINT CODE DIRECTORY

WEIGHTED CONSTRAINTS

C 1 CONSTRAIN THE COORDINATES OF A STATION AT A PRIORI VALUES*(I.E. WEIGHT IT)
C 2 IMPOSE CHORD DISTANCE CONSTRAINT*.
C 3 IMPOSE RELATIVE POSITION CONSTRAINT*
C 4 IMPOSE DIRECTION CONSTRAINT*
C 5 CONSTRAIN THE GEODETIC LATITUDE, LONGITUDE AND HEIGHT OF A STATION.*

ABSOLUTE CONSTRAINTS

C 11 DEFINE THE ORIGIN OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 12 DEFINE THE ORIENTATION OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT
C 13 DEFINE THE SCALE OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 14 COMPLETELY FIX ONE OR MORE COORDINATES OF A STATION * \$
C 15 COMPLETELY FIX ONE OR MORE COORDINATES OF RELATIVE POSITION**

*IF THE COORDINATES, RELATIVE POSITION, DISTANCE, OR DIRECTION, TO BE CONSTRAINED ARE NOT GIVEN, THE CONSTRAINT IS COMPUTED FROM THE APPROXIMATE COORDINATES OF THE STATION(S) INVOLVED

\$THE DIAGONAL ELEMENTS OF THE WEIGHT MATRIX ARE WHICH COORDINATES ARE TO BE FIXED. A NON-ZERO CODE MEANS TO FIX THE COORDINATE.

```
IMPLICIT REAL*8(A-H,O-Z)
INTEGER ENDSIG/1HE/,CONTIN,STANAM,
INTEGER*2 IDS
DIMENSION XI(3),XJ(3),W(3,3),DIS(3),DXB(3),DXC(3)
COMMON/STALOC/STAUVH(3,150),DATPRM(2,15),DATNAM(4,15),
```

```
1STANAM(5,150),IDS(150)
COMMON/NSTA/NSTA/STAORD/KORDER(150)
DATA DPR/57.295779513D0/
10 READ(5,5000,END=1000)KODE,CONTIN
5000 FORMAT(I2,77X,A1)
      WRITE(3) KODE,CONTIN
      IF(CONTIN.EQ.ENDSIG) GO TO 1000
      IF(KODE.LE.0) GO TO 950
      IF(KODE.GT.19) GO TO 950
      GO TO (100,200,300,400,500,600,700,800,900,950,
1100,1200,1300,1400,1500,1600,1700,1800,1900).KODE
```

```
100 CONTINUE
    READ(5,5001) IS
5001 FORMAT(14I5)
    READ(5,5002) XI
5002 FORMAT(3D16.8)
    READ(5,5002) (W(I,I),I=1,3)
    ISTA=KSID2(IS)
    DO 110 I=1,3
        IF(XI(I).EQ.0.0) XI(I)=STAUVW(I,ISTA)
110 CONTINUE
    WRITE(3) IS,ISTA,XI,W
    GO TO 10
```

```

C                               CHORD CONSTRAINT
200 CONTINUE
  READ(5,5001) IS,JS
  READ(5,5002) CD,RELUNC
  ISTA=KSID2(IS)
  JSTA=KSID2(JS)
  CDC=0.0
  DO 205 I=1,3
    DIS(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)
205 CDC=CDC+DIS(I)**2
  CDC=DSQRT(CDC)
  IF(CD.EQ.0.0)CD=CDC
  WRITE(3) IS,ISTA,JS,JSTA,CD,RELUNC
  GO TO 10
C                               RELATIVE POSITION CONSTRAINT
300 CONTINUE
  READ(5,5001) IS,JS
  READ(5,5002) DXB
  READ(5,5002) (W(I,I),I=1,3)
  ISTA=KSID2(IS)
  JSTA=KSID2(JS)
  DO 310 I=1,3
    DXC(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)
    IF(DXB(I).EQ.0.0) DXB(I)=DXC(I)
310 CONTINUE
  WRITE(3) IS,ISTA,JS,JSTA,DXB,W
  GO TO 10
C                               400 CONTINUE
C                               DIRECTION CONSTRAINTS
C                               ALPHA IS LONGITUDE-LIKE ANGLE
C                               BETA IS LATITUDE-LIKE ANGLE
  READ(5,5001) IS,JS
  ISTA=KSID2(IS)
  JSTA=KSID2(JS)
C                               READ ANGLES IN DEGREES AND UNCERTAINTIES IN SECONDS OF ARC
  READ(5,5002) ALF,BETA
  READ(5,5002) VARA,VARB,COVAB
  DO 405 I=1,3
405 DXC(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)
  IF(ALF)412,411,412
411 ALF=DATAN2(DXC(2),DXC(1))
  ALF=ALF*DPR
412 IF(BETA)414,413,414
413 RSCSB=DXC(1)**2+DXC(2)**2
  BETA=DATAN(DXC(3)/DSQRT(RSCSB))
  BETA=BETA*DPR
414 CONTINUE
  WRITE(3) IS,ISTA,JS,JSTA,ALF,BETA,VARA,VARB,COVAB
  GO TO 10
C                               500 CONTINUE
C                               CONSTRAINT ON GEODETIC LATITUDE, LONGITUDE, AND HEIGHT
  READ(5,5001) IS
  ISTA=KSID2(IS)
  IDTS=IDS(ISTA)
C                               READ LATITUDE AND LONGITUDE IN DEGREES AND HEIGHT IN METERS

```

C AN INPUT COORDINATE OF ZERO INDICATES THAT THE APPROXIMATE VALUE OF
C THE COORDINATE IS TO BE USED.
READ(5,5002) PHIO,FLAMO,HO
C READ UNCERTAINTIES IN SECONDS OF ARC AND METERS
READ(5,5002) SDP,SDL,SDH
C AN INPUT UNCERTAINTY OF ZERO INDICATES THAT A ZERO WEIGHT IS TO BE USED.
CALL UVWTG2(STAUWV(1,ISTA),DATPRM(1,1DTS),PHI,FLAM,H)
IF(PHIO.EQ.0.0) PHIO=PHI*DPR
IF(FLAMO.EQ.0.0) FLAMO=FLAM*DPR
IF(HO.EQ.0.0) HO=H
WRITE(3) IS,ISTA,1DTS,PHIO,FLAMO,HO,SDP,SDL,SDH
GO TO 10
C
600 CONTINUE
700 CONTINUE
800 CONTINUE
900 CONTINUE
GO TO 950
C
C INNER ADJUSTMENT CONSTRAINTS
1100 CONTINUE
1200 CONTINUE
1300 CONTINUE
1600 CONTINUE
GO TO 10
C
1400 CONTINUE
C FIX A STATION
GO TO 100
C
1500 CONTINUE
GO TO 300
C
1700 CONTINUE
1800 CONTINUE
1900 CONTINUE
GO TO 950
950 WRITE(6,6095) KODE
6095 FORMAT('OILLEGAL CONSTRAINT CODE IN CONAP IGNORED',15)
GO TO 10
C
1000 CONTINUE
REWIND 3
RETURN
END

```

C
C           CONSTRAINT CODE DIRECTORY
C
C           WEIGHTED CONSTRAINTS
C
C 1  CONSTRAIN THE COORDINATES OF A STATION AT A PRIORI VALUES*(I.E.WEIGHT IT)
C 2  IMPOSE CHORD DISTANCE CONSTRAINT*.
C 3  IMPOSE RELATIVE POSITION CONSTRAINT*
C 4  IMPOSE DIRECTION CONSTRAINT*
C 5  CONSTRAIN THE GEODETIC LATITUDE, LONGITUDE AND HEIGHT OF A STATION.*

C           ABSOLUTE CONSTRAINTS
C
C 11  DEFINE THE ORIGIN OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 12  DEFINE THE ORIENTATION OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT
C 13  DEFINE THE SCALE OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 14  COMPLETELY FIX ONE OR MORE COORDINATES OF A STATION * $
C 15  COMPLETELY FIX ONE OR MORE COORDINATES OF RELATIVE POSITION**

C
C *IF THE COORDINATES, RELATIVE POSITION, DISTANCE, OR DIRECTION, TO BE
C CONSTRAINED ARE NOT GIVEN, THE CONSTRAINT IS COMPUTED FROM THE
C APPROXIMATE COORDINATES OF THE STATION(S) INVOLVED
C
C
C $THE DIAGONAL ELEMENTS OF THE W MATRIX ARE USED AS CODES TO INDICATE
C WHICH COORDINATES ARE TO BE FIXED. A NON-ZERO CODE MEANS TO FIX
C THE COORDINATE.

C
C SUBROUTINE PSOLN
C IMPLICIT REAL*8(A-H,O-Z)
C COMMON/NSTA/NSTA/STAORD/KORDER(150)
C INTEGER*2 L, LSOLVE, IDS
C INTEGER STANAM
C COMMON/STALOC/STAUVW(3,150), DATPRM(2,15), DATNAM(4,15),
C 1STANAM(5,150), IDS(150)
C COMMON/WPW/WPW, XPU, IDEGF, IFSTA
C INTEGER*2 PCODE
C COMMON/PCODES/PCODE(20)
C DIMENSION ADX(3)
C DIMENSION EIG(6), EV(3,3), COVX(3,3), VARX(3,40)
C REAL*8 LAM
C DIMENSION Q(3,3,40), DELCOV(3,3), UNC(3)
C DIMENSION DX(3,40), UNCE(3,40)
C DIMENSION RLX(3,3), RLD(3,3), RXD(3,3)
C EQUIVALENCE(RXD(1,1), EV(1,1))
C LOC(K)=(K*(K+1))/2
C REWIND 2
C WRITE(6,6001) WPW, XPU
6001 FORMAT(/10X, "W**PW=", D16.8, " -X**U=", D16.8)
C VPV=WPW-XPU
C IDEGF=IDEgf-3*NSTA
C VARO=VPV/DFLOAT(IDEgf)
C SIGO=DSQRT(VARO)
C WRITE(6,6002) IDEGF, VPV, VARO, SIGO
6002 FORMAT(1H1/////////T50, "NUMBER OF DEGREES OF FREEDOM =", I8/1H0, T38
1, "QUADRATIC SUM OF ALL THE RESIDUALS (VPV) =", F13.4/1H0, T55,
2"VARIANCE OF UNIT WEIGHT =", F13.4/1H0, T45,
3"STANDARD DEVIATION OF UNIT WEIGHT =", F13.4)

```

C

```

READ(2) ((VARX(I,ISTA),I=1,3),ISTA=1,NSTA)
DO 80 ISTA=1,NSTA
DO 80 I=1,3
IF(VARX(I,ISTA).LE.0.0) GO TO 79
VARX(I,ISTA)=DSQRT(VARX(I,ISTA))
GO TO 80
79 VARX(I,ISTA)=0.0
80 CONTINUE
C
DO 200 ISTA=1,NSTA
READ(2)(DX(I,ISTA),I=1,3),(((Q(I,J,JSTA),I=1,3),J=1,3),
1JSTA=ISTA,NSTA)
C
IF(ISTA-2*(ISTA/2).EQ.1.OR.PCODE(20).EQ.1) WRITE(6,6011)
6011 FORMAT(1H1)
IDAT=IDS(ISTA)
WRITE(6,6003) KORDER(ISTA),(STANAM(I,ISTA),I=1,5)
1, IDAT,(DATNAM(I, IDAT),I=1,4)
6003 FORMAT(3(/),*0STATION NUMBER - ',I8,10X,4A4,A2,5X,
1'ELLIPSOID',I4,4X,4A8)
WRITE(6,6010)
6010 FORMAT('0',25X,'X',15X,'Y',15X,'Z',21X,'LAT.',12X,'LONG.(+E)
1ELL. HT.')
CALL UVWTG(STAUWV(1,ISTA),DATPRM(1, IDAT),PHI,LAM,H)
CALL DANG (PHI,ISGNP,IDEGL,IMINP,SECP)
CALL DANG (LAM,ISGNL,IDEGL,IMINL,SECL)
WRITE(6,6005) (STAUWV(I,ISTA),I=1,3),ISGNP,IDEGL,IMINP,SECP,
1ISGNL,IDEGL,IMINL,SECL,H
6005 FORMAT('OPREL. COORD. - ',3F16.4,7X,2(3X,A1,2I3,F8.4),F12.4)
DO 100 IK=1,3
100 ADX(IK)=DX(IK,ISTA)
CALL DELL(ADX
1 DATPRM(1, IDAT),DP,DL,DH,DELCOV,RLX),
, Q(1,1,ISTA),PHI,LAM,H,
DO 120 I=1,3
DO 110 J=1,3
DELCOV(I,J)=DELCOV(I,J)*VARO
110 COVX(I,J)=Q(I,J,ISTA)*VARO
UNC(I)=VARX(I,ISTA)*SIGO
IF(DELCOV(I,I).GT.0.0) GO TO 115
UNCE(I,ISTA)=0.0
GO TO 116
115 CONTINUE
UNCE(I,ISTA)=DSQRT(DELCOV(I,I))
116 CONTINUE
STAUWV(I,ISTA)=STAUWV(I,ISTA)+DX(I,ISTA)
120 CONTINUE
C
CALL DANG(DP,ISGNP,IDEGL,IMINP,SECP)
CALL DANG(DL,ISGNL,IDEGL,IMINL,SECL)
DO 125 I=1,3
IF(I.LT.3) UNCE(I,ISTA)=UNCE(I,ISTA)*206264.8062
DO 125 J=1,3
IF(I.LT.3) DELCOV(I,J)=DELCOV(I,J)*206264.8062
IF(J.LT.3) DELCOV(I,J)=DELCOV(I,J)*206264.8062
125 CONTINUE
WRITE(6,6006) (DX(I,ISTA),I=1,3),
1,ISGNP,IDEGL,IMINP,SECP,ISGNL,IDEGL,IMINL,SECL,DH
6006 FORMAT('OCORRECTIONS - ',3F16.4,7X,2(3X,A1,2I3,F8.4),F12.4)

```

```

CALL UVWTG(STAUVW(1,ISTA),DATPRM(1,1DAT),PHI,LAM,H)
CALL DANG(PHI,ISGNP,IDEGL,IMINP,SECP)
CALL DANG(LAM,ISGNL,IDEGL,IMINL,SECL)
WRITE(6,6007) (STAUVW(I,ISTA),I=1,3)
1,ISGNP,IDEGL,IMINP,SECP,ISGNL,IDEGL,IMINL,SECL,H
6007 FORMAT('ADJ. COORD. - ',3F16.4,7X,2(3X,A1,2I3,F8.4),F12.4)
  IF(PCODE(17).EQ.1) WRITE(7,7001) KORDER(ISTA),(STAUVW(I,ISTA),
  1  I=1,3),(COVX(1,I),I=1,3)
7001 FORMAT(14,4X,3F16.6/3F10.3)
  IF(PCODE(18).EQ.1) WRITE(7,5005) KORDER(ISTA),1DAT,
  1(STANAM(I,ISTA),I=1,5),ISGNP,IDEGL,IMINP,SECP,IDEGL,IMINL,SECL,H,
  2(UNCE(I,ISTA),I=1,3)
5005 FORMAT(14,I2,4A4,A2,A1,2(2I3,F8.4), F10.2,2F3.1,F3.0,7X,A1)
  WRITE(6,6008) ((COVX(I,J),J=1,3),(DELCOV(I,J),J=1,3),I=1,3)
6008 FORMAT('COVARIANCE-COVARIANCE MATRIX OF THE STATION POSITION//'
  13(14X,3F16.6, 10X,3F16.6))
  WRITE(6,6009) UNCE,(UNCE(I,ISTA),I=1,3)
6009 FORMAT('STANDARD. DEV. -',3F16.4,10X,3F16.4)
C
C      IF(PCODE(19).NE.1) GO TO 150
C          COMPUTE EIGENVALUES
  WRITE(6,6100)
6100 FORMAT('DIRECTIONS OF EIGENVECTORS AND SQUARE ROOTS OF EIGENVALUE
  1S OF VARIANCE-COVARIANCE MATRIX -',T20,'LATITUDE',T40,'LONGITUDE',
  2 T60,'ELEVATION',T80,'AZIMUTH',T100,'AXIS LENGTH')
  NB=0
  DO 135 J=1,3
  DO 135 I=1,J
  NB=NB+1
135 EIG(NB)=COVX(1,J)
  CALL DEIGEN(EIG,EV,3,0)
  CALL DGMPRD(RLX,RXD,RLD,3,3,3)
  DO 140 I=1,3
  PHI=DATAN(EV(3,I)/DSQRT(EV(1,I)**2+EV(2,I)**2))
  LAM=DATAN2(EV(2,I),EV(1,I))
  ELEV=DATAN(RLD(3,I)/DSQRT(RLD(1,I)**2+RLD(2,I)**2))
  AZ=DATAN2(RLD(2,I),RLD(1,I))
  CALL DANG(PHI,ISGNP,IDEGL,IMINP,SECP)
  CALL DANG(LAM,ISGNL,IDEGL,IMINL,SECL)
  CALL DANG(ELEV,ISGNEL,IDEGL,IMINEL,SECL)
  CALL DANG(AZ,ISGNAZ,IDEGL,IMINAZ,SECAZ)
  EVAL=EIG(LOC(1))
  IF(EVAL.LE.0.0) GO TO 137
  EVAL=DSQRT(EVAL)
  GO TO 139
137 EVAL=0.0
139 CONTINUE
  WRITE(6,6101) ISGNP,IDEGL,IMINP,SECP,ISGNL,IDEGL,IMINL,SECL,
  1 ISGNEL,IDEGL,IMINEL,SECL,ISGNAZ,IDEGL,IMINAZ,SECAZ,EVAL
6101 FORMAT(1H0,14X,4(A1,2I3,F8.4,5X),F12.4)
  140 CONTINUE
  150 CONTINUE
C
C      IF(PCODE(20).NE.1) GO TO 200
C          COMPUTE CORRELATION COEFFICIENTS
  WRITE(6,6105)
6105 FORMAT(1H0,15X,'3X3 WEIGHT COEFFICIENT MATRICES',43X,'CORRELATION
  1COEFFICIENTS'/1H0,2(25X,'X',15X,'Y',15X,'Z'))

```

```

DO 160 JSTA=ISTA,NSTA
DO 155 I=1,3
DO 155 J=1,3
DENOM=VARX(I,ISTA)*VARX(J,JSTA)
IF(DENOM) 154,154,153
153 EV(I,J)=Q(I,J,JSTA)/DENOM
GO TO 155
154 EV(I,J)=0.0
155 CONTINUE
      WRITE(6,6106) KORDER(ISTA),KORDER(JSTA)
6106 FORMAT(1H0,5X,' STA. NO.',15,' WITH STA. NO.',15)
      WRITE(6,6107) ((Q(I,J,JSTA),J=1,3),(EV(I,J),J=1,3),I=1,3)
6107 FORMAT(14X,3F16.6,10X,3F16.6)
160 CONTINUE
200 CONTINUE
C      PRINT OUT SUMMERY OF RESULTS
IF(PCODE(10).EQ.0) GO TO 300
      WRITE(6,6011)
DO 300 ISTA=1,NSTA
DO 219 I=1,3
219 UNC(I)=VARX(I,ISTA)*SIGO
      WRITE(6,6108) KORDER(ISTA),(STANAM(I,ISTA),I=1,5)
IDAT=IDS(ISTA)
JD=PCODE(10)
GO TO(220,240,260,260),JD
220 WRITE(6,6109) (DX(I,ISTA),I=1,3),UNC
GO TO 300
240 WRITE(6,6110) (STAUVW(I,ISTA),I=1,3),UNC
GO TO 300
260 CALL UVWTG(STAUVW(1,ISTA),DATPRM(1,1DAT),PHI,LAM,H)
CALL DANG (PHI,ISGNP,IDEGP,IMINP,SECP)
CALL DANG (LAM,ISGNL,IDEGL,IMINL,SECL)
IF (PCODE(10).EQ.4) GO TO 270
      WRITE(6,6111) ISGNP,IDEGP,IMINP,SECP,ISGNL,IDEGL,IMINL,SECL,H,
1(UNCE(I,ISTA),I=1,3)
GO TO 300
270 WRITE(6,6112) (STAUVW(I,ISTA),I=1,3),ISGNP,IDEGP,IMINP,SECP,IDEGL,
1IMINL,SECL,H,UNC,(UNCE(I,ISTA),I=1,3)
6108 FORMAT(1B,1X,4A4,A2)
6109 FORMAT (1H+,27X,3F10.4,/28X,3F10.4/)
6110 FORMAT (1H+,27X,3F16.4,/28X,3F16.4/)
6111 FORMAT (1H+,27X,2(3X,A1,213,F8.4),F12.4,/38X,F8.4,10X,F8.4,F12.4/)
6112 FORMAT(1H+,27X,3F16.4, 3X,A1,213,F8.4,3X,213,F8.4,F12.4,/28X,
1 3F16.4,2(10X,F8.4),F12.4/)
300 CONTINUE
      RETURN
      END

```

```

SUBROUTINE CONAP1(KODE)
C
C           CONSTRAINT CODE DIRECTORY
C
C           WEIGHTED CONSTRAINTS
C
C 1  CONSTRAIN THE COORDINATES OF A STATION AT A PRIORI VALUES*(1.E.WEIGHT IT)
C 2  IMPOSE CHORD DISTANCE CONSTRAINT*.
C 3  IMPOSE RELATIVE POSITION CONSTRAINT*
C 4  IMPOSE DIRECTION CONSTRAINT*
C 5  CONSTRAIN THE GEODETIC LATITUDE, LONGITUDE AND HEIGHT OF A STATION.*

C
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 L,LSOLVE,IDS
INTEGER CONTIN,STANAM
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1STANAM(5,150),IDS(150)
COMMON/NSTA/NSTA,NBLOCK
COMMON/STAORD/KORDER(150)
COMMON/NORMEQ/REDN(3,3,2485),U(3,70),L(2485),LSOLVE
COMMON/WPW/WPW,XPU,IDEGF,IFSTA
DIMENSION XI(3),XJ(3),W(3,3),DIS(3),DXB(3),DXC(3)
EQUIVALENCE (XI(1),DXB(1)),(XJ(1),DXC(1))
DIMENSION G(2,3)
DATA SPR,DPR/206264.80625,57.295779513/
LOC(K)=(K*(K+1))/2
6150 FORMAT(10X,3D16.8)
GO TO (100,200,300,400,500,600,700,800,900),KODE
100 CONTINUE
READ (3) IS,ISTA,X1,W
NB=LOC(ISTA)
K=0
DO 110 I=1,3
DIS(I)=XI(I)-STAUVW(I,ISTA)
REDN(I,I,NB)=REDN(I,I,NB)+W(I,I)
U(I,ISTA)=U(I,ISTA)+W(I,I)*DIS(I)
IF(W(I,I).EQ.0.0) GO TO 110
K=K+1
WPW=WPW+DIS(I)*W(I,I)*DIS(I)
IDEGF=IDEGF+1
110 CONTINUE
IF(K.EQ.3) L(NB)=L(NB)+1
WRITE(6,6100) IS,(STANAM(I,ISTA),I=1,5),XI,(W(I,I),I=1,3)
6100 FORMAT(//15X,'A PRIORI CONSTRAINT ON STATION',I5,2X,5A4/
115X,'COORDINATES',3F16.2/15X,'WEIGHTS',6X,3F16.4)
GO TO 10
C           CHORD CONSTRAINT
200 CONTINUE
READ (3) IS,ISTA,JS,JSTA,CD,RELUNC
IF(JSTA.GE.ISTA) GO TO 210
C           SWITCH SUBSCRIPTS IF NECESSARY
NB=ISTA
ISTA=JSTA
JSTA=NB
210 CONTINUE
CDC =0.0
DO 205 I=1,3
DIS(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)

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205 CDC=CDC+DIS(I)**2
CDC=DSQRT(CDC)
DO 215 I=1,3
215 DIS(I)=DIS(I)/CDC
CDD=CD-CDC
WCD=(RELUNC/CD)**2
IB=LOC(ISTA)
JB=LOC(JSTA)
NB=LOC(JSTA-1)+ISTA
DO 220 I=1,3
U(I,ISTA)=U(I,ISTA)+DIS(I)*WCD*CDD
U(I,JSTA)=U(I,JSTA)-DIS(I)*WCD*CDD
DO 220 J=1,3
TERM=DIS(I)*WCD*DIS(J)
REDN(I,J,IB)=REDN(I,J,IB)+TERM
REDN(I,J,JB)=REDN(I,J,JB)+TERM
REDN(I,J,NB)=REDN(I,J,NB)-TERM
220 CONTINUE
WRITE(6,6200) KORDER(ISTA),(STANAM(I,ISTA),I=1,5),KORDER(JSTA),
1(STANAM(I,JSTA),I=1,5),CD,RELUNC
6200 FORMAT(///15X,'CHORD DISTANCE CONSTRAINT IMPOSED BETWEEN STATION'
1,15,2X,5A4/53X,'AND STATION',I5,2X,5A4/15X,'CONSTRAINED DISTANCE='
2,F16.2/15X,'THE WEIGHT IS COMPUTED FROM A RELATIVE UNCERTAINTY OF'
3'ONE PART IN',F16.2)
WRITE(6,6150) CDD
L(IB)=L(IB)+1
L(JB)=L(JB)+1
L(NB)=L(NB)+1
WPW=WPW+ CDD*WCD*CDD
IDEGF=IDEGF+1
GO TO 10
C
      RELATIVE POSITION CONSTRAINT
300 CONTINUE
READ (3) IS,ISTA,JS,JSTA,DXB,W
IB=LOC(ISTA)
JB=LOC(JSTA)
NB=LOC(JSTA-1)+ISTA
IF(ISTA.GT.JSTA) NB=LOC(ISTA-1)+JSTA
DO 310 I=1,3
DXC(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)
DIS(I)=DXB(I)-DXC(I)
IF(W(I,I).EQ.0.0) GO TO 310
WPW=WPW+DIS(I)*W(I,I)*DIS(I)
IDEGF=IDEGF+1
U(I,ISTA)=U(I,ISTA)+W(I,I)*DIS(I)
U(I,JSTA)=U(I,JSTA)-W(I,I)*DIS(I)
REDN(I,I,IB)=REDN(I,I,IB)+W(I,I)
REDN(I,I,JB)=REDN(I,I,JB)+W(I,I)
REDN(I,I,NB)=REDN(I,I,NB)-W(I,I)
310 CONTINUE
L(IB)=L(IB)+1
L(JB)=L(JB)+1
L(NB)=L(NB)+1
WRITE(6,6300) KORDER(ISTA),(STANAM(I,ISTA),I=1,5),KORDER(JSTA),
1(STANAM(I,JSTA),I=1,5),DXB,(W(I,I),I=1,3)
6300 FORMAT(///15X,'RELATIVE POSITION CONSTRAINT'/15X,'BETWEEN STATION'
1',I5,3X,5A4,' AND STATION',I5,2X,5A4//15X,'RELATIVE COORDINATES
2 ARE'/15X,3F16.2//15X,'WEIGHTS ARE'/17X,3F16.4)

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      WRITE(6,6150) (DIS(I),I=1,3)
      GO TO 10
 400 CONTINUE
C      DIRECTION CONSTRAINTS
C      ALPHA IS LONGITUDE-LIKE ANGLE
C      BETA IS LATITUDE-LIKE ANGLE
      READ (3) IS,ISTA,JS,JSTA,ALF,BETA,VARA,VARB,COVAB
      WRITE(6,6400) KORDER(ISTA),(STANAM(I,ISTA),I=1,5),KORDER(JSTA),
      1(STANAM(I,JSTA),I=1,5),ALF,BETA,VARA,VARB,COVAB
 6400 FORMAT(///15X,'DIRECTION CONSTRAINT IMPOSED BETWEEN STATION',
      1I5,2X,5A4,/4X,'AND STATION',I5,2X,5A4/15X,'ANGLES(DEGREES)',4X,
      22F16.8/15X,'UNCERTAINTIES(SECONDS)',3F16.3)
      DO 405 I=1,3
 405 DXC(I)=STAUVW(I,ISTA)-STAUVW(I,JSTA)
      RSCSB=DXC(1)**2+DXC(2)**2
      TA=DXC(2)/DXC(1)
      CSA=1.0/(1.0+TA*TA)
      RCB=DSQRT(RSCSB)
      TB=DXC(3)/RCB
      CSB=1.0/(1.0+TB*TB)

C      AO=DATAN2(DXC(2),DXC(1))
      DIS(1)=ALF/DPR-AO
      PI=180.0/DPR
      IF(DIS(1).GT.PI) DIS(1)=DIS(1)-2.0*PI
      IF(DIS(1).LT.(-PI)) DIS(1)=DIS(1)+2.0*PI
      BO=DATAN(DXC(3)/RCB)
      DIS(2)=BETA/DPR-BO
      WRITE(6,6150) DIS(1),DIS(2)

C      G(1,1)=CSA*TA/DXC(1)
      G(1,2)=-CSA/DXC(1)
      G(1,3)=0.0
      G(2,1)=CSB*TB*DXC(1)/RSCSB
      G(2,2)=G(2,1)*TA
      G(2,3)=-CSB/RCB

C      VARA=(VARA/SPR)**2
      VARB=(VARB/SPR)**2
      DET=VARA*VARB-COVAB*COVAB
      W(1,1)=VARB/DET
      W(2,2)=VARA/DET
      W(1,2)=-COVAB/DET
      W(2,1)=W(1,2)

C      1B=LOC(ISTA)
      JB=LOC(JSTA)
      NB=LOC(JSTA-1)+ISTA
      IF(ISTA.GT.JSTA) NB=LOC(ISTA-1)+JSTA
      DO 445 I=1,3
      SUM=0.0
      DO 443 II=1,2
      DO 443 JJ=1,2
 443 SUM=SUM+G(II,I)*W(II,JJ)*DIS(JJ)
      U(I,ISTA)=U(I,ISTA)-SUM
      U(1,JSTA)=U(1,JSTA)+SUM
      DO 445 J=1,3
      SUM=0.0

```

```

DO 444 II=1,2
DO 444 JJ=1,2
444 SUM=SUM+G(II,I)*W(II,JJ)*G(JJ,J)
REDN(I,J,IB)=REDN(I,J,IB)+SUM
REDN(I,J,JB)=REDN(I,J,JB)+SUM
IF(ISTA.GT.JSTA) GO TO 446
REDN(I,J,NB)=REDN(I,J,NB)-SUM
GO TO 445
446 REDN(J,I,NB)=REDN(J,I,NB)-SUM
445 CONTINUE
DO 450 II=1,2
IF(W(II,II).EQ.0.0) GO TO 450
L(IB)=L(IB)+1
L(JB)=L(JB)+1
L(NB)=L(NB)+1
IDEGF=IDEGF+1
DO 450 JJ=1,2
WPH=WPH+DIS(II)*W(II,JJ)*DIS(JJ)
450 CONTINUE
GO TO 10
C
500 CONTINUE
READ (3) IS,ISTA,IDTS,PHIO,FLAMO,HO,SDP,SDL,SDH
WRITE(6,6500) IS,(STANAM(I,ISTA),I=1,5),PHIO,FLAMO,HO,IDTS,
I (DATNAM(I,IDTS),I=1,4), SDP,SDL,SDH
6500 FORMAT(//15X,'THE ELLIPSOIDAL COORDINATES (LAT.,LONG.,HEIGHT) OF
1 STATION',I7,3X,5A4/15X,'ARE CONSTRAINED AT'//2(F20.9,' DEGREES'),
2 F20.3,' METERS'//15X,'ON DATUM',I5,3X,4A8//'
3 15X,'THE WEIGHTS FOR THESE CONSTRAINTS ARE COMPUTED FROM OBSERVAT
4IONAL STANDARD DEVIATIONS OF'//
5 2(F20.3,' SECONDS'),F20.3,' METERS')
CALL UVWTG3(STAUUV(1,ISTA),DATPRM(1,IDTS),PHI,FLAM,H)
IB=LOC(ISTA)
SP=DSIN(PHI)
CP=DCOS(PHI)
SL=DSIN(FLAM)
CL=DCOS(FLAM)
AE=DATPRM(1, IDTS)
E2=1.0-(DATPRM(2, IDTS)/AE)**2
EW=DSQRT(1.0-E2*SP*SP)
EN=AE/EW
EM=AE*(1.0-E2)/EW**3
IF(SDP.EQ.0.0) GO TO 510
WT=1.0/(SDP/SPR)**2
DXB(1)=-SP*CL/(EM+H)
DXB(2)=-SP*SL/(EM+H)
DXB(3)= CP/(EM+H)
DISC=PHIO/DPR-PHI
ASSIGN 510 TO J5
GO TO 550
510 CONTINUE
IF(SDL.EQ.0.0) GO TO 520
WT=1.0/(SDL/SPR)**2
DENOM=(EN+H)*CP
DXB(1)=-SL/DENOM
DXB(2)= CL/DENOM
DXB(3)=0.0
DISC=FLAMO/DPR-FLAM

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ASSIGN 520 TO J5
GO TO 550
520 CONTINUE
IF(SDH.EQ.0.0) GO TO 560
WT=1.0/SDH**2
DXB(1)=CP*CL
DXB(2)=CP*SL
DXB(3)=SP
DISC=HQ-H
WRITE(6,6150) DISC
ASSIGN 560 TO J5
GO TO 550
550 CONTINUE
DO 555 I=1,3
U(I,ISTA)=U(I,ISTA)+DXB(I)*WT*DISC
DO 555 J=1,3
REDN(I,J,IB)=REDN(I,J,IB)+DXB(I)*WT*DXB(J)
555 CONTINUE
L(IB)=L(IB)+1
IDEGF=IDEGF+1
WPW=WPW+DISC*WT*DISC
GO TO J5,(510,520,560)
560 CONTINUE
GO TO 10
C
600 CONTINUE
700 CONTINUE
800 CONTINUE
900 CONTINUE
10 CONTINUE
RETURN
END
```

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SUBROUTINE CONAP2(KODE2)
      ABSOLUTE CONSTRAINTS
C
C 11  DEFINE THE ORIGIN OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 12  DEFINE THE ORIENTATION OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT
C 13  DEFINE THE SCALE OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
C 14  COMPLETELY FIX ONE OR MORE COORDINATES OF A STATION *
C 15  COMPLETELY FIX ONE OR MORE COORDINATES OF RELATIVE POSITION**
C
C *IF THE COORDINATES, RELATIVE POSITION, DISTANCE, OR DIRECTION, TO BE
C CONSTRAINED ARE NOT GIVEN, THE CONSTRAINT IS COMPUTED FROM THE
C APPROXIMATE COORDINATES OF THE STATION(S) INVOLVED
C
C $THE DIAGONAL ELEMENTS OF THE W MATRIX ARE USED AS CODES TO INDICATE
C WHICH COORDINATES ARE TO BE FIXED. A NON-ZERO CODE MEANS TO FIX
C THE COORDINATE.
C
C PROCESS A PRIORI CONSTRAINTS ON, AND BETWEEN, STATIONS
      IMPLICIT REAL*8(A-H,O-Z)
      INTEGER*2 L,LSOLVE,IDS
      INTEGER ENDSIG/1HE/,CONTIN,STANAM
      COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
      1STANAM(5,150),IDS(150)
      COMMON/NSTA/NSTA,NBLOCK
      COMMON/STAORD/KORDER(150)
      COMMON/NORMEQ/REDN(3,3,2485),U(3,70),L(2485),LSOLVE
      COMMON/WPW/WPW,XPU,IDEKF,IFSTA
      DIMENSION XI(3),XJ(3),W(3,3),DIS(3),DXB(3),DXC(3)
      EQUIVALENCE (XI,DXB),(XJ,DXC)
      DIMENSION G(2,3)
      DATA SPR/206264.80625/
      LOC(K)=(K*(K+1))/2
      MAXBLK=70
C
C ABSOLUTE CONSTRAINTS THAT REQUIRE EXPANSION OF THE NORMAL EQUATION
C MATRIX BY THE ADDITION OF LAGRANGE MULTIPLIERS
C
C KODE2 IS KODE-10
      GO TO (1100,1200,1300,1400,1500,1600,1700,1800,1900),KODE2
C 11  DEFINE THE ORIGIN OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
1100 CONTINUE
      ASSIGN 1110 TO JRTN
      GO TO 960
1110 CONTINUE
      DO 1120 ISTA=1,NSTA
      IB=LOC(NBLOCK-1)+ISTA
      L(IB)=1
      DO 1120 I=1,3
1120 REDN(I,I,IB)=1.0
      IDEGF=IDEKF+3
      WRITE(6,6011)
6011 FORMAT('THE ORIGIN OF THE COORDINATE SYSTEM IS DEFINED BY INNER
      ADJUSTMENT.')
      GO TO 10
C 12  DEFINE THE ORIENTATION OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT
1200 CONTINUE
      ASSIGN 1210 TO JRTN
      GO TO 960

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1210 CONTINUE
  DO 1220 ISTA=1,NSTA
    IB=LOC(NBLOCK-1)+ISTA
    REDN(1,2,IB)= STAUUVW(3,ISTA)/SPR
    REDN(1,3,IB)=-STAUUVW(2,ISTA)/SPR
    REDN(2,1,IB)=-STAUUVW(3,ISTA)/SPR
    REDN(2,3,IB)= STAUUVW(1,ISTA)/SPR
    REDN(3,1,IB)= STAUUVW(2,ISTA)/SPR
    REDN(3,2,IB)=-STAUUVW(1,ISTA)/SPR
    L(IB)=1
1220 CONTINUE
  IDEGF=IDEgf+3
  WRITE(6,6012)
6012 FORMAT('ORIENTATION OF THE COORDINATE SYSTEM DEFINED BY INNER
1ADJUSTMENT PROCEDURE')
  GO TO 10
C 13  DEFINE THE SCALE OF THE COORDINATE SYSTEM BY INNER ADJUSTMENT EQUATIONS
1300 CONTINUE
  ASSIGN 1310 TO JRTN
  GO TO 960
1310 CONTINUE
  DO 1320 ISTA=1,NSTA
    IB=LOC(NBLOCK-1)+ISTA
    L(IB)=1
    DO 1320 I=1,3
      REDN(I,1,IB)=STAUUVW(I,ISTA)/1000000.0
1320 CONTINUE
C  FILL IN EXTRA TWO ROWS IN BLOCK WITH DUMMY EQUATIONS
  NB=LOC(NBLOCK)
  REDN(2,2,NB)=1.0
  REDN(3,3,NB)=1.0
  L(NB)=1
  IDEGF=IDEgf+1
  WRITE(6,6013)
6013 FORMAT('SCALE OF THE COORDINATE SYSTEM DEFINED BY INNER ADJUSTMENT
1T PROCEEDURE.')
  GO TO 10
C 14  COMPLETELY FIX ONE OR MORE COORDINATES OF A STATION
1400 CONTINUE
  ASSIGN 1410 TO JRTN
  GO TO 960
1410 CONTINUE
  READ(3) IS,ISTA,XI,W
  DO 1420 I=1,3
1420 DIS(I)=XI(I)-STAUUVW(I,ISTA)
  IB=LOC(NBLOCK-1)+ISTA
  NB=LOC(NBLOCK)
  DO 1430 I=1,3
C  THE DIAGONAL ELEMENTS OF W ARE USED AS INDICATORS TO SHOW
C  WHICH COORDINATES ARE TO BE FIXED.
  IF(W(I,I).EQ.0.0) GO TO 1424
  REDN(I,1,IB)=1.0
  U(I,NBLOCK)=DIS(I)
  IDEGF=IDEgf+1
  L(IB)=1
  GO TO 1430
1424 REDN(I,1,NB)=1.0
  L(NB)=1

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```

1430 CONTINUE
  WRITE(6,6014) IS,(STANAM(I,ISTA),I=1,5),XI,(W(I,I),I=1,3)
6014 FORMAT(///15X,'CARTESIAN COORDINATES OF STATION',I5,2X,5A4,3X,
  1'FIXED AT'//15X,3F16.2//15X,'FIXED COORDINATES ARE INDICATED BY
  2NON ZERO ENTRY BELOW'//15X,3F16.2)
  L(LOC(ISTA))=1
  GO TO 10
C
1500 CONTINUE
C 15 COMPLETELY FIX ONE OR MORE COORDINATES OF RELATIVE POSITION
  ASSIGN 1510 TO JRTN
  GO TO 960
1510 CONTINUE
  READ (3) IS,ISTA,JS,JSTA,DXB,W
  IB=LOC(NBLOCK-1)+ISTA
  JB=LOC(NBLOCK-1)+JSTA
  NB=LOC(NBLOCK)
  L(IB)=1
  L(JB)=1
  L(LOC(ISTA))=1
  L(LOC(JSTA))=1
  DO 1530 I=1,3
  IF(W(I,I).EQ.0.0) GO TO 1524
  REDN(I,I,IB)=1.0
  REDN(I,I,JB)=-1.0
  U(I,NBLOCK)=DXB(I)-(STAUVW(I,ISTA)-STAUVW(I,JSTA))
  IDEGF=IDEGF+1
  GO TO 1530
1524 REDN(I,I,NB)=1.0
  L(NB)=1
1530 CONTINUE
  WRITE(6,6015) IS,(STANAM(I,ISTA),I=1,5),JS,(STANAM(I,JSTA),I=1,5),
  1 DXB,(W(I,I),I=1,3)
6015 FORMAT(///15X,'RELATIVE POSITION BETWEEN STATION',I5,3X,5A4/
  137X,'AND STATION',I5,3X,5A4/           15X,
  2'FIXED AT'//15X,3F16.2//15X,'RELATIVE COORDINATES WHICH ARE FIXED
  3ARE INDICATED BY A NON-ZERO ENTRY BELOW'//15X,3F16.2)
  GO TO 10
1600 CONTINUE
C  SET UP A BLOCK OF 3 DUMMY EQUATIONS
  ASSIGN 1610 TO JRTN
  GO TO 960
1610 CONTINUE
  NB=LOC(NBLOCK)
  DO 1615 I=1,3
1615 REDN(I,I,NB)=1.0
  L(NB)=1
  WRITE(6,6016)
6016 FORMAT('OBLOCK OF 3 DUMMY EQUATIONS ADDED TO REDUCED NORMALS.')
  GO TO 10
C
C  EXPAND REDUCED NORMALS BY ADDING A BLOCK OF THREE LAGRANGE MULTIPIERS
960 CONTINUE
  NBLOCK=NBLOCK+1
  IF(NBLOCK.LE.MAXBLK) GO TO 961
  WRITE(6,6096)
6096 FORMAT('ATTEMPTED CONSTRAINT RESULTS IN AN ATTEMPT TO EXPAND THE
  1REDUCED NORMAL EQUATION MATRIX'//15X,'BEYOND ITS DIMENSIONS.'/>

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```
215X,'PROGRAM STOPS')
STOP
961 CONTINUE
  DO 965 ISTA=1,NBLOCK
  NB=LOC(NBLOCK-1)+ISTA
  L(NB)=0
  DO 965 I=1,3
  DO 965 J=1,3
965 REDN(I,J,NB)=0.0
  DO 966 I=1,3
966 U(I,NBLOCK)=0.0
  GO TO JRTN,(1110,1210,1310,1410,1510,1610)
C
C
1700 CONTINUE
1800 CONTINUE
1900 CONTINUE
10 CONTINUE
  RETURN
  END
```

```

SUBROUTINE CONAP
IMPLICIT REAL*8(A-H,O-Z)
C PROCESS A PRIORI CONSTRAINTS ON, AND BETWEEN, STATIONS
COMMON/NSTA/NSTA,NBLOCK
INTEGER*2 L,LSOLVE,IDS
INTEGER ENDSIG/1HE/,CONTIN,STANAM
COMMON/NORMEQ/REDN(3,3,2485),U(3,70),L(2485),LSOLVE
COMMON/WPW/WPW,XPU,IDEKF,IFSTA
LOC(K)=(K*(K+1))/2
IFSTA=0
NBLOCK=NSTA
10 CONTINUE
C WRITE(6,6801) NSTA,NBLOCK
6801 FORMAT(7I7)
READ (3) KODE,CONTIN
IF(CONTIN.EQ.ENDSIG) GO TO 1000
IF(KODE.LE.0) GO TO 950
IF(KODE.GT.19) GO TO 950
IF(KODE.GT.9) GO TO 11
GO TO (100,200,300,400,500,600,700,800,900),KODE
11 IF(KODE.LT.11) GO TO 950
KODE2=KODE-10
GO TO (1100,1200,1300,1400,1500,1600,1700,1800,1900),KODE2
C
C
100 CONTINUE
200 CONTINUE
300 CONTINUE
400 CONTINUE
500 CONTINUE
CALL CONAPI(KODE)
GO TO 10
600 CONTINUE
700 CONTINUE
800 CONTINUE
900 CONTINUE
GO TO 950
C
1100 CONTINUE
1200 CONTINUE
1300 CONTINUE
1400 CONTINUE
1500 CONTINUE
1600 CONTINUE
CALL CONAP2(KODE2)
GO TO 10
1700 CONTINUE
1800 CONTINUE
1900 CONTINUE
GO TO 950
C
950 WRITE(6,6095) KODE
6095 FORMAT('!ILLEGAL CONSTRAINT CODE IN CONAP IGNORED',15)
GO TO 10
C
CHECK TO SEE IF NORMALS ARE SOLVABLE
1000 CONTINUE
LSOLVE=1

```

```
DO 1010 ISTA=1,NSTA
NB=LOC(ISTA)
C      WRITE(6,6801) NB,L(NB)
      IF(L(NB).NE.0) GO TO 1010
      LSOLVE=0
      IDEGF=ISTA
1010 CONTINUE
C
      RETURN
      END
```

```

SUBROUTINE SOLVE
C SOLVE NORMAL EQUATIONS AND COMPUTE INVERSE FOR UP TO 40 STATIONS
C BY THE METHOD OF TRIANGULAR MATRICES.
C SEE DEPT. OF GEODETIC SCIENCE REPORT NO. 86, SECTION 5
C THE SCHEME USED TO ADDRESS THE UPPER TRIANGULAR PART OF THE REDUCED
C NORMALS IS THE SAME AS THAT USED IN FORMRN.
C
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*2 L,LSOLVE
COMMON/NORMEQ/REDN(3,3,2485),U(3,70),L(2485),LSOLVE
COMMON/WPW/WPH,XPU,IDEFG,NFSTA
COMMON/NSTA/NSTA1,NBLOCK
DIMENSION DX(3,70),TEMP(3,3,70),TA(3,3),TB(3,3),TC(3)
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
INTEGER*2 PIVOT(210)
DIMENSION RN(22365),RU(210),TEM(210),DXI(210)
EQUIVALENCE (RN,REDN),(RU,U,DXI),(TEM,TEMP)
EQUIVALENCE (DX,U)
C
LOC(K)=(K*(K+1))/2
C
IF(PCODE(16).LT.2) GO TO 5
DO 2 ISTA=1,NBLOCK
DO 1 JSTA=ISTA,NBLOCK
NB=LOC(JSTA-1)+ISTA
WRITE(6,6803) ISTA,JSTA,NB,L(NB)
6803 FORMAT(1H0,7I7)
WRITE(6,6801) ((REDN(I,J,NB),J=1,3),I=1,3)
6801 FORMAT(//3(3D20.8/))
1 CONTINUE
WRITE(6,6802) (U(I,ISTA),I=1,3)
6802 FORMAT(/3D20.8)
2 CONTINUE
5 CONTINUE
CHECK TO SEE IF THIS SET OF EQUATIONS HAS BEEN MARKED SOLVABLE
IF(LSOLVE.GE.1) GO TO 10
9 WRITE(6,6001) IDEFG
6001 FORMAT('REDUCED NORMALS MARKED UNSOLVABLE. PROGRAM STOPS.',15)
STOP
10 CONTINUE
REWIND 2
C NSTA1 GIVES THE NUMBER OF GROUND STATIONS IN THE ADJUSTMENT
C NBLOCK GIVES THE TOTAL NUMBER OF BLOCKS OF UNKNOWNS IN THE REDUCED-NORMALS,
C INCLUDING BOTH STATION COORDINATES AND BLOCKS OF LAGRANGE MULTIPLIERS.
C
C THE EXPANDED (NBLOCK SQUARE) SET OF REDUCED NORMALS IS SOLVED
NSTA=NBLOCK
REWIND 2
DO 20 ISTA=1,NBLOCK
DO 20 JSTA=ISTA,NBLOCK
NB=LOC(JSTA-1)+ISTA
19 WRITE(2) ((REDN(I,J,NB),I=1,3),J=1,3)
20 CONTINUE
REWIND 2
NUNK=3*NBLOCK
NB=LOC(NUNK)
DO 25 I=1,NB

```

```

25 RN(I)=0.0
DO 30 ISTA=1,NBLOCK
DO 30 JSTA=ISTA,NBLOCK
READ(2) TA
DO 30 I=1,3
II=3*(ISTA-1)+I
DO 30 J=1,3
JJ=3*(JSTA-1)+J
IF(II.GT.JJ) GO TO 30
IB=LOC(JJ-1)+II
RN(IB)=TA(I,J)
30 CONTINUE
C
IF(PCODE(16).LT.3) GO TO 41
DO 40 I=1,NUNK
DO 35 J=I,NUNK
NB=LOC(J-1)+I
TEM(J)=RN(NB)
35 CONTINUE
40 WRITE(6,6805) I,(TEM(J),J=I,NUNK)
6805 FORMAT(I5,6D19.10/250(5X,6D19.10/))
41 CONTINUE
C
C PERFORM FIRST REDUCTION - COMPUTE R AND C MATRICES -EQ.5-9
C
DO 100 I=1,NUNK
C
C FIND PIVOT ELEMENT
IP=0
PMAX=0.0
DO 55 J=I,NUNK
NB=LOC(J)
IF(DABS(RN(NB)).LE.PMAX) GO TO 55
PMAX=DABS(RN(NB))
IP=J
55 CONTINUE
PIVOT(I)=IP
IF(PCODE(16).GT.0) WRITE(6,6806) I,IP,PMAX
6806 FORMAT(2I7,D20.10)
IF(IP.EQ.0) GO TO 9
C
SWITCH ROWS AND COLUMNS
CALL SWITCH(I,IP)
C
IB=LOC(I)
RN(IB)=1.0/RN(IB)
IF(I.EQ.NUNK) GO TO 100
IP1=I+1
C
OUTER LOOP ---REDUCE ROW K
DO 80 K=IP1,NUNK
KIB=LOC(K-1)+I
C
GET MULTIPLIER
TD=RN(KIB)*RN(IB)
IF(TD.EQ.0.0)GO TO 80
C
REDUCE CONSTANT COLUMN
RU(K)=RU(K)-TD*RU(I)
C
INNER LOOP
DO 80 J=K,NUNK

```

```

NB=LOC(J-1)+I
KJB=LOC(J-1)+K
RN(KJB)=RN(KJB)-TD*RN(NB)
80 CONTINUE
100 CONTINUE
C
C
XPU=0.0
DO 300 NMIP1=1,NUNK
I=NUNK-NMIP1+1
NB=LOC(I)
C ACCUMULATE XPU
XPU=XPU+RN(NB)*RU(I)**2
IP1=I+1
C
C BACK SUBSTITUTION
IF(I.EQ.NUNK) GO TO 190
DO 150 J=IP1,NUNK
IB=LOC(J-1)+I
RU(I)=RU(I)-RN(IB)*DXI(J)
150 CONTINUE
190 CONTINUE
DXI(I)=RN(NB)*RU(I)
C
C
C DEVELOP ROW I OF INVERSE MATRIX
IF(I.EQ.NUNK) GO TO 300
DO 280 NMJP1=1,NMIP1
J=NUNK-NMJP1+1
TD=0.0
IF(I.EQ.J) TD=1.0
DO 270 K=IP1,NUNK
IB=LOC(K-1)+I
IF(J.EQ.I) GO TO 245
IF(K.GT.J) GO TO 240
JB=LOC(J-1)+K
TE=RN(JB)
GO TO 250
240 CONTINUE
JB=LOC(K-1)+J
TE=RN(JB)
GO TO 250
245 CONTINUE
TE=TEM(K)
250 CONTINUE
TD=TD-RN(IB)*TE
270 CONTINUE
C STORE ITH ROW TEMPORARILY IN TEM
TEM(J)=RN(NB)*TD
280 CONTINUE
C
COPY ITH ROW OUT OF TEM
DO 290 NMJP1=1,NMIP1
J=NUNK-NMJP1+1
JB=LOC(J-1)+I
RN(JB)=TEM(J)
290 CONTINUE

```

```

C UNDO PIVOTING
  IP=PIVOT(I)
  CALL SWITCH(I,IP)
300 CONTINUE
  REWIND 2
C
  DO 305 I=1,NUNK
  DO 304 J=I,NUNK
  IB=LOC(J-1)+I
304 TEM(J)=RN(IB)
  WRITE(2) (TEM(J),J=I,NUNK)
  IF(PCODE(16).LT.3) GO TO 305
  WRITE(6,6805) I,(TEM(J),J=I,NUNK),DXI(I)
305 CONTINUE
C
C
  REWIND 2
  DO 310 I=1,NUNK
  READ(2) (TEM(J),J=I,NUNK)
  DO 310 J=I,NUNK
  ISTA=(I-1)/3+1
  JSTA=(J-1)/3+1
  NB=LOC(JSTA-1)+ISTA
  II=I-3*(ISTA-1)
  JJ=J-3*(JSTA-1)
  REDN(II,JJ,NB)=TEM(J)
  IF(ISTA.EQ.JSTA) REDN(JJ,II,NB)=TEM(J)
310 CONTINUE
  REWIND 2
C
C OUTPUT THE SOLUTION AND COVARIANCE BLOCKS CORRESPONDING TO THE STATION UNKNOWN
  DO 320 ISTA=1,NSTA1
  NB=LOC(ISTA)
  DO 320 I=1,3
  TEMP(I,I,ISTA)=REDN(I,I,NB)
320 CONTINUE
  WRITE(2) ((TEMP(I,I,ISTA),I=1,3),ISTA=1,NSTA1)
  DO 350 ISTA=1,NSTA
  IF(ISTA.GT.NSTA1) GO TO 339
  DO 330 JSTA=ISTA,NSTA1
  NB=LOC(JSTA-1)+ISTA
  DO 330 I=1,3
  DO 330 J=1,3
  TEMP(I,J,JSTA)=REDN(I,J,NB)
330 CONTINUE
  WRITE(2) (DX(I,ISTA),I=1,3),(((TEMP(I,J,JSTA),I=1,3),J=1,3),
  1JSTA=ISTA,NSTA1)
339 CONTINUE
  IF(PCODE(16).LT.2) GO TO 350
  IB=LOC(ISTA)
  WRITE(6,6002) ISTA,(DX(I,ISTA),I=1,3),((REDN(I,J,IB),
  1 J=1,3),I=1,3)
6002 FORMAT(//I5/3F16.8//3(3D16.8/))
  DO 340 JSTA=ISTA,NSTA
  NB=LOC(JSTA-1)+ISTA
  340 WRITE(6,6003) JSTA,((REDN(I,J,NB
6003 FORMAT(//I5//3(3D16.8/))
  1 ,J=1,3),I=1,3)
350 CONTINUE

```

REWIND 2
RETURN
END

```

SUBROUTINE SWITCH(I,IP)
C   SWITCH ROW AND COLUMN I WITH ROW AND COLUMN IP
  IMPLICIT REAL*8(A-H,O-Z)
  INTEGER*2 L,LSOLVE
  COMMON/NORMEQ/REDN(3,3,2485),U(3,70),L(2485),LSOLVE
  COMMON/NSTA/NSTA1,NBLOCK
  DIMENSION RN(22365),RU(210)
  EQUIVALENCE (RN,REDN),(RU,U)
  LOC(K)=(K*(K+1))/2
  NSTA=NBLOCK
  NUNK=3*NBLOCK
  IF(IP.EQ.1) RETURN
  DO 25 J=1,NUNK
    NB=LOC(J-1)+I
    IF(J.EQ.I) GO TO 22
    IF(J-IP).EQ.16,25,18
16   IB=LOC(IP-1)+J
    GO TO 24
18   IB=LOC(J-1)+IP
    GO TO 24
22   IB=LOC(IP)
    GO TO 24
24   CONTINUE
    TD=RN(IB)
    RN(IB)=RN(NB)
    RN(NB)=TD
25   CONTINUE
    TD=RU(IP)
    RU(IP)=RU(I)
    RU(I)=TD
    RETURN
  END

```

```
DOUBLE PRECISION FUNCTION ANRADD(ISGN, IDEG, MIN, SEC)
INTEGER*2 MINUS/1H-/, PLUS/1H+/, AMPSAN/1H&/, ISGN, IDEG, MIN
DOUBLE PRECISION SEC
IF(IDEGL.GE.0) GO TO 10
ISGN=MINUS
IDEGL=-IDEGL
10 CONTINUE
ANRADD=(DFLOAT((IDEGL*60+MIN)*60)+SEC)/206264.80625
IF(ISGN.EQ.MINUS) ANRADD=-ANRADD
IF(ISGN.EQ.AMPSAN) ISGN=PLUS
RETURN
END
```

```
INTEGER FUNCTION KSTAID(ID)
COMMON/STAORD/KORDER(150)
COMMON/NSTA/NSTA
KSTAID=0
C SEARCH TABLE OF STATION IDENTIFIERS FOR THE INTERNAL NUMBER OF THIS STATION
DO 10 I=1,NSTA
  IF(KORDER(I).NE.ID) GO TO 10
  KSTAID=I
  RETURN
10 CONTINUE
/ RETURN
END
```

```
SUBROUTINE UVWTG3(UVW,DATUM,PHI,LAM,H)
C  CONVERT RECTANGULAR TO GEODETIC COORDINATES
C  ALIAS FOR UVWTG
  IMPLICIT REAL*8(A-Z)
  DIMENSION UVW(3),DATUM(2)
  LAM=DATAN2(UVW(2),UVW(1))
  IF(LAM.LT.0.0) LAM=LAM+6.28318530717958
  OME2=(DATUM(2)/DATUM(1))**2
  E2=1.0-OME2
  P=DSQRT(UVW(1)**2+UVW(2)**2)
  WP=UVW(3)/P
  TP1=WP/OME2
  PHI1=DATAN(TP1)
5   TTP=TP1*TP1
  SECP=DSQRT(1.0+TTP)
  N=DATUM(1)*SECP/DSQRT(1.0+OME2*TTP)
  H=P*SECP-N
  TP2=WP/(1.0-E2*N/(N+H))
  PHI=DATAN(TP2)
  IF(DABS(PHI-PHI1).LT.1.D-12) RETURN
  PHI1=PHI
  TP1=TP2
  GO TO 5
END
```

```
SUBROUTINE UVWTG2(UVW,DATUM,PHI,LAM,H)
C   CONVERT RECTANGULAR TO GEODETIC COORDINATES
C   ALIAS FOR UVWTG
      IMPLICIT REAL*8(A-Z)
      DIMENSION UVW(3),DATUM(2)
      LAM=DATAN2(UVW(2),UVW(1))
      IF(LAM.LT.0.0) LAM=LAM+6.28318530717958
      OME2=(DATUM(2)/DATUM(1))**2
      E2=1.0-OME2
      P=DSQRT(UVW(1)**2+UVW(2)**2)
      WP=UVW(3)/P
      TP1=WP/OME2
      PHI1=DATAN(TP1)
5   TTP=TP1*TP1
      SECP=DSQRT(1.0+TTP)
      N=DATUM(1)*SECP/DSQRT(1.0+OME2*TTP)
      H=P*SECP-N
      TP2=WP/(1.0-E2*N/(N+H))
      PHI=DATAN(TP2)
      IF(DABS(PHI-PHI1).LT.1.D-12) RETURN
      PHI1=PHI
      TP1=TP2
      GO TO 5
      END
```

```

SUBROUTINE DELL(DX,DXCOV, PHI,LAM,H,DATUM,DP,DL,DH,DELCOV,RLX)
IMPLICIT REAL*8(A-H,O-Z)
DIMENSION DX(3),DXCOV(3,3),DATUM(2),DELCOV(3,3),DE(3),GE(3,3)
DIMENSION RLX(3,3)
REAL*8 LAM
ESQ=1.0-(DATUM(2)/DATUM(1))**2
CP=DCOS(PHI)
SP=DSIN(PHI)
SL=DSIN(LAM)
CL=DCOS(LAM)
EW=DSQRT(1.0-ESQ*SP**2)
EN=DATUM(1)/EW
EM=EN*(1.0-ESQ)/EW**2
H1=EM+H
H2=(EN+H)*CP
GE(1,1)=-SP*CL/H1
GE(1,2)=-SP*SL/H1
GE(1,3)=CP/H1
GE(2,1)=-SL/H2
GE(2,2)=CL/H2
GE(2,3)=0.0
GE(3,1)=CP*CL
GE(3,2)=CP*SL
GE(3,3)=SP
C
CALL DGMPRD(GE,DX,DE,3,3,1)
DP=DE(1)
DL=DE(2)
DH=DE(3)
C
DO 14 I=1,3
DO 14 J=1,3
SUM=0.0
DO 12 K=1,3
DO 12 L=1,3
12 SUM=SUM+GE(I,K)*DXCOV(K,L)*GE(J,L)
14 DELCOV(I,J)=SUM
C
DO 15 J=1,3
RLX(1,J)=GE(1,J)*H1
RLX(2,J)=GE(2,J)*H2
15 RLX(3,J)=GE(3,J)
C
RETURN
END

```

```
SUBROUTINE DANG(ANGR,ISGN,IDEG,MIN,SEC)
IMPLICIT REAL*8(A-H,O,Z)
INTEGER BLANK/1H /,MINUS/1H-
ISGN=BLANK
IF(ANGR.LT.0.0) ISGN=MINUS
ANGD=57.295779513082D0*DABS(ANGR)
IDEG=IDINT(ANGD)
FMIN=ANGD-DFLOAT(IDEG)
FMIN=FMIN*60.0
MIN=IDINT(FMIN)
SEC=(FMIN-DFLOAT(MIN))*60.0
RETURN
END
```

```
SUBROUTINE UVWTG(UVW,DATUM,PHI,LAM,H)
C  CONVERT RECTANGULAR TO GEODETIC COORDINATES
  IMPLICIT REAL*8(A-Z)
  DIMENSION UVW(3),DATUM(2)
  LAM=DATAN2(UVW(2),UVW(1))
  IF(LAM.LT.0.0) LAM=LAM+6.28318530717958
  OME2=(DATUM(2)/DATUM(1))**2
  E2=1.0-OME2
  P=DSQRT(UVW(1)**2+UVW(2)**2)
  WP=UVW(3)/P
  TP1=WP/OME2
  PHI1=DATAN(TP1)
  5 TTP=TP1*TP1
  SECP=DSQRT(1.0+TTP)
  N=DATUM(1)*SECP/DSQRT(1.0+OME2*TTP)
  H=P*SECP-N
  TP2=WP/(1.0-E2*N/(N+H))
  PHI=DATAN(TP2)
  IF(DABS(PHI-PHI1).LT.1.0D-12) RETURN
  PHI1=PHI
  TP1=TP2
  GO TO 5
END
```

```
FUNCTION KSID2(IS)
  KSID2=KSTAID(IS)
  IF(KSID2.GT.0) RETURN
  WRITE(6,6000) IS
6000 FORMAT(//10X,'STATION NUMBER NOT FOUND IN INPUT LIST',15)
  STOP
  END
```

```

SUBROUTINE SATXYZ (XS,YS,ZS)
IMPLICIT REAL*8(A-H,O-Z)
INTEGER STANAM,IDS*2
COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
1STANAM(5,150),IDS(150)
COMMON/STAPLH/STAPLH(2,150)
COMMON/RANGED/RAN(50),VARRA(50),RMSMC,NSTE,KSTATE(50),ITEST,NITR
DIMENSION A( 3),AN(3,3),C(3),L(3),M(3),AX(3)
C   GET THE FIRST APPROXIMATION TO THE SATELLITE POSITION
PI=3.14159265358
TPI=2.*PI
PHI=0.0
ALAM=0.0
IS=NSTE
DO 20 I=1,IS
PHI=PHI+STAPLH(1,KSTATE(I))
STALNG=STAPLH(2,KSTATE(I))
IF(I.EQ.1) GO TO 19
IF(STALNG-STAPLH(2,KSTATE(1)).GT.PI)STALNG=STALNG-TPI
IF(STAPLH(2,KSTATE(1))-STALNG.GT.PI)STALNG=STALNG+TPI
19 ALAM=ALAM + STALNG
20 CONTINUE
PHI=PHI/IS
ALAM=ALAM/IS
H=1.6D06
IDTS=IDS(KSTATE(1))
CALL UVWD(DATPRM(1,1DTS),DATPRM(2,1DTS),PHI,ALAM,H,XS,YS,ZS)
C
NITR=0
25 CONTINUE
C   START ANOTHER ITERATION
NITR=NITR+1
WPW=0.0
DO 30 I=1,3
C(I)=0.0
DO 30 J=1,3
30 AN(I,J)=0.0
C
DO 50 IS=1,NSTE
DX=XS-STAUVW(1,KSTATE(IS))
DY=YS-STAUVW(2,KSTATE(IS))
DZ=ZS-STAUVW(3,KSTATE(IS))
R=DSQRT(DX*DX+DY*DY+DZ*DZ)
AL=RAN(IS)-R
WPW=AL*AL/VARRA(IS)**2+WPW
A(1)=DX/R
A(2)=DY/R
A(3)=DZ/R
DO 40 I=1,3
C(I)=C(I)+A(I)*AL/VARRA(IS)**2
DO 40 J=1,3
40 AN(I,J)=AN(I,J)+A(I)*A(J)/VARRA(IS)**2
50 CONTINUE
CALL DMINV(AN,3,DET,L,M)
CALL DGMPRD(AN,C,AX,3,3,1)
RMSMC=WPW/(NSTE-3)
C   TEST FOR CONVERGENCE
ICONVR=1

```

```
DO 55 I=1,3
  IF(DABS(AX(I)).GT.0.01) ICONVR=0
55 CONTINUE
C
C UPDATE
  XS=XS+AX(1)
  YS=YS+AX(2)
  ZS=ZS+AX(3)
  IF(ICONVR.EQ.1) RETURN
  IF(NITR.LT.20) GO TO 25
C  SET ITEST =2 INDICATED THAT CONVERGENCE WAS NOT OBTAINED IN 20 ITERATIONS
  ITEST=2
  RETURN
  END
```

```

SUBROUTINE RNG360
  IMPLICIT REAL*8 (A-H,O-Z)
C   S/360 VERSION OF SAR PROGRAM FOR SATELLITE DISTANCES
  INTEGER*2 PCODE(20)
  COMMON/PCODES/PCODE
  INTEGER*4 ENDSIG/1HE/,CONTIN,DELCOD(2)/1H ,1H*/,ECODE
  INTEGER*2 ID(50),IYR(50),IDAY(50),IHR(50),MIN(50)
  DIMENSION SEC(50),RAN(50),VARRA(50),MONTH(50),KSTATE(50)
  DIMENSION DN(3,3,150),BN(3,3,50),D0N(3,3),DK(3,150),DDK(3),A(3)
  COMMON/NSTA/NSTA
  DIMENSION NOBSTA(150),VPVSTA(150)
  COMMON/RANGED/RAN,VARRA,RMSMC,NSTE,KSTATE,ITEST,NITR
  DIMENSION POSSAT(3),DX(3)
  DIMENSION L1(3),L2(3)
  INTEGER STANAM,IDS*2
  COMMON/STALOC/STAUVW(3,150),DATPRM(2,15),DATNAM(4,15),
  1STANAM(5,150),IDS(150)
  COMMON/STAORD/KORDER(150)
  COMMON/WPW/WPW,XPU,1DEGF,NFSTA
  COMMON/STAPLH/STAPLH(2,150)
  DATA RPD/57.295779513/
C
  REWIND 2
  REWIND 3
  READ (3) TD,STAPLH
  WRITE(6,6004) TD
  6004 FORMAT(//20X,'TEST VARIANCE =',F20.2)

  WRITE(6,6001)
  6001 FORMAT(///'0STATION',T12,'DATE',T24,'TIME',T43,'RANGE',T60,
  1  'UNCERTAINTY',T76,'MISCLUSION')
C
C
  DO 70 KSTA=1,NSTA
  NOBSTA(KSTA)=0
  VPVSTA(KSTA)=0.0
  DO 70 I=1,3
  DK(I,KSTA)=0.0
  DO 70 J=1,3
  DN(I,J,KSTA)=0.0
  70 CONTINUE
C
  KEVENT=0
  EPR=0.0
  210 CONTINUE
  READ (3) IEVENT,NSTE,EPR, (ID(IS),IYR(IS),MONTH(IS),IDAY(IS),
  X IHR(IS),
  1MIN(IS),SEC(IS),RAN(IS),VARRA(IS),KSTATE(IS),IS=1,NSTE),CONTIN
  WRITE(6,6008) IEVENT
  6008 FORMAT(/ 1X,'EVENT',16)
  ITEST=0
  IF(NSTE.GT.3) GO TO 220
C   SET ITEST =1 TO INDICATE THAT LESS THAN FOUR STATIONS WERE OBSERVING
  ITEST=1
  DO 280 IS=1,NSTE
  280 WRITE(6,6009) ID(IS),IYR(IS),MONTH(IS),IDAY(IS),IHR(IS),
  1MIN(IS),SEC(IS),RAN(IS),VARRA(IS)

```

```

6009 FORMAT(17,2X,12,1X,A3,1X,I2,2X,2I3,F8.4,F18.3,F15.2,F15.2)
GO TO 630
220 CONTINUE
CALL SATXYZ(XS,YS,ZS)
IF(RMSMC.GT.TD) ITEST=3
IF(ITEST) 311,300,311
C
C      SET UP OBSERVATION EQUATIONS FOR THIS EVENT AND COMPUTE
C      CONTRIBUTIONS TO THE NORMAL EQUATIONS
300 CONTINUE
KEVENT=KEVENT+1
DO 310 I=1,3
DDK(I)=0.0
DO 310 J=1,3
DDN(I,J)=0.0
310 CONTINUE
C
311 CONTINUE
POSSAT(1)=XS
POSSAT(2)=YS
POSSAT(3)=ZS
DO 390 IS=1,NSTE
KSTA=KSTATE(IS)
DO 305 I=1,3
DX(I)=POSSAT(I)-STAUVW(I,KSTA)
RC=DSQRT(DPDOT(DX,DX,3))
AL=RAN(IS)-RC
WRITE(6,6009) ID(IS),IYR(IS),MONTH(IS),IDAY(IS),IHR(IS),
1MIN(IS),SEC(IS),RAN(IS),VARRA(IS),AL
IF(ITEST) 390,307,390
307 CONTINUE
COMPUTE WEIGHT
WT=1./VARRA(IS)**2
DO 306 I=1,3
306 A(I)=DX(I)/RC
C
COMPUTE VPK OF MISCLOSURES
VPVTO=AL*WT*AL
VPVSTA(KSTA)=VPVSTA(KSTA)+VPVTO
NOBSTA(KSTA)=NOBSTA(KSTA)+1
COMPUTE CONTRIBUTIONS TO NORMAL EQUATIONS
DO 330 I=1,3
TERM=A(I)*WT*AL
DK(I,KSTA)=DK(I,KSTA)-TERM
DDK(I)=DDK(I)+TERM
DO 330 J=1,3
TERM=A(I)*WT*A(J)
BN(I,J,IS)=-TERM
DN(I,J,KSTA)=DN(I,J,KSTA)+TERM
DDN(I,J)=DDN(I,J)+TERM
330 CONTINUE
390 CONTINUE
C
IF(ITEST) 600,391,600
391 CONTINUE
CALL DMINV(DDN,3,DET,L1,L2)
WRITE(2) NSTE, DDN,DDK,((BN(I,J,IS),I=1,3),J=1,3),KSTATE(IS),

```

```

1IS=1,NSTE),CONTIN
600 CONTINUE
  WRITE(6,6011)RMSMC,NITR
6011 FORMAT(' RMSMC = ',F10.2,
           ' NITR = ',I3)
           ' VARIANCE OF EVENT ADJUSTMENT = ',F10.2,
           ' AFTER',I3,' ITERATIONS')
           IF(PCODE(11)) 610,630,610
610 IF(PCODE(11)=3) 611,612,611
611 WRITE(6,6022) POSSAT
6022 FORMAT(' SATELLITE POSITION',3F15.3)
           IF(PCODE(11)=2) 612,630,612
612 IDTS=IDS(KSTATE(1))
           CALL UVWTG2(POSSAT,DATPRM(1,1),PHI,FLAM,H)
           PHI=PHI*RPD
           FLAM=FLAM*RPD
           WRITE(6,6023) PHI,FLAM,H
6023 FORMAT(' GEOD. COORD. OF SATELLITE',2F14.6,F14.1)
630 CONTINUE
           IF(ITEST) 290,640,290
290 WRITE(6,6015) ITEST
6015 FORMAT(1H ,27X,'ENTIRE EVENT DELETED. DELETION CODE = ',I3)

```

```

C   ITEST=0 MEANS A GOOD EVENT
C   ITEST=1 MEANS NOT ENOUGH OBSERVATIONS
C   ITEST=2 MEANS MORE THAN 20 ITERATIONS WERE REQUIRED TO GET APPROXIMATE
C   SATELLITE POSITION
C   ITEST=3 MEANS THE EVENT IS REJECTED BECAUSE THE EVENT VARIANCE IS GREATER
C       THAN THE TEST VARIANCE
640 CONTINUE
C
C   IF(CONTIN.EQ.ENDSIG) GO TO 700
C   GO TO 210
C
C   700 CONTINUE
C
CHECK TO SEE IF END SIGNAL HAS BEEN WRITTEN ON DATA SET FT02
  IF(ITEST.EQ.0) GO TO 710
  BACKSPACE 2
C   READ AND WRITE LAST RECORD FROM LSST GOOD EVENT
  READ (2) NSTE,  DDN,DDK,(((BN(I,J,IS),I=1,3),J=1,3),KSTATE(IS),
  1IS=1,NSTE)
  WRITE(2) NSTE,  DDN,DDK,(((BN(I,J,IS),I=1,3),J=1,3),KSTATE(IS),
  1IS=1,NSTE),CONTIN
710 CONTINUE
  WRITE(2) (((DN(I,J,KSTA),I=1,3),DK(J,KSTA),J=1,3),
  XKSTA=1,NSTA)
C   WRITE(6,6018)(KORDER(KSTA),((DN(I,J,KSTA),J=1,3),I=1,3),
C   1KSTA=1,NSTA)
  6018 FORMAT(1I5/3(3D18.7/))
  WPW=0.0
  NOBS=0
  WRITE(6,6019)
6019 FORMAT(1H1,8(/),10X,'ANALYSIS OF MISCLOSURES BY STATION'//,
  1T10,'STATION',T20,'NUMBER OF OBSERVATIONS',T50,'RMS MISCLOSURE')
  DC 750 KSTA=1,NSTA

```

```
NOBS=NOBS+NOBSTA(KSTA)
WPW=WPW+VPVSTA(KSTA)
RMSMC=0.0
IF(NOBSTA(KSTA).GT.0) RMSMC=DSQRT(VPVSTA(KSTA)/DFLOAT(NOBSTA(KSTA))
1)
WRITE(6,6020) KORDER(KSTA),NOBSTA(KSTA),RMSMC
6020 FORMAT(T10,I7,T35,I7,T50,F14.2)
750 CONTINUE
IDEGF=NOBS-3*KEVENT
RMSMC=DSQRT(WPW/DFLOAT(IDEGF))
WRITE(6,6021) NOBS,KEVENT,IDEGF,WPW,RMSMC
6021 FORMAT(///10X,'TOTAL NUMBER OF GOOD OBSERVATIONS',T60,I8//
110X,'TOTAL NUMBER OF GOOD EVENTS',T60,I8,/)
210X,'CORRESPONDING DEGREES OF FREEDOM',T60,I8//
310X,'TOTAL SUM OF SQUARES OF MISCLOSURES',T60,F11.2//
410X,'CORRESPONDING STANDARD DEVIATION OF UNIT WEIGHT',T60,F11.2)
RETURN
END
```

```

SUBROUTINE RRDATA
IMPLICIT REAL*8(A-H,O-Z)
INTEGER*4 ENDSIG/1HE/,CONTIN
INTEGER*2 ID(50),IYR(50),IDAY(50),IHR(50),MIN(50)
DIMENSION SEC(50),RAN(50),VARRA(50),MONTH(50),KSTATE(50)
COMMON/STAPLH/STAPLH(2,150)
COMMON/OBSD/OBSD(150),OVOBSD
INTEGER*2 PCODE(20)
COMMON/PCODES/PCODE
MAXSTE=50
SPR=206264.80625
PI=3.14159265358
PI2=2.0*PI
WPWSP=0.0
READ(5,5004) TD,OVOBSD
WRITE(6,6004) TD
5004 FORMAT(F20.2,F10.2)
6004 FORMAT(//20X,TEST VARIANCE =,F20.2)

```

```

C
C   START DATA INPUT
    REWIND 3
    WRITE(3) TD,STAPLH
    IS=0
    IEVENT=0
    EPR=0.0
C   ENTER HERE FOR A NEW OBSERVATION
C
200 IS=IS+1
205 CONTINUE
    READ(5,1022,END=901) ID(IS),IYR(IS),MONTH(IS),IDAY(IS),IHR(IS),
    1MIN(IS),SEC(IS),SEC1,RAN(IS),RA1,VARRA(IS),CONTIN
    RAN(IS)=RAN(IS)+RA1/1000.
1022 FORMAT(14X,14,5I2,F2.0,F4.0,F16.0,F3.0,11X,F6.3,9X,A1)
    IF(SEC1.LT.1.) GO TO 201
    SEC(IS)=SEC(IS)+SEC1/10000.
    GO TO 202
201 SEC(IS)=SEC(IS)+SEC1
202 CONTINUE
    IF(CONTIN.EQ.ENDSIG) GO TO 250
    DDT=DFLOAT(MJD(IDAY(IS),MONTH(IS),IYR(IS)))
    DDT=DDT+(DFLOAT((IHR(IS)*60+MIN(IS))*60)+SEC(IS))/864.0D2
    IF(IS.LE.1) GO TO 210
C   THIS TEST SHOULD BE TRUE ONLY FOR THE FIRST CARD OF THE FIRST EVENT.
C
CHECK FOR END OF EVENT, ALLOWING 0.5 MS DISCREPANCY
    IF(DABS(DDT-EPR).GT.0.58D-8) GO TO 250
C
C   ENTER HERE TO BEGIN A NEW EVENT
C   THE FIRST ENTRY OF THE EVENT SHOULD ALWAYS BE MADE WITH IS=1
210 CONTINUE
    IDD=ID(IS)
    KSTA=KSTAID(IDD)
    IF(KSTA.GT.0) GO TO 220
    WRITE(6,6042) ID(IS),IHR(IS),MIN(IS),SEC(IS),IDAY(IS),MONTH(IS),
    1IYR(IS)
6042 FORMAT(5X, 'STATION NUMBER NOT FOUND IN INPUT LIST',15,3X,2I3,

```

```
1F8.4,3X,I3,A3,I2,'OBSERVATION IGNORED')
GO TO 205
220 CONTINUE
IF(PCODE(12).EQ.1) VARRA(IS)=OBSD(KSTA)
IF(PCODE(12).EQ.2) VARRA(IS)=OVOBSD
KSTATE(IS)=KSTA
EPR=DDT
GO TO 200
C
C   END OF INPUT FOR THIS EVENT. BEGIN PROCESSING
250 CONTINUE
NSTE=IS-1
IEVENT=IEVENT+1
IF(IHR(NSTE+1).EQ.99) CONTIN=ENDSIG
WRITE(3) IEVENT,NSTE,EPR, (ID(IS),IYR(IS),MONTH(IS),IDAY(IS),
X IHR(IS),
1MIN(IS),SEC(IS),RAN(IS),VARRA(IS),KSTATE(IS),IS=1,NSTE),CONTIN
C   TEST FOR END OF INPUT
IF(CONTIN.EQ.ENDSIG) GO TO 700
C   PREPARE FOR NEXT EVENT
ID(1)=ID(NSTE+1)
IYR(1)=IYR(NSTE+1)
MONTH(1)=MONTH(NSTE+1)
IDAY(1)=IDAY(NSTE+1)
IHR(1)=IHR(NSTE+1)
MIN(1)=MIN(NSTE+1)
SEC(1)=SEC(NSTE+1)
RAN(1)=RAN(NSTE+1)
VARRA(1)=VARRA(NSTE+1)
C   RETURN TO START A NEW EVENT
IS=1
GO TO 210
C
700 RETURN
C
C   ERROR EXITS
901 CONTINUE
C   ENTER HERE IF END SIGNAL CARD IS MISSING FROM INPUT DATA DECK
CONTIN=ENDSIG
GO TO 250
END
```